

SiPM: almost an Ideal Low Light Level Sensor for CTA SST

Razmik Mirzoyan

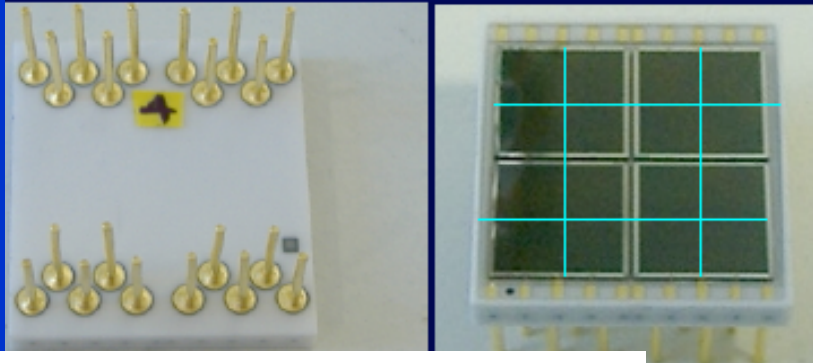
Max-Planck-Institute for Physics
(Werner-Heisenberg-Institute)
Munich, Germany

What LLL sensor can we dream about ?

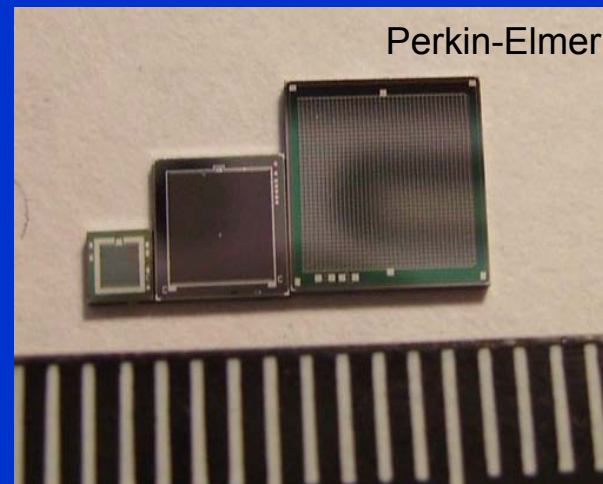
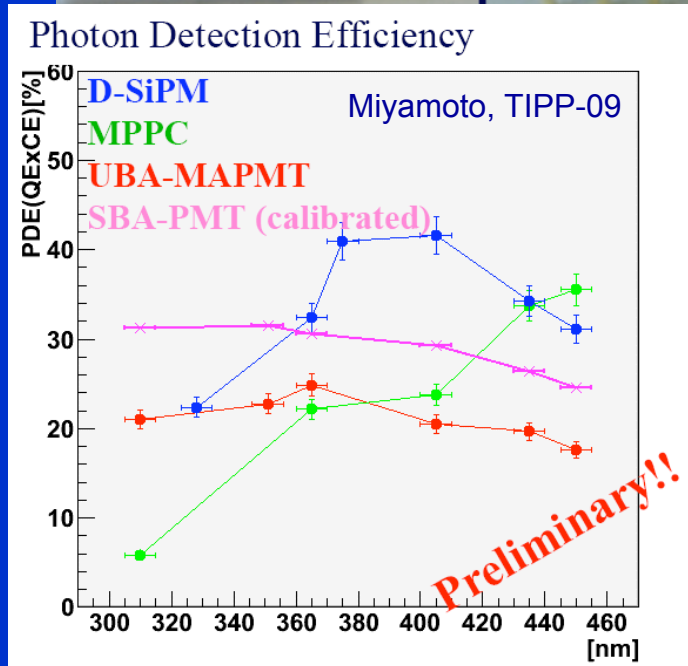
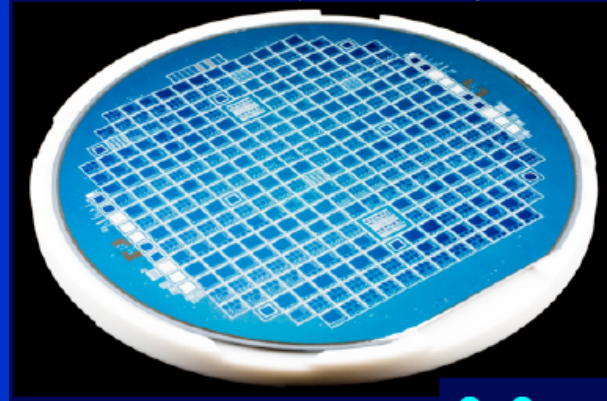
- Nearly 100 % QE and photon detection efficiency (PDE)
- Could be made in very large and in very small sizes
- Few ps fast (in air and in many materials the light speed is usually 20-30 cm/ns; in 5 ps it will make 1-1.5 mm)
- Signal amplification $\times 10^6$
- Noiseless amplification: F-factor - 1.001
- Few % amplitude resolution
- No fatigue, no degradation in lifetime
- Low power consumption
- Operation at ambient temperatures
- No danger to expose to light
- Insensitive to magnetic fields
- No vacuum, no HV, lightweight,...

Few examples of SiPMs, still under development

Hamamatsu (MPPC)



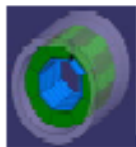
MPI-HLL (SiMPL)



SiPMs with ~ 50-60% PDE and low cross-talk (< 1%) could be anticipated within 1-2 years

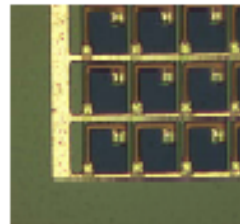
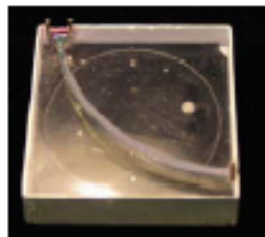
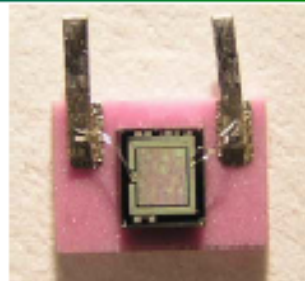
ILC: Potential Consumer of $(5-200) \times 10^6$ SiPMs

- Scintillation Calorimetry- for instance a SciTile Imagine Hadron Calorimeter for ILC (CALICE Collaboration), sci tile size: a few cm
- Typical threshold is $\sim 5-7$ phe

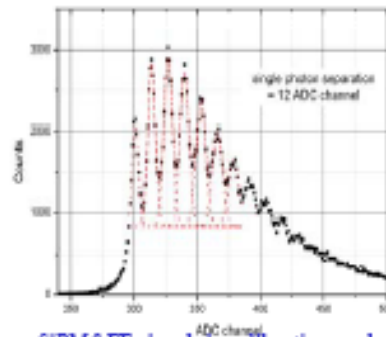


SiPM tile fibre system

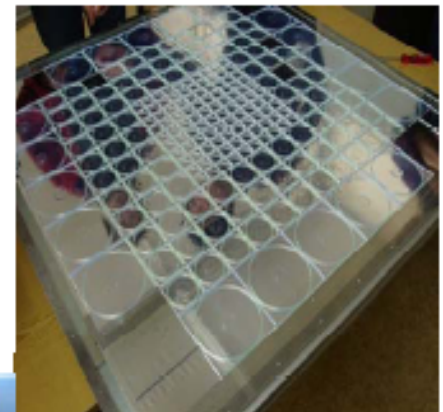
- SiPM developed by MEPHI/PUSAR
 - Gain $\sim 10^6$, bias ~ 50 V, size 1 mm^2 , 1156 pixels
 - Eff (green) $\sim 15\%$, quenching R $\sim 1 - 10 \text{ M}\Omega$
- SiPM tile fibre system integration: ITEP
 - $3 \times 3 \times 0.5 \text{ cm}^3$ tiles from UNIPLAST, Russia
 - WLS fibre Kuraray Y11(300) 1mm
 - Matted edges, 2% light xtalk per edge
 - Faces covered with EM mirror foil



A big 8000 channel HCAL prototype with tail catcher is constructed by CALICE (DESY, ITEP, LAL, MEPHI, NIU, Prague, UK) for analogue and semidigital modes

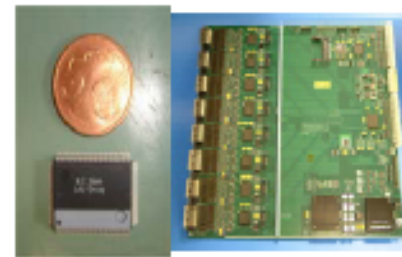


SiPM&FE signals in calibration mode



One plane with SiPMs and WLS fibers installed into 3×3 , 6×6 and $12 \times 12 \text{ cm}^2$ 0.5 cm thick tiles

