



# ELEC @ SST

General considerations  
and recommendations



# General wishes of ELEC WP

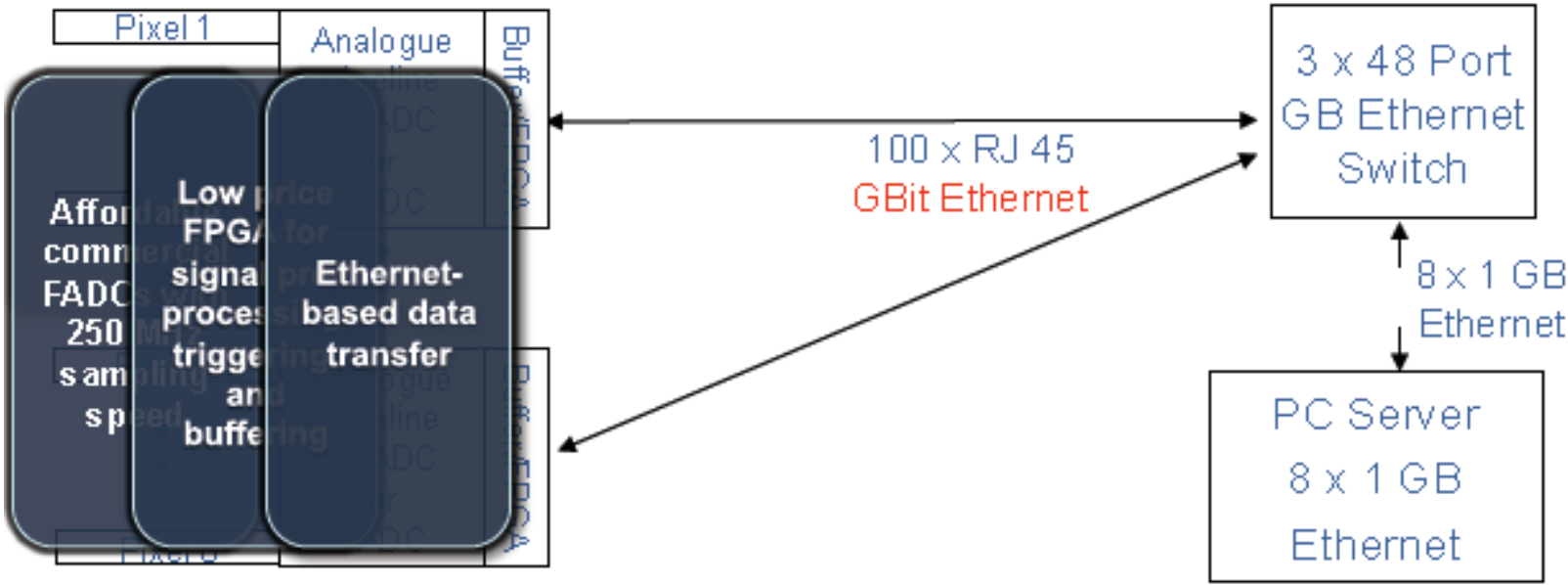
- Good performance to achieve physics goal
- homogenous design across various telescope size; without losing performance.
- Same electronics, or at least similar components if possible, even with other technical groups, prefer homogeneous devices (hardware and software).  
Lower the cost also for maintenance
- low electric power consumption, man power on site ...
- Easy to maintain (accessible in safety condition)
- Known technology
- Can be upgrade (15-20 years of operation is a long time)