CERN test beam, 2006



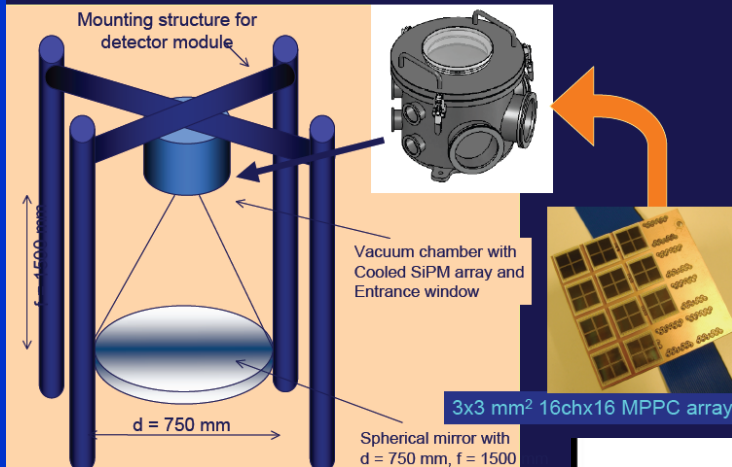
LAL 18 ch. SiPM FE chip

B. Dolgoshein, SiPM review

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Outlook

- Telescope with MPPC array camera



4-SiPMs of $5 \times 5 \text{ mm}^2$, includes cooling, signal shaping



A 22mmx22mm SiPM based pixel for a telescope

The same as on the left but 4-times larger



PROCEEDINGS OF THE 31st ICRC, LODZ 2009

SiPM development and application for astroparticle physics experiments

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Abstract. A Silicon Photomultiplier (SiPM, G-APD) is a novel solid state photodetector which has an outstanding photon counting ability. The device has excellent features such as high quantum efficiency, good charge resolution, fast response ($< 100 \text{ ps}$), very compact size, high gain (up to $2 \cdot 3 \cdot 10^6$), very low power consumption with low bias voltages (30-70V), immunity to the magnetic field. In the last few years, UV sensitive SiPMs with a p-on-n structure have been developed by a few companies such as Hamamatsu, Photonique, Zecotek, Photonics Inc., and institutes such as the MPiHL (Max-Planck-Institute for Physics - Max-Planck-Institute Semiconductor Laboratory) as well as the MPI-MEPH (Max-Planck-Institute for Physics - Moscow Engineering Physics Institute) for astroparticle physics applications. Here the current status of the SiPM development in MPI and HLL, MPI and MEPH, and the study of the application to imaging atmospheric Cherenkov telescopes (IACTs) MAGIC/MAGIC-II [1] and CTA [2], and a fluorescence telescope in the space JEM-EUSO [3] will be reported.

Keywords: Imaging Cherenkov, Imaging fluorescence, SiPM

I. INTRODUCTION

The high PDE of these devices will allow us to lower the threshold energy of gamma ray detection down to 10 - 20 GeV in case of MAGIC telescopes, and ensure the detection efficiency of UHECRs above $(2-3) \cdot 10^{19} \text{ eV}$

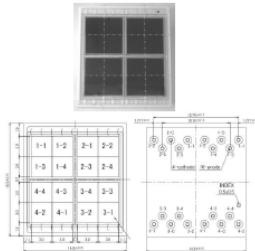


Fig. 1. Top: Left/Center: Blue print of 16ch (4x4) of $3 \times 3 \text{ mm}^2$ MPPC array device (front/back). Bottom: Photo of 16 ch MPPC array device.

SOME EXAMPLES OF SHOWERS RECORDED BY MAGIC AND THE G-APD PIXEL

EVALUATION OF IMPROVEMENT

NOTE: G-APD SIGNAL MUST BE CORRECTED FOR OPTICAL CROSS-TALK

Shower Signals: G-APD vs PMT

FINDINGS

- The tests IN 2000-2008 have confirmed that Cherenkov light from air showers can be detected
- P-on-n type G-APDs are available now with high sensitivity in the "blue", matched to Cherenkov spectrum (but UV sensitivity can still be raised, by design or use of WLS)
- Tests confirmed 2x gain compared to flat window, standard bakelite PMTs (about a factor 1.6 improvement compared to advanced heterostructural joints with diffuse lacquer coating and special light collection as in the MAGIC camera (the 50chx16 MPPC))
- No cooling necessary: intrinsic noise ~ night sky illumination rate
- Clip cable or diT. Amplifier allows to shorten pulse width
- The currently available densely packed arrangement of 16×16 of $3 \times 3 \text{ mm}$ each is already suitable for grids of a high resolution imaging

PDE collection efficiency

CE G-APD	0.65
CE at flat PMT	0.9
CE Shg PMT	0.9
CE ET RBCS PMT	0.95
CE MAGIC	0.95
CE Shg PMT	0.95
Mesh dynode	0.65

NOTE THE MAIN PROBLEM: G-APDs CAN HAVE A HIGH QE OVER WIDE SPECTRAL RANGE BUT THE CURRENTLY TOO HIGH GAIN OF LARGE CELL TYPE DEVICES PREVENTS THE OPERATION AT HIGH OVERVOLTAGE

- PDE IS WELL BELOW QE BECAUSE HIGH GAIN CAUSES HIGH OPTICAL CROSS-TALK (\approx PHOTONS/HP)
- MUST OPERATE G-APDS WITH LOW OVERVOLTAGE
- THE KEY REQUIREMENT: LOWER THE GAIN PER CELL OF $(100 \times 100) \mu\text{m}$ TO OPERATE AT $\approx 4 \text{ V}$

CONCLUSIONS:

Air shower measurements by using a 256 ch. matrix of $3 \times 3 \text{ mm}^2$ Hamamatsu MPPCs in the focus of a 60cm F/2.5 mirror from the roof of the MPI building

FACT = First G-APD Cherenkov Telescope for TeV Gamma Astronomy

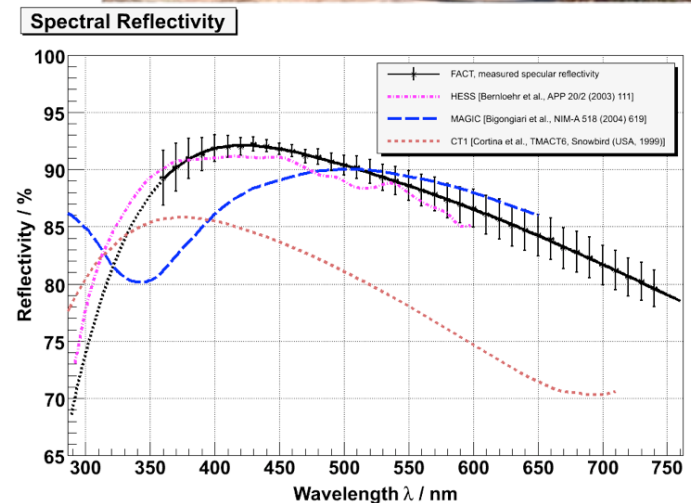
TU Dortmund, EPF Lausanne, U Würzburg, ETH Zürich

■ Goal:

- Crab-observations in the coming period
- Test the technology (**CTA**, **AGIS**, **MAGIC**, ...)
- Will be operated for the DWARF-projects physics program !

■ Ready availability

- HEGRA-3 Telescope-Mount (La Palma):
- 8.5 m² mirror of good reflectivity
- Microcontroller-based Drive system
- Experience with 36 Pixel-Test camera M0

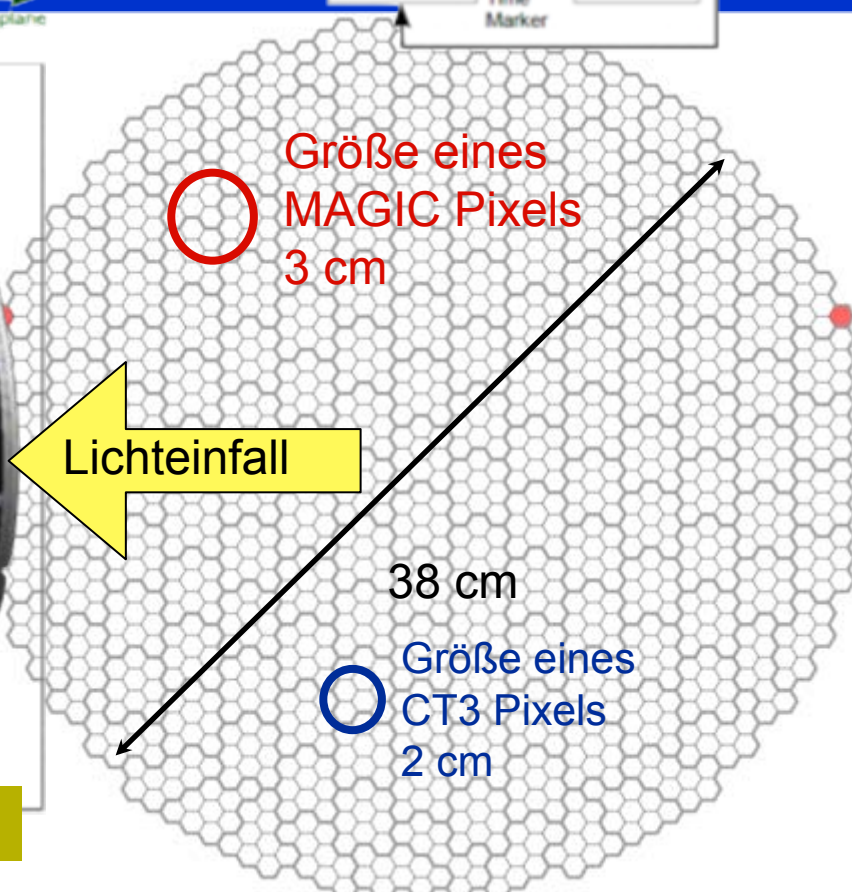
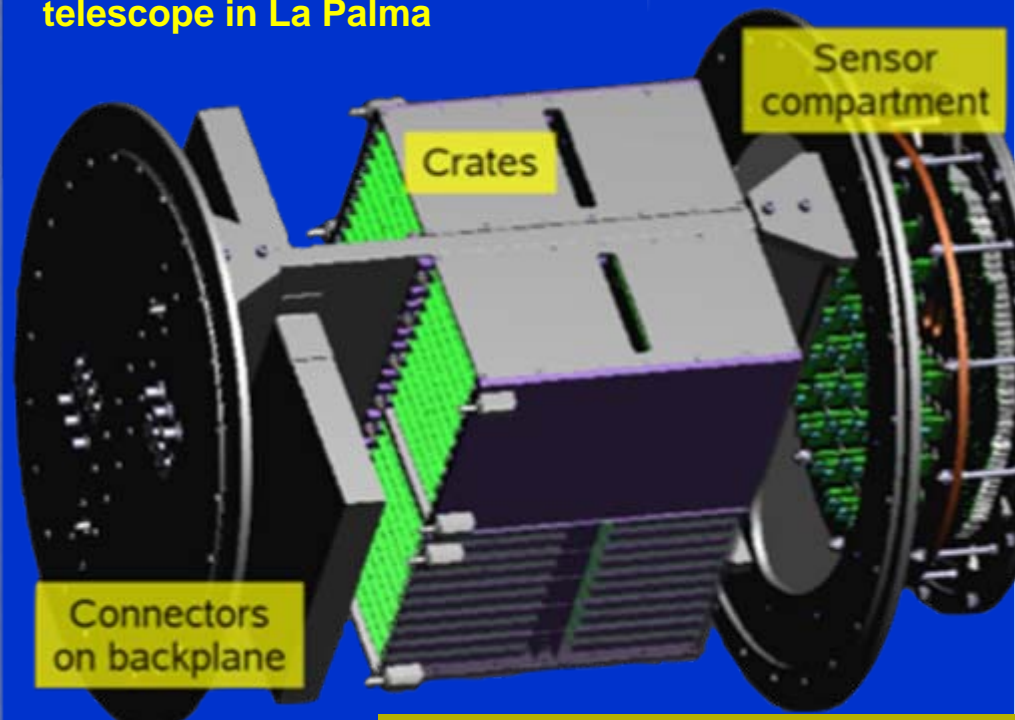
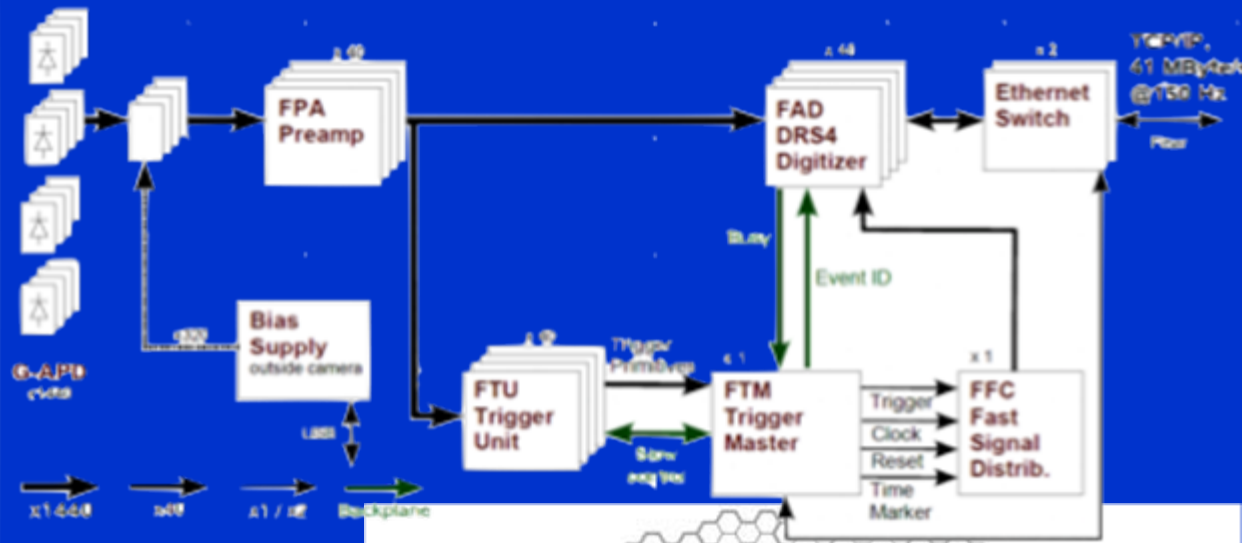


2010: 1440-Pixel G-APD-Kamera

DAQ in Kamera integriert

Produced in Zurich

In fall 2010 will be installed in the Focus of the 4.2m HEGRA 3 telescope in La Palma



Detail structure of the camera

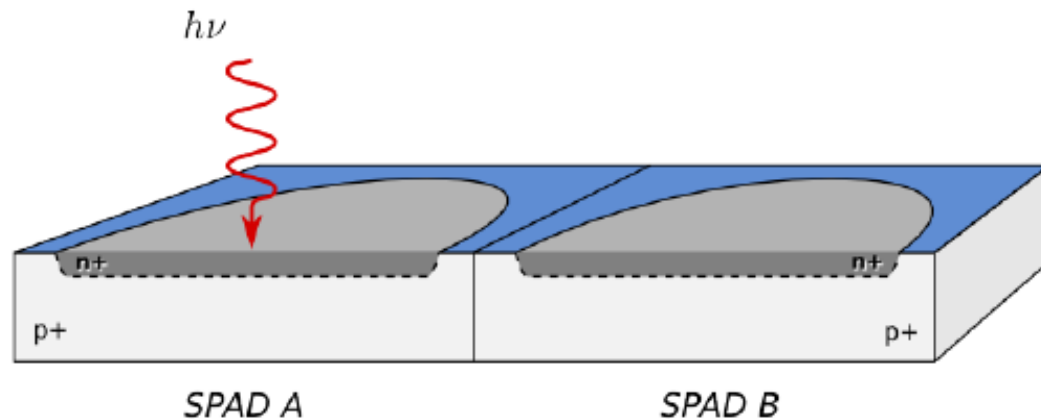
7-8 September 2010

SST meeting, LI

Why the light emission from Si avalanches is so important

- First observation of the light emission from reversed-biased Si p-n junction in 1955 (Newman)
- Revived interest about the effect in recent years because of:
- Cross-talk in SiPMs (GAPD, MPPC, micro-channel APD,...) spoils the amplitude resolution
- The light emission is proportional to the number of e⁻ in the avalanche. This puts a limit to the maximum gain under which one can operate the SiPMs
- If no measures are taken against the cross-talk, then the F-factor is worse than in classical PMTs
- As a consequence one encounters major problems in self-trigger schemes when measuring very low light level signals

Cross-Talk

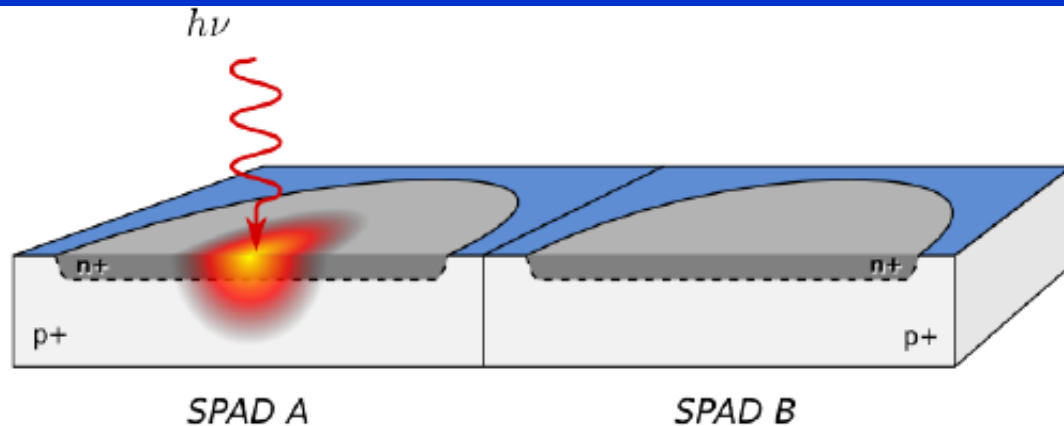


Ingargiola NDIP-08

When an avalanche is triggered in one SPAD we have:

- Secondary photons **emission** due to the avalanche current
- Photons **propagation** throughout the chip
- Secondary photon **detection** by a nearby detector