# General considerations

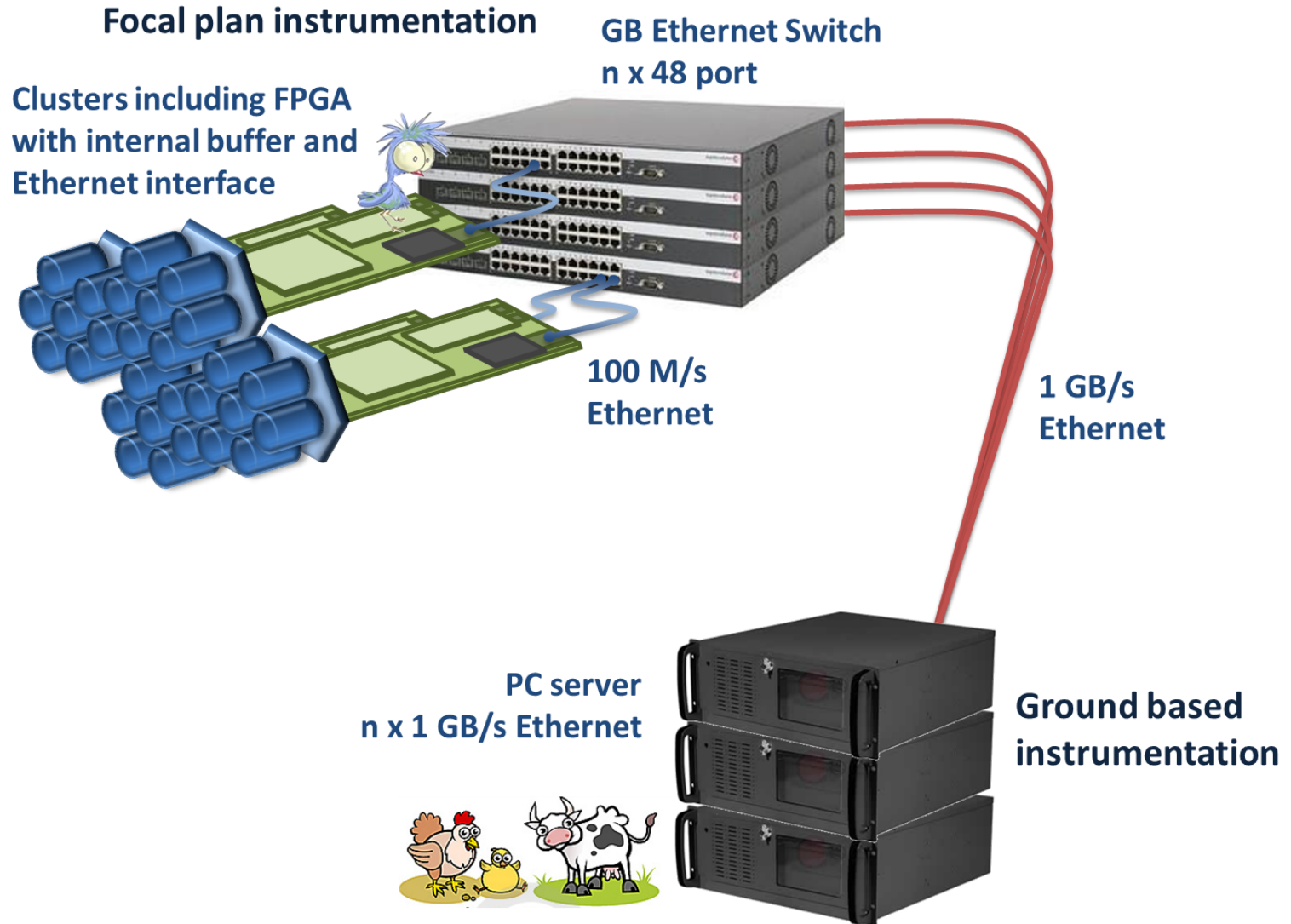
		FE	L0	L1
1	D	DESY/MPIK Heidelberg/AIT/ETH/U. Zurich/AGH/Krakow/Leeds.		
2	A/D	Pisa	Zeuthen	Zeuthen
3	A	Pisa	IFAE-UAB	UM/CIEMAT
4	A	IRFU/LPNHE/LPTA/ICC-UB	IFAE-UAB	UM/CIEMAT
5	A/D	CTA-US (AGIS)		
6	-	-		

	DAQ	Central Trigger
7	MPIK Heidelberg	APC
8		Iowa state U.