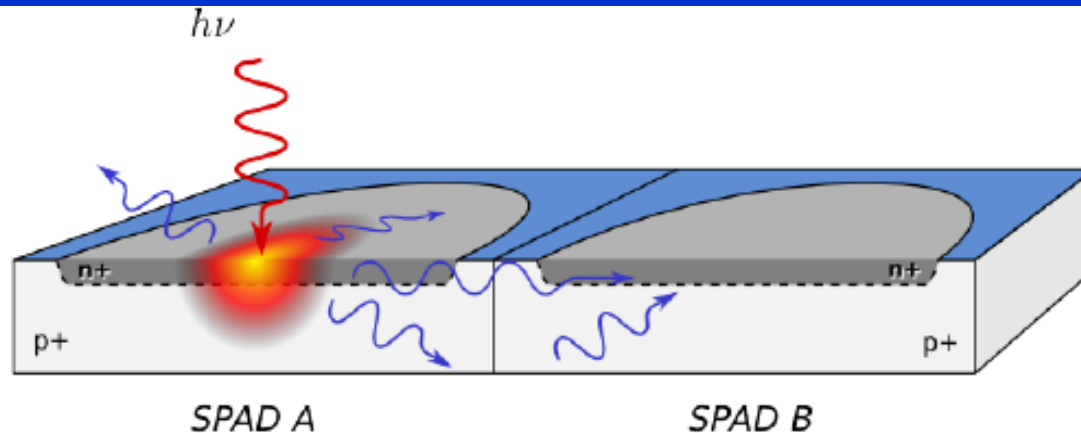
Cross-Talk



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Cross-Talk

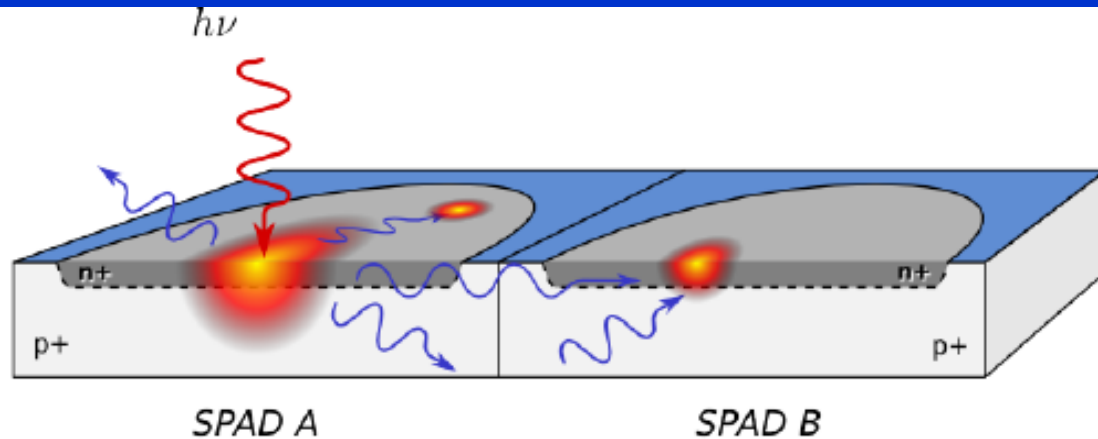


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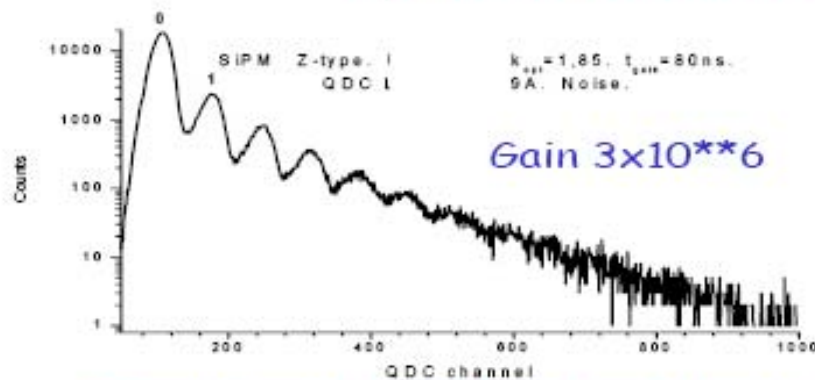
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Current status of SiPM and the prospects

- Currently there is a lot of enthusiasm about the new devices but the deep understanding is not simple, it comes only slowly
- One of the main problems of SiPMs is the low PDE, that is not easy to measure. It shall be disentangled from the cross-talk and afterpulsing.
- The afterpulsing in PMTs is a $\sim 1\%$ effect on single ph.e. level, while for example, for currently existing MPPC's from Hamamatsu it is a 20-30 % effect. This shall strongly manifest itself in self-trigger mode
- Usually the real value of PDE is much lower than the claimed (advertised) one. The reason is the low applied overvoltage.
- For $\sim 100\%$ Geiger efficiency and a high PDE one needs to apply an overvoltage that is 15-20 % higher than the breakdown one. The commercially available devices cannot do this yet (because of their design they do not quench above an applied overvoltage of 2.5 %).
- Already during this year some type of SiPMs with good UV response and a low cross-talk level could become available.
- Hamamatsu, Philips, Perkin-Elmer and some other companies are working on it.

→ Long tail in SiPM pulse height distribution vs threshold

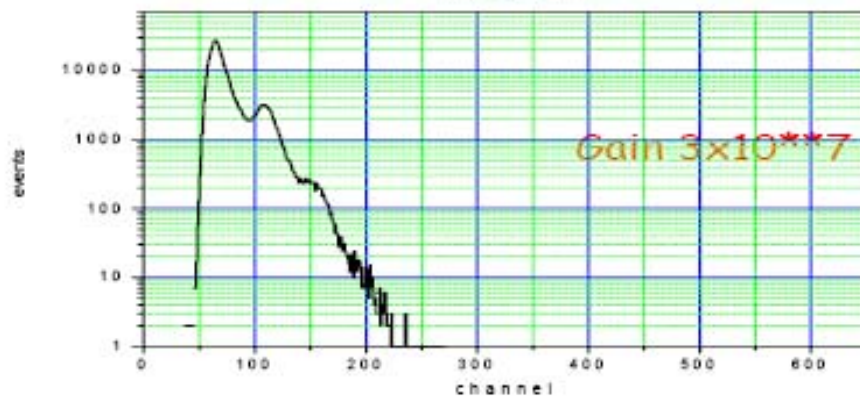
Optical crosstalk, SiPM 1x1 mm², dark noise



Crosstalk → non-Poissonian distribution:

pixel fired/phe=1.7

ENF=1.6



Crosstalk suppression by special SiPM topology:

Poisson distribution:

pixel fired/phe= ~1

ENF= ~1

Optical Crosstalk OC

P. Buzhan, B. Dolgoshein, et al., 2009

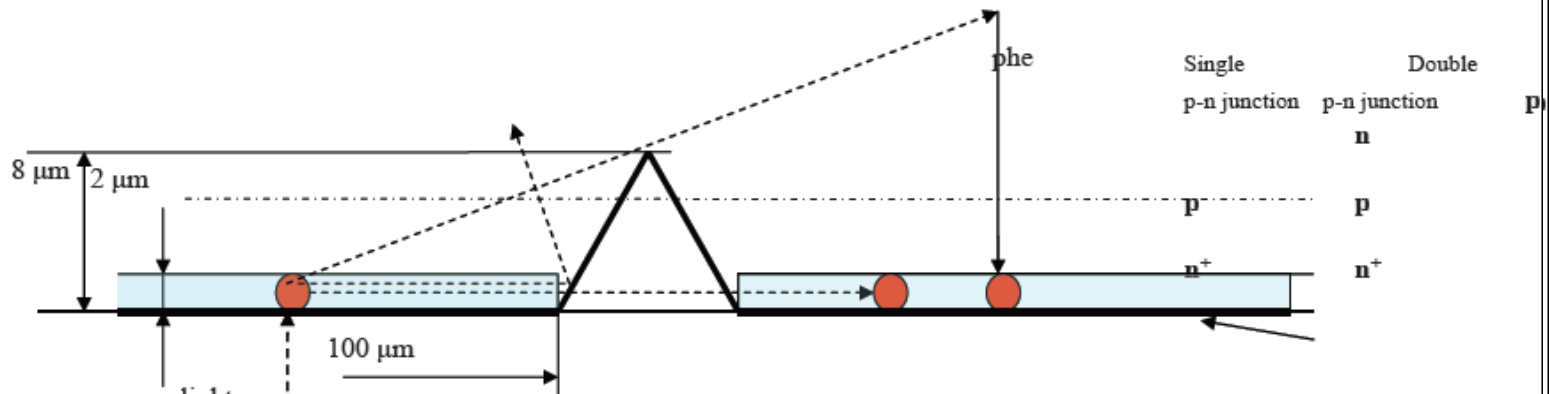
OC has two components

FIRST: phe's are induced in high electric field depletion region of neighbouring pixels

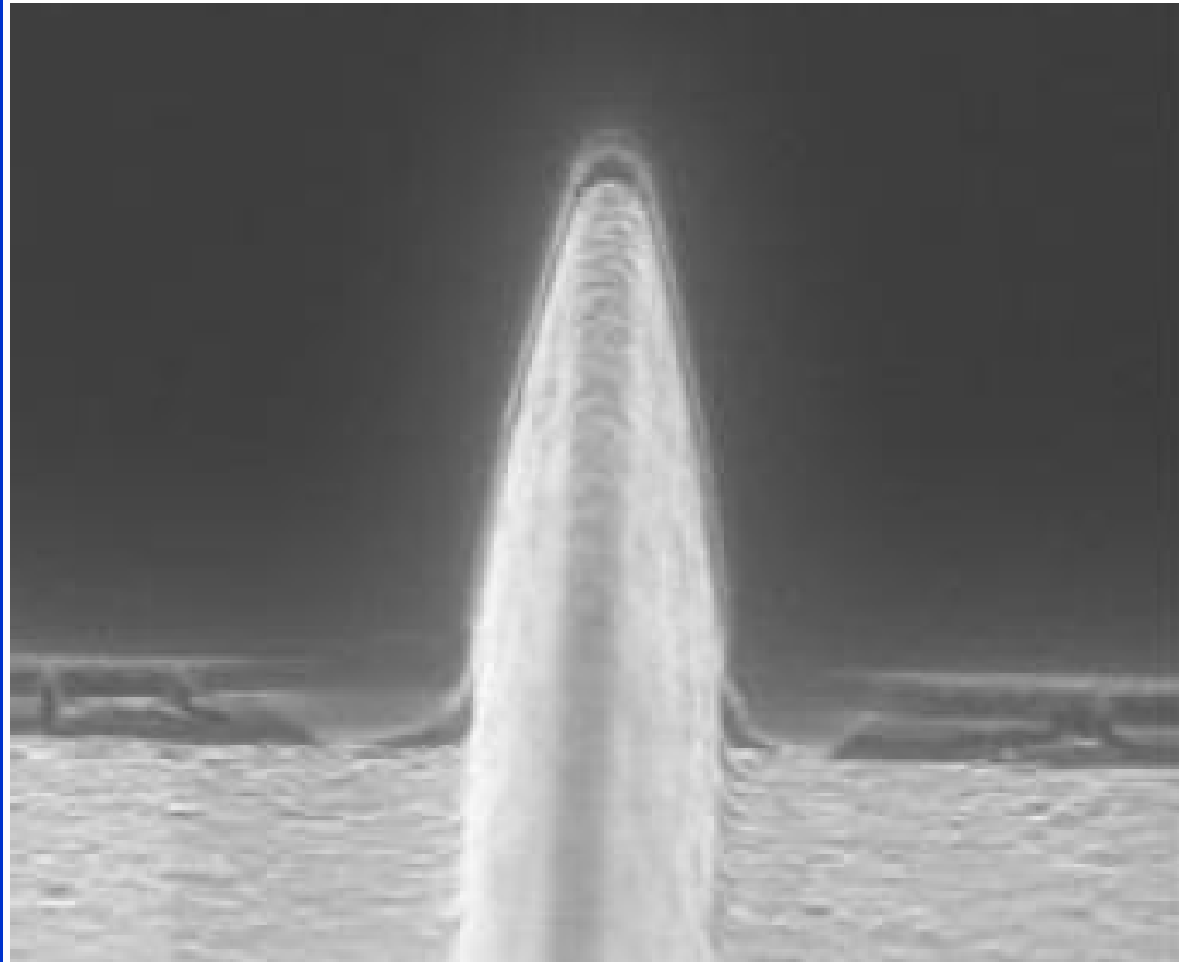
→ this mechanism is very fast: $\sim 1\text{ns}$ (prompt OC)