# General considerations



# General considerations



# General considerations

	Electronics		
Power consumption (all camera)	< 3 W/channel		SITE
Power consumption (per camera)	< 1 W/channel		SITE
Band width before digitalization	> 300 MHz		MC
Sampling rate	250 MHz to 2 GHz		MC
Dynamic range vs telescope size	5 000 photo-electron?		MC
Single photo-electron resolution (electronics noise)	pic/valley > 1.5		FPI
Dead time at 10 kHz	< 5 %		ELEC
Maximum acquisition rate (1-fold)	10. kHz		MC
stereoscopy	(20% of GHz network)		ELEC
crosstalk for the electronics	< 0.1%		ELEC
Trigger region size (sector or super cluster)	14-28 pixels (investigate more)		MC
Trigger window as function of telescope size	2 – 5 ns		MC
Time stamp accuracy for telescope	< 80 ns		DATA
Local time stamp accuracy in case of autonomous clusters	< 15 ns		ELEC
Response stability of operation	signal calibrated at 2%		ELEC
Maximum temperature	< 50°C		FPI
Temperature stability under acquisition	Assuming < +/- 1°C during a run of 25 mn		FPI
Data needed to be saved on disk			MC
LST	Sampling		
MST	Charge or parameterization (? Mb/s)		
SST	Charge		
Data reduction			MC
Pixel readout window			MC
Full camera readout window as function of the energy			MC
Manpower for characterization/monitoring	less than 2 persons		
Manpower for operation	Low maintenance		
Cost (from PM to network)	? € / channel vs telescope size		MC
Peripheral infrastructure needed			

# SST should achieve

- ~32 telescopes of about 1 000 channels

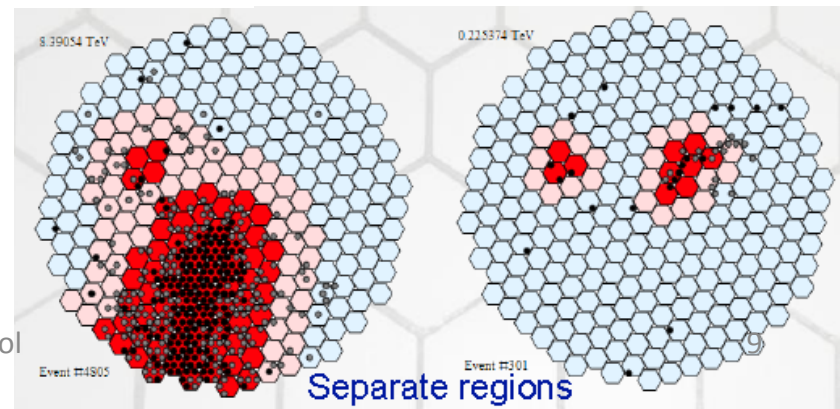
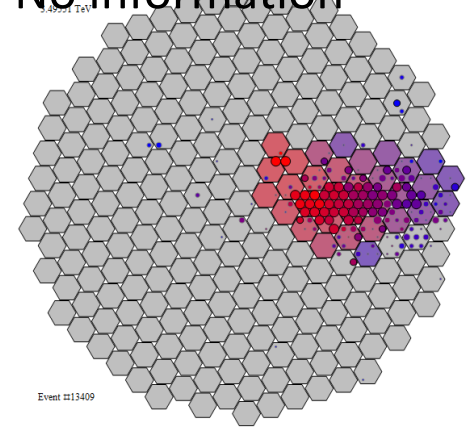
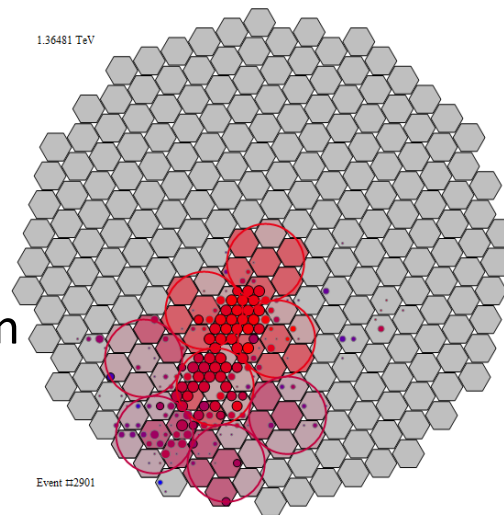
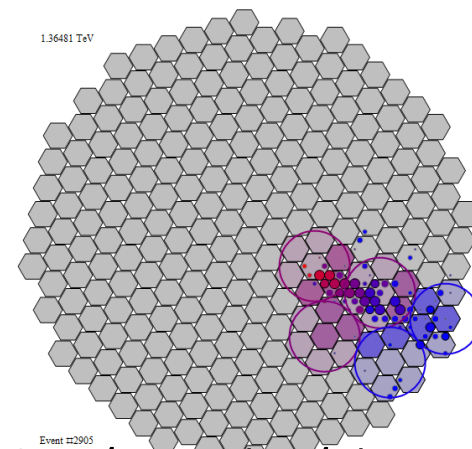
# SST should achieve – data flow