SECOND: The same in undepleted region and then the diffusion (or drift) to high electric field Geiger region of neighbouring pixels

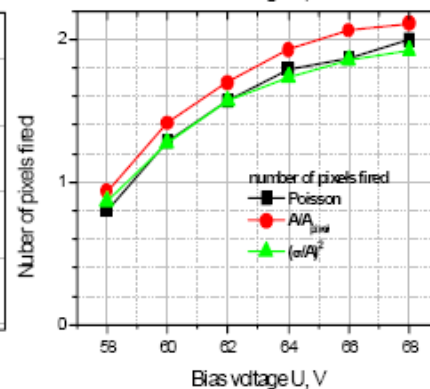
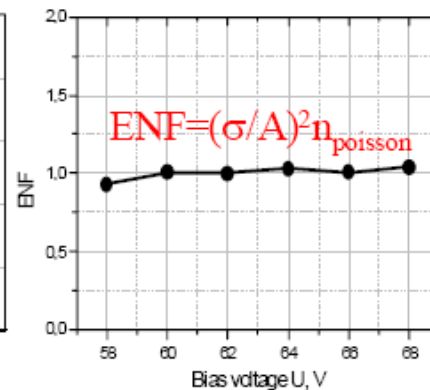
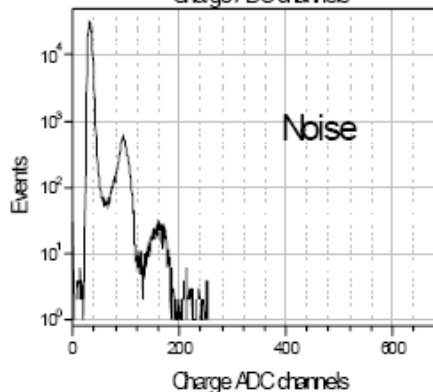
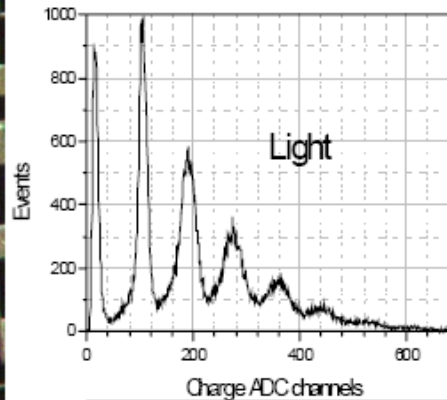
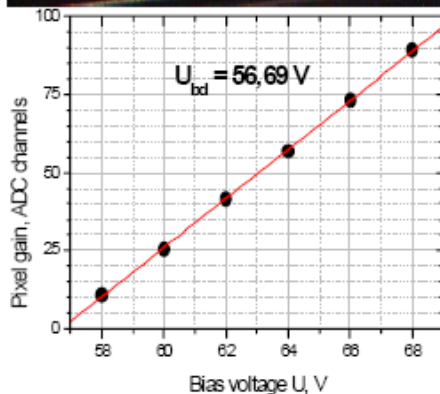
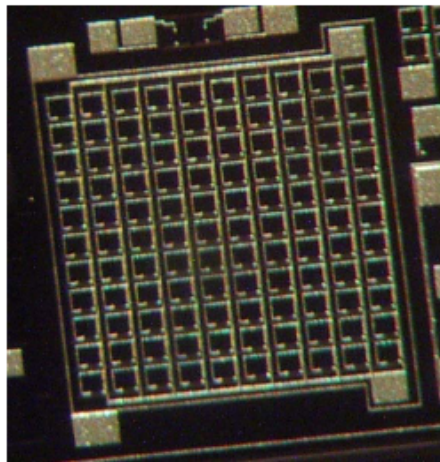
→ this process is delayed: later than 1ns



A filled in trench

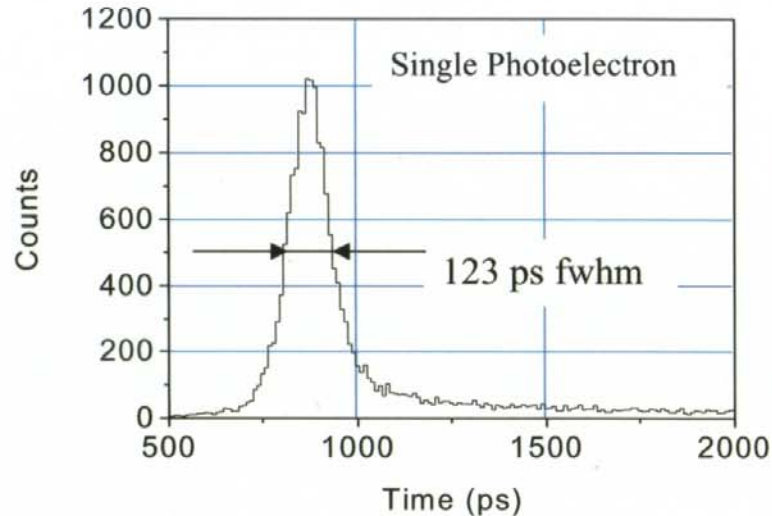


First step: SiPM 1.4x1.4 mm² with OC suppression topology

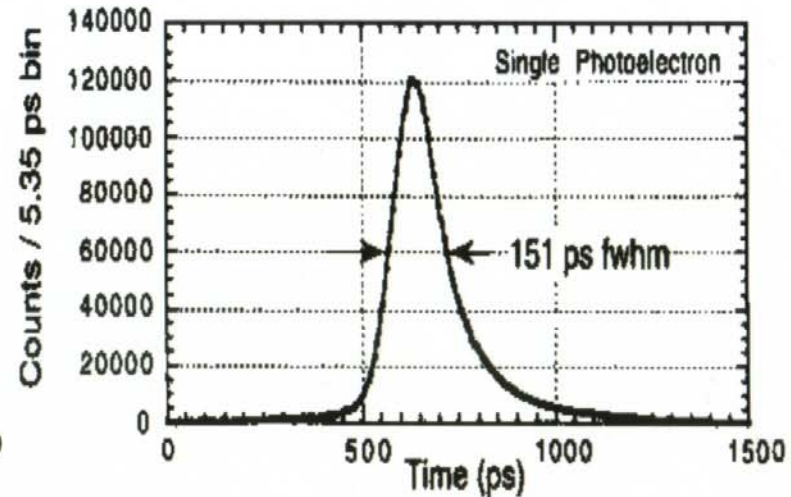


Timing by SiPM: possible application for Cherenkov Imaging Counters

SiPM



PMT R-5320



SiPM:

- position sensitive ($\sim 1 \text{ mm}^2$)
- a single photon detection capability with background hits density : $2 \cdot 10^{-3} \text{ 1/ns} \cdot \text{mm}^2$ (room temperature)
 $3 \cdot 10^{-4} \text{ 1/ns} \cdot \text{mm}^2$ (-50°C)

- insensitive to magnetic field
- good time resolution ($\sim 50 \text{ ns rms}$)

FWHM: Laser (40 ps) + electronics (60 ps) => SiPM (100 ps)

Second step: 5x5mm² SiPM with OC and AP suppression

SiPM parameters:

→ size	5x5mm ²
→ double junction structure with optical barriers 6mkm	
→ number of pixels	1600
→ pixel size	100mkm
→ gain	2×10^7
→ geometrical eff.(filling factor)	64%
→ pixel capacitance	~1pF
→ output SiPM capacitance	~160pF
→ antireflection entrance window	
→ single pixel recovery time	~ .5mks