- $1\,000 \text{ pixels} \times 1 \text{ kHz} \times 16 \text{ bits} = 16 \text{ Mb/s}$
- $1\,000 \text{ pixels} \times 1 \text{ kHz} \times 16 \text{ bits} \times 2 \text{ gains} = 32 \text{ Mb/s}$
- $1\,000 \text{ pixels} \times 1 \text{ kHz} \times 16 \text{ bits} \times 2 \text{ gains} \times 50 \text{ ns} = 1,6 \text{ Gb/s}$
- $100 \text{ tel.} \times 1\,000 \text{ pixels} \times 1 \text{ kHz} \times 16 \text{ bits} \times 2 \text{ gains} \times 50 \text{ ns} = 160 \text{ Gb/s}$
- 600 TBytes/night



# General considerations

- Data reduction
  - Data compression algorithm 16- $\rightarrow$ 8 bits/word ( $\sim 1/2$ ). No information lose.
  - zero suppression
  - readout selected region

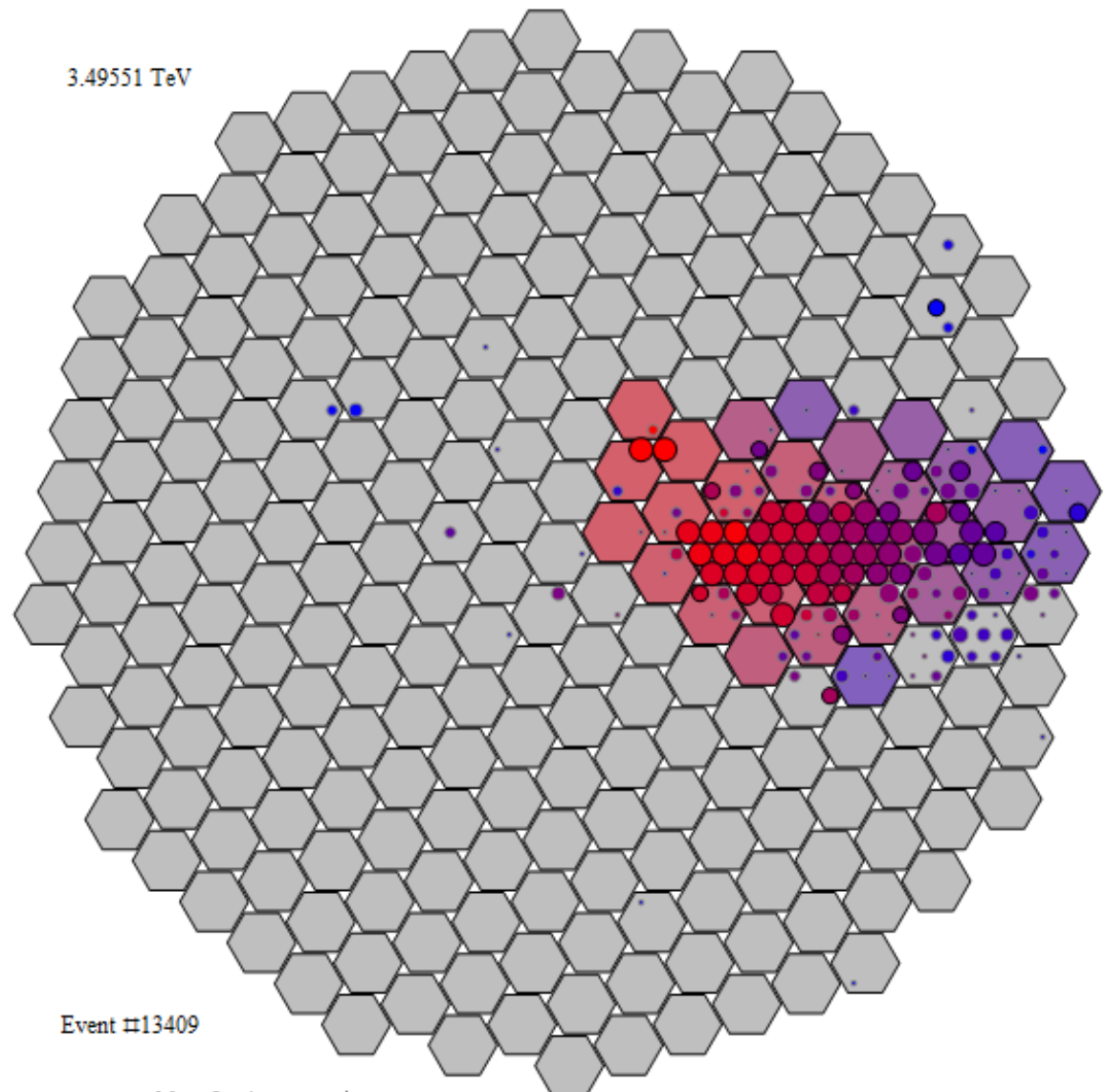


# SST should achieve

- ~32 telescopes of about 1 000 channels