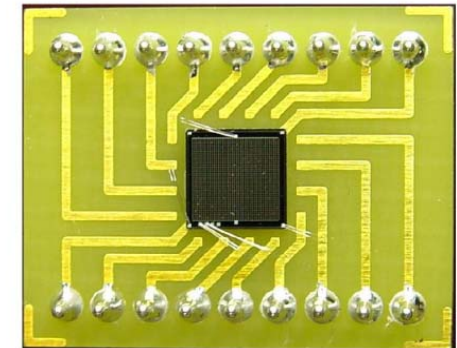
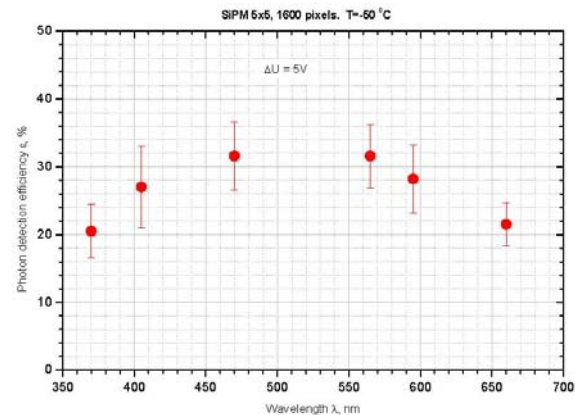


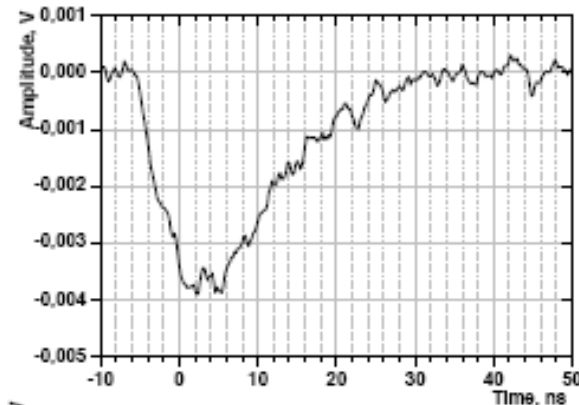
Figure 3: $25(5 \times 5)\text{mm}^2$ SiPM. It consists of the array of 1600(40×40) micropixels with $100 \times 100\mu\text{m}^2$ size.

Timing by 5x5mm² SiPM: signal shape

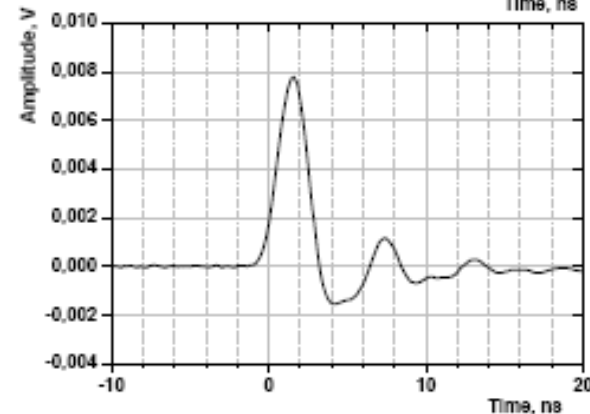
→ Because high SiPM output capacitance (~160pF)

a special FE electronics has been developed:

low input impedance (a few Ohm)
current amplifier+shaper

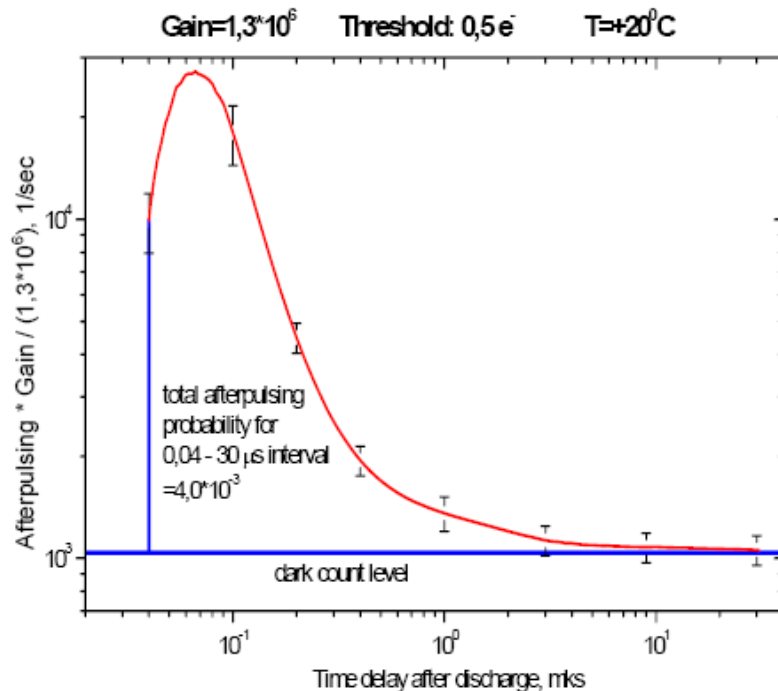


50 Ohm
FWHM
15ns



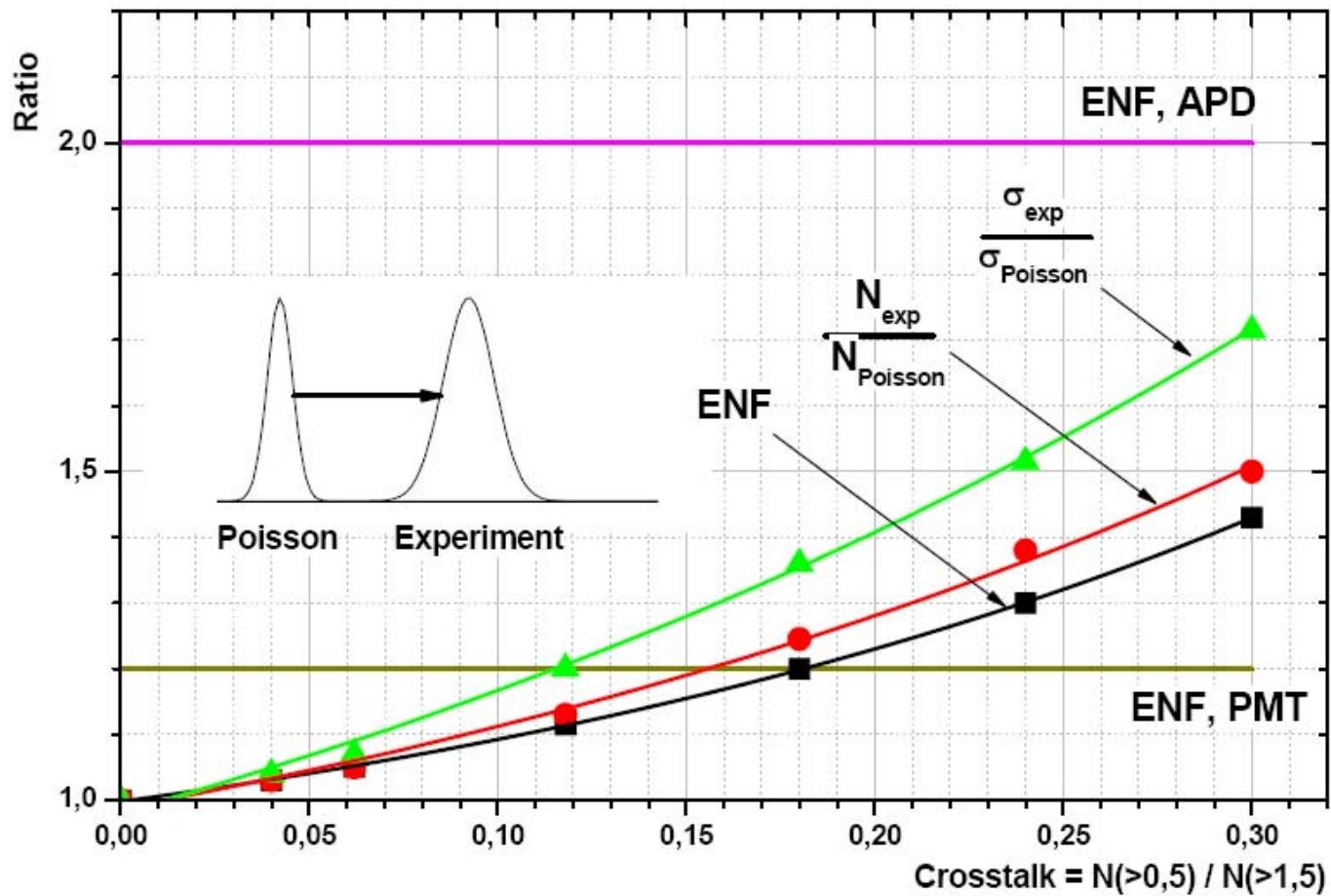
~ 7 Ohm
+shaper
FWHM
2,5ns

→ The lifetimes of trapped electron are mostly rather small:
less than ~100 ns

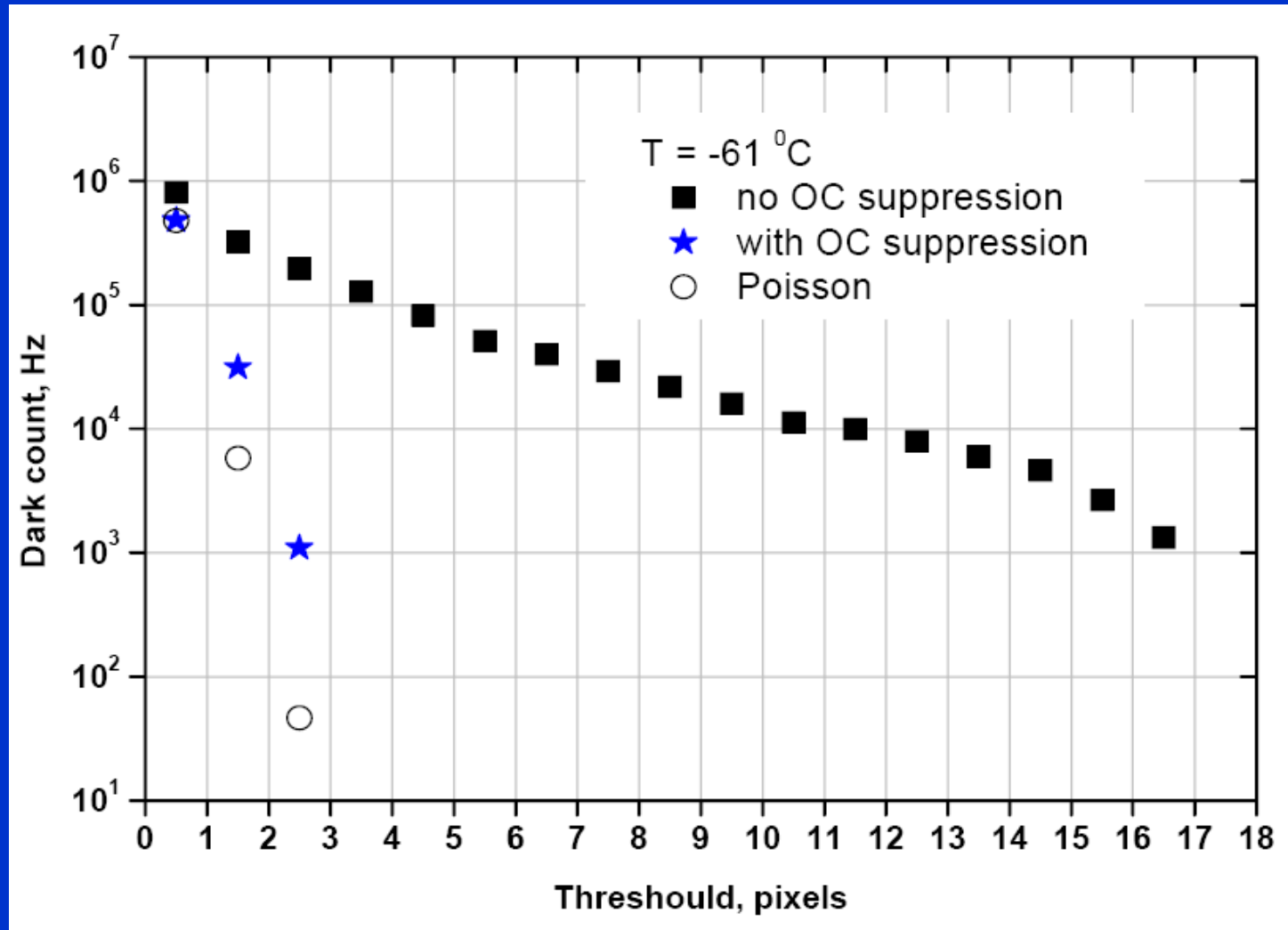


Therefore a single pixel recovery time $R_{quench} \times C_{pixel}$ should not be not very small and recommended at level of .5-1 mks

→ Even for high Gain x PDE the Afterpulsing has to be small enough:
 $AP(\text{Gain} = 10 \cdot 7) = \sim 1\%$
for recovery time of $> 500 \text{ ns}$

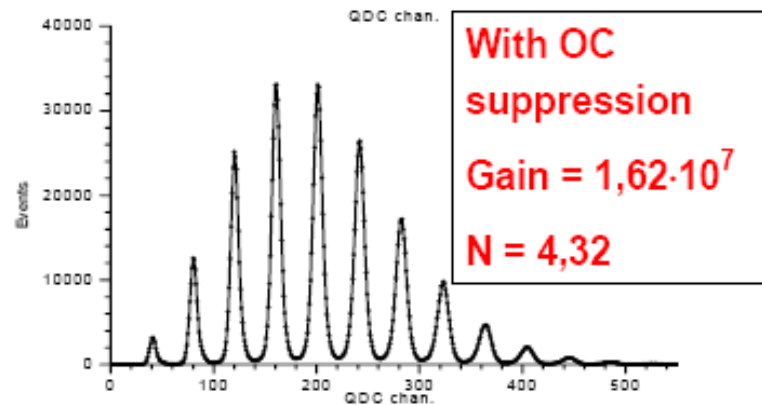
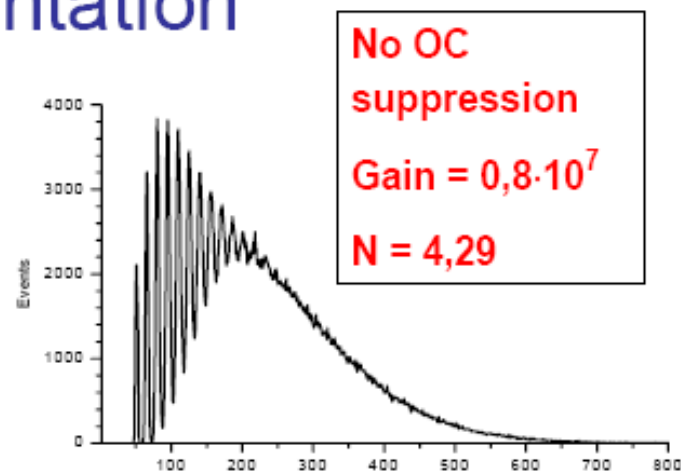
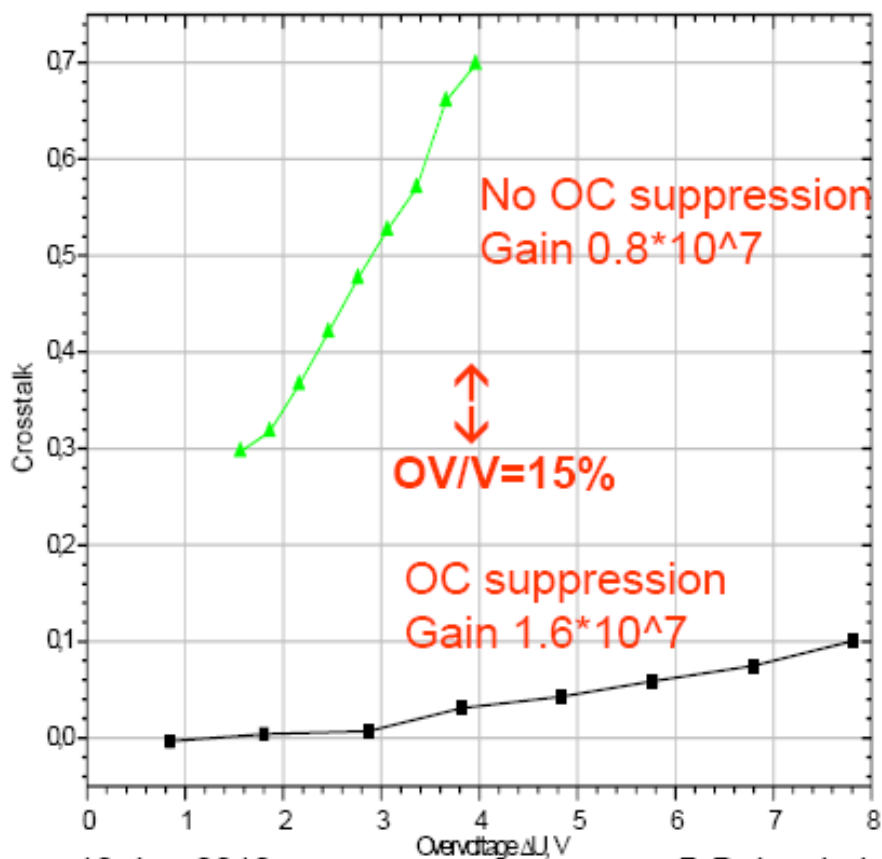


The cross-talk effect (optical coupling: OC) has a major impact in self-trigger schemes, it can prohibit obtaining a low threshold setting



Prompt OC suppression using Si damaged by ion implantation

(Patent pending)



SiPM $1 \times 1 \text{ mm}^2$

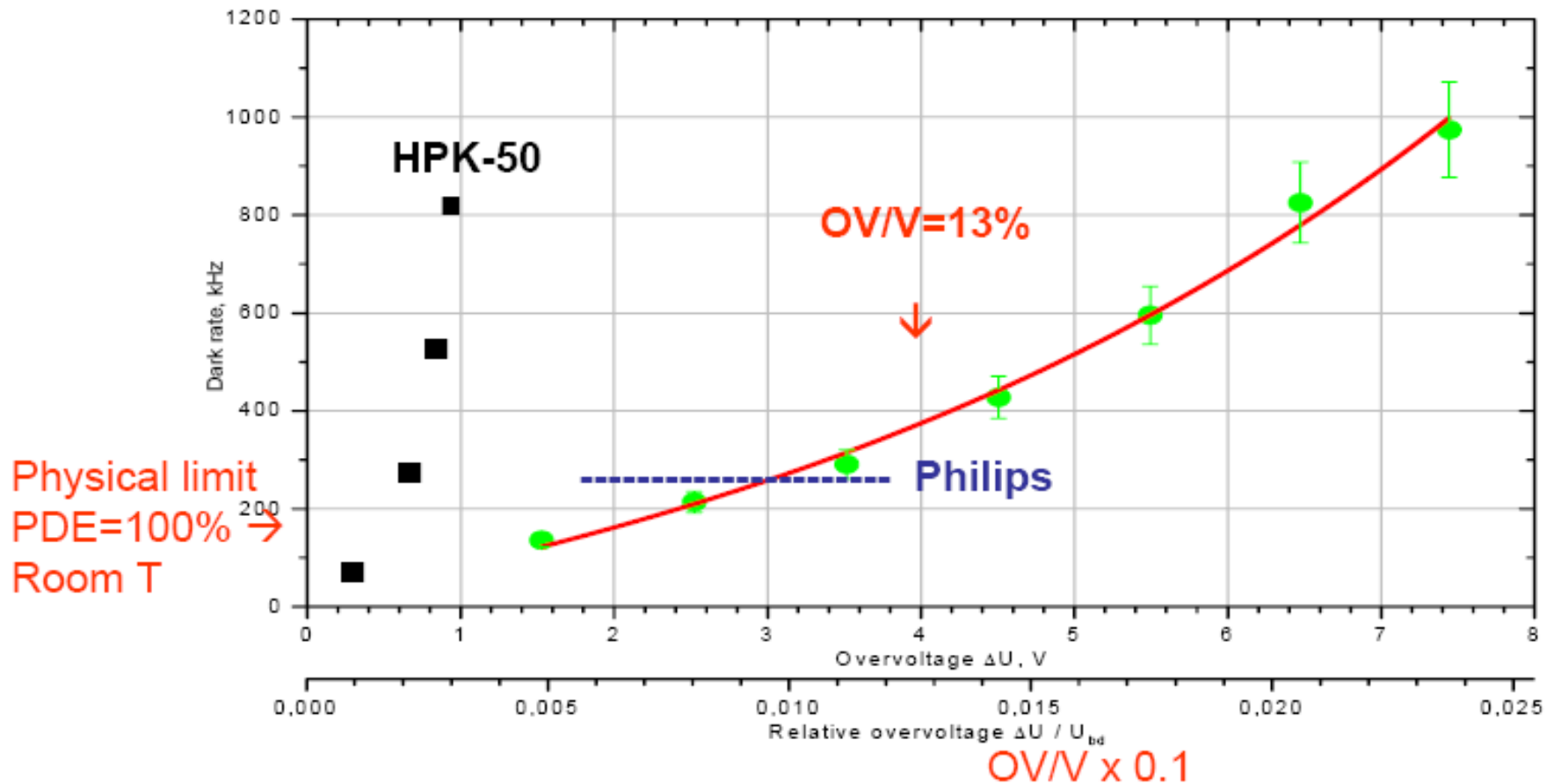
10-Jun-2010

B.Dolgoshein Silicon PM

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Dark count rate: SiPM 1x1mm², OC=4%, AP=1% room temperature

SiPM 1x1 mm² P on N, (pixel size size 100x100 μm) with OC and AP suppression



Digital SiPM

→ single pixel dark count rate is lower by factor of 1.5-2 (~physical limit)
→ digital output is more convenient for system integration

→ PDE loss (filling factor is less due to electronics on chip)
→ problems with Optical Crosstalk and Afterpulsing have to be solved

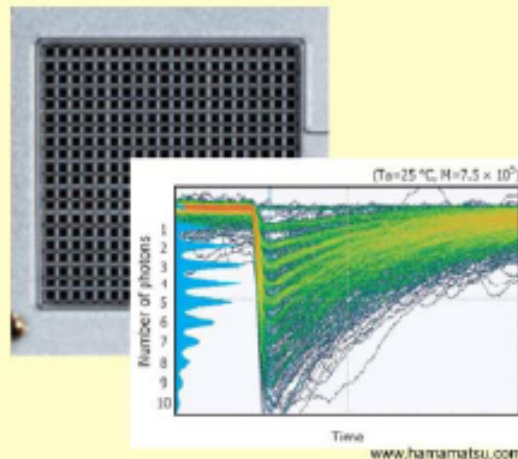
Fabrication cost?

PHILIPS

Digital Photon Counting – The Concept

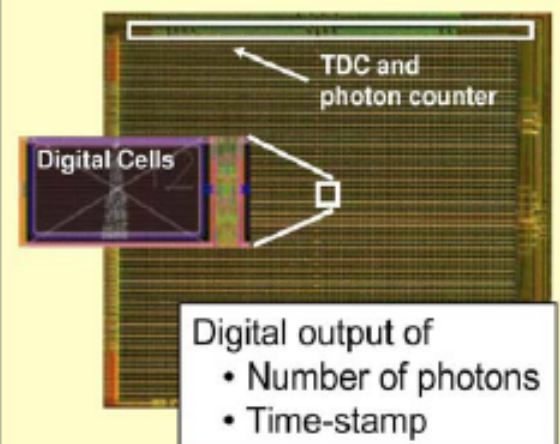
Intrinsically, the SiPM is a digital device: a single cell breaks down or not

analog SiPM



Summing all cell outputs leads to an analog output signal and limited performance

digital SiPM (dSiPM)

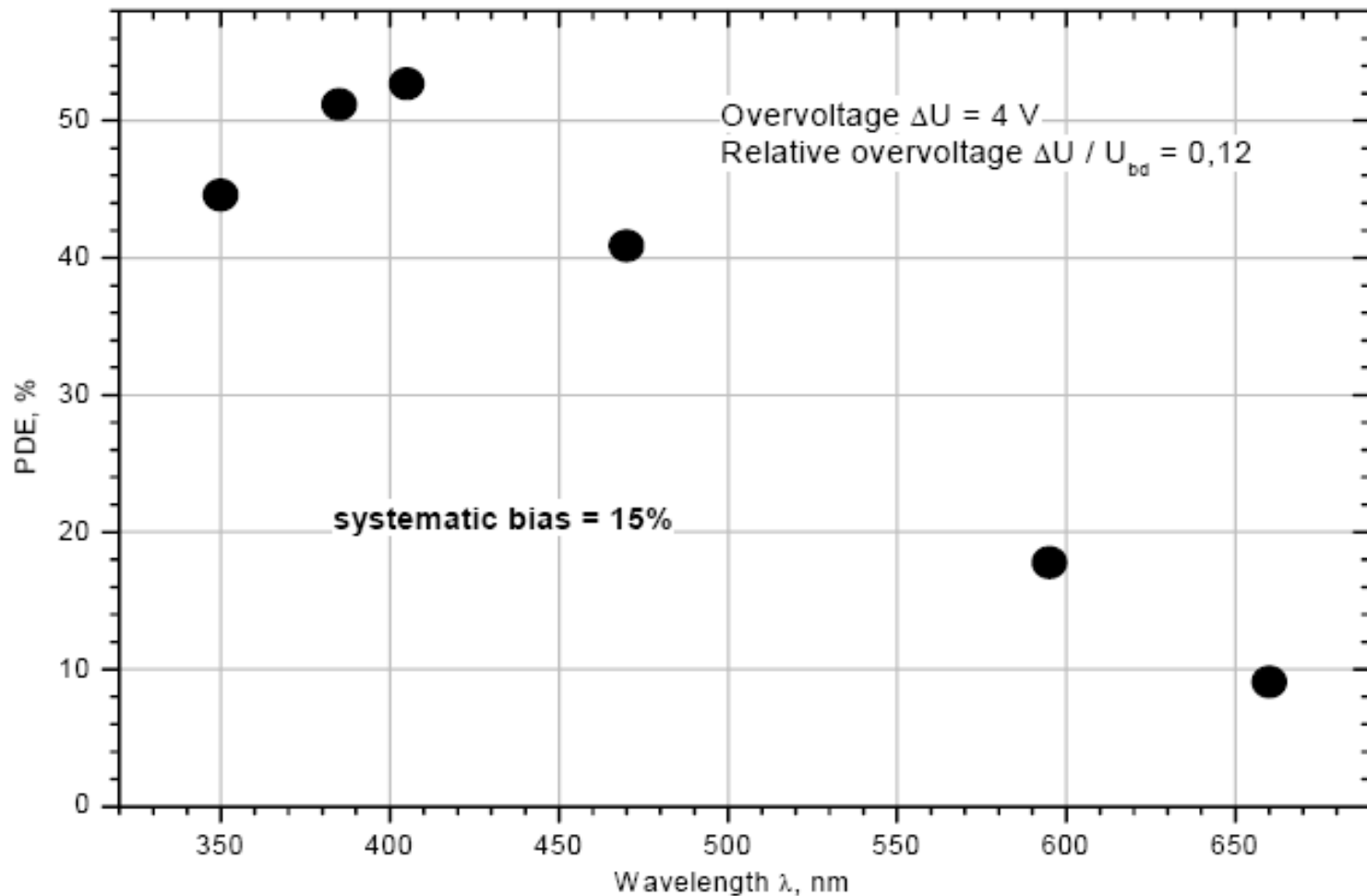


Integrated readout electronics is the key element to superior detector performance

PDE ,SiPM p on n,3x3 mm², OV/V=12% +OC suppression

Test-product of PEI, to become soon commercial product

SiPM 3x3 mm² P on N (pixel size 100x100 μm, geom. eff. = 0,6) T = -50 °C



Conclusions

- In a time scale 1-2 years from now one can buy SiPMs with outstanding characteristics, probably from several manufacturers.
- Their sizes could span 1-10 mm.
- SiPM cost will be reduced due to the availability of full CMOS designs. Several USD per mm² is not unrealistic.
- They could offer PDE of 60-65 %, x-talk < 1% and low temperature and voltage dependences.
- These devices are going to substitute classical PMTs and APD in many applications, including those in physics instrumentation in, for example, nuclear medicine (time-of-flight PET,...).
- Realistic candidates for SST telescopes, especially for 2-optical element designs