- High energy : few x 100 TeV
  - Large dynamics range
  - sophisticate readout and trigger

# General considerations

- Data readout
  - MC Trigger studies



# SST should achieve

- ~32 telescopes of about 1 000 channels
- High energy : few x 100 TeV
  - Large dynamics range
  - sophisticate readout and trigger
- reliability
  - Complete monitoring and slow control system
  - Accurate calibration
  - Compact design allows full construction and test in laboratory  
Only power & network connection
  - Modularity facilitates the installation and maintenance

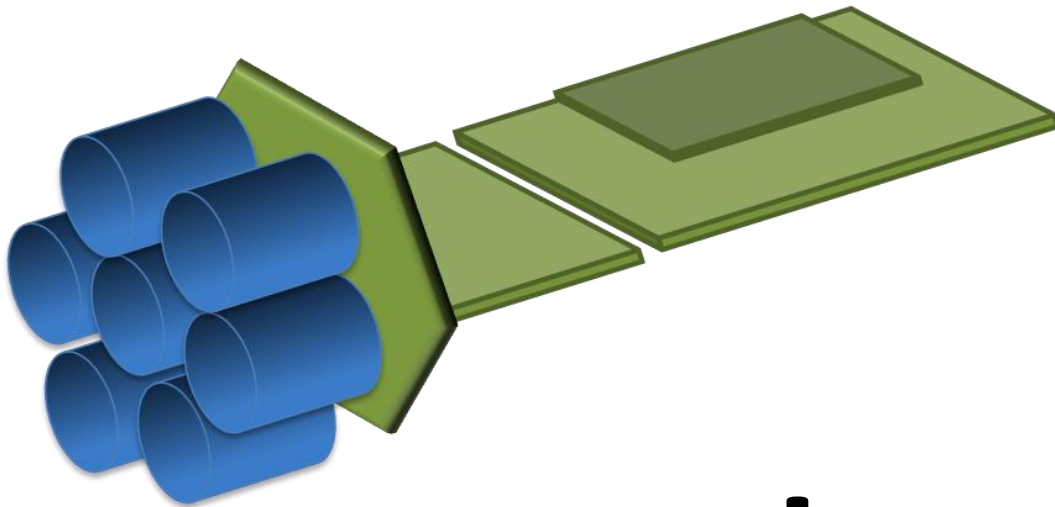
# General considerations

- Something prepared and tested at home institutes or manufactories



# General considerations

- Accessible (at the ground level ?) and easy to maintain system (modular).

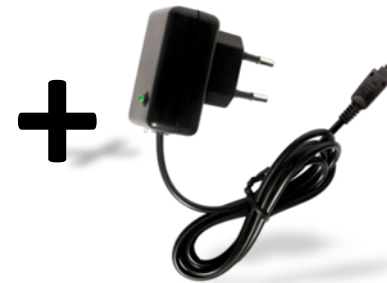
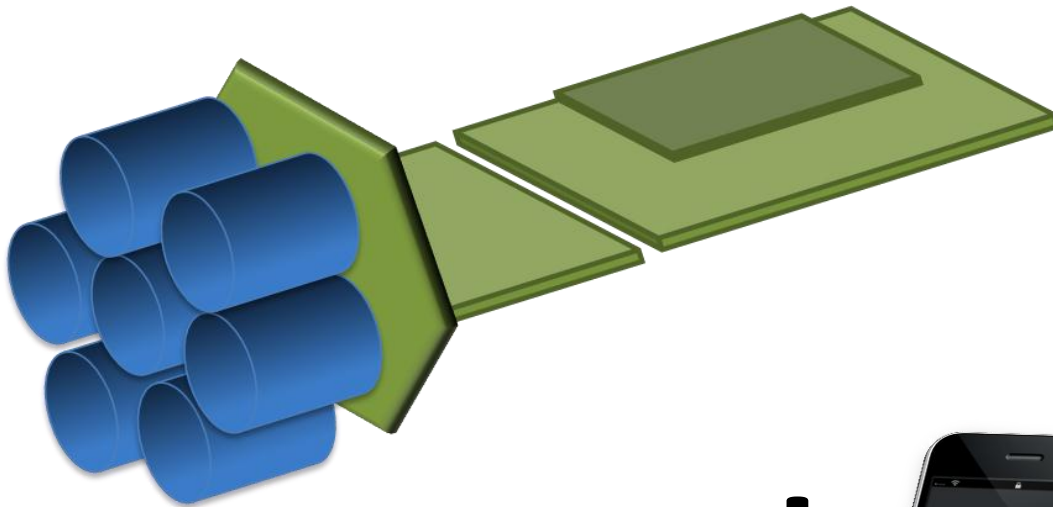


- standardized interfaces and software



# General considerations

- Accessible (at the ground level ?) and easy to maintain: system modular.



- standardized interfaces and software



CTA Store



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# SST should achieve

- ~32 telescopes of about 1 000 channels
- High energy : few x 100 TeV
  - Large dynamics range
  - sophisticate readout and trigger
- reliability
  - Complete monitoring and slow control system
  - Accurate calibration
  - Compact design allows full construction and test in laboratory  
Only power & network connection
  - Modularity facilitates the installation and maintenance
- **Capability to evolve:** maybe less critical for this telescope size.



# Figures of merit

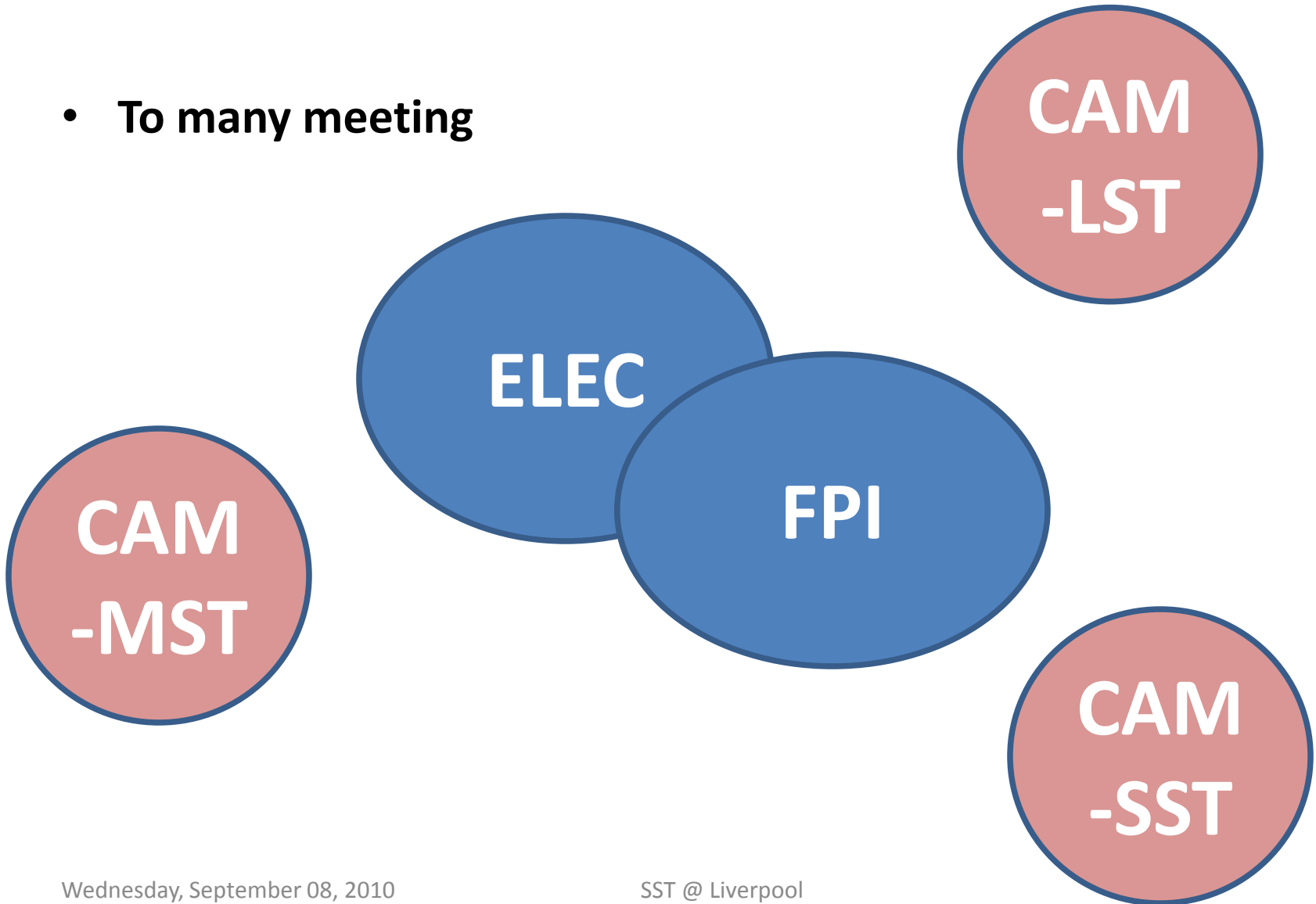
	Telescope size	SST	MST	LST
Dead time		2	1	1
Linearity		1	1	1
Dynamics		1	1	2
Crosstalk		2	1	1
Electronics bandwidth		2	1	1
Trigger		3	2	1
Readout		1	2	1
Power consumption		2	1	3
Power dissipated		2	2	1
Modularity (evolution)		3	2	1
Modularity (maintenance)		1	1	2
Access to camera		1	1	2
Weight		3	2	1
Upgrade		2	2	1
Cost/channel		1	1	2

# CAM-SST should achieve

- constraints and need for High Energy
- specifications for the electronics
- cost model of a full telescope
- central triggering ?
- installation, maintenance model

# Remark

- To many meeting



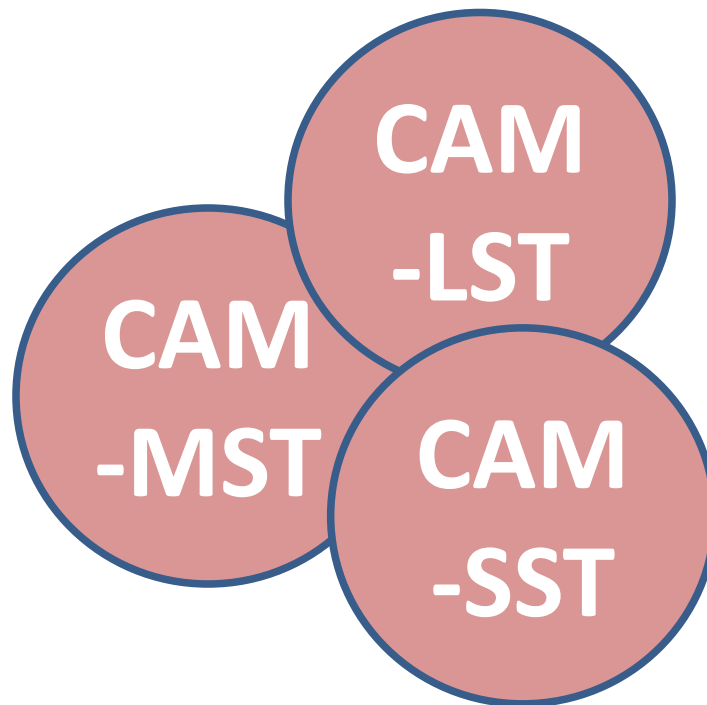
# Remark

- **Inhomogeneous system**
- **expensive**



# Remark

- **Back to the futur**



**End**