

First G-APD Cherenkov Telescope

The FACT Collaboration:

- TU Dortmund, Germany
- ISDC, Geneva, Switzerland
- EPFL, Lausanne, Switzerland
- University of Würzburg, Germany
- ETH Zurich, Switzerland

08.09.2010

Isabel Braun
ETH Zurich

for the FACT Collaboration

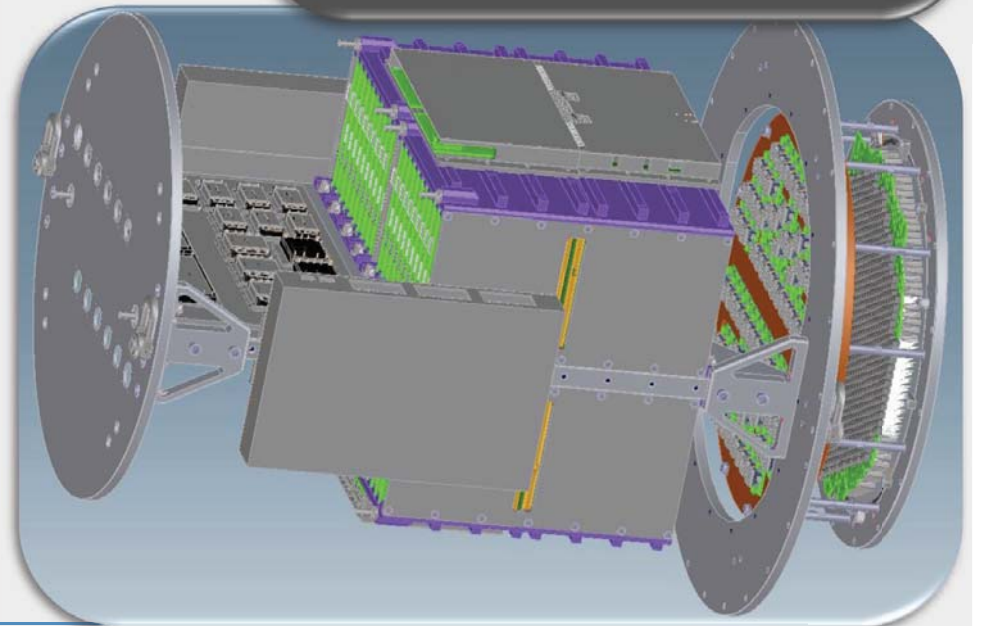
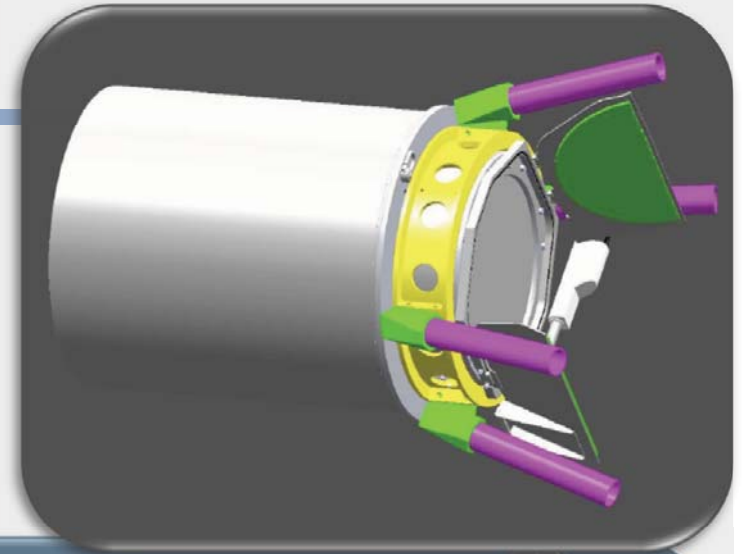
Telescope

- Former Hegra CT3
- CT1 mirrors
 - spherical mirrors with hexagonal shape
 - diameter (inner circle) 60.5 cm
 - Al honeycomb, weight ~ 6 kg
- refurbished @ LT ultra
 - diamond milled
 - mean focal length (4.890 \pm 0.008) m
 - av. spot size @ 2f: 16.0 mm² ()
 - av. reflectivity $\sim 90\%$ (300-500 nm)
 - maximum @ 450 nm
- Drive (downscaled MAGIC)



Camera Mechanics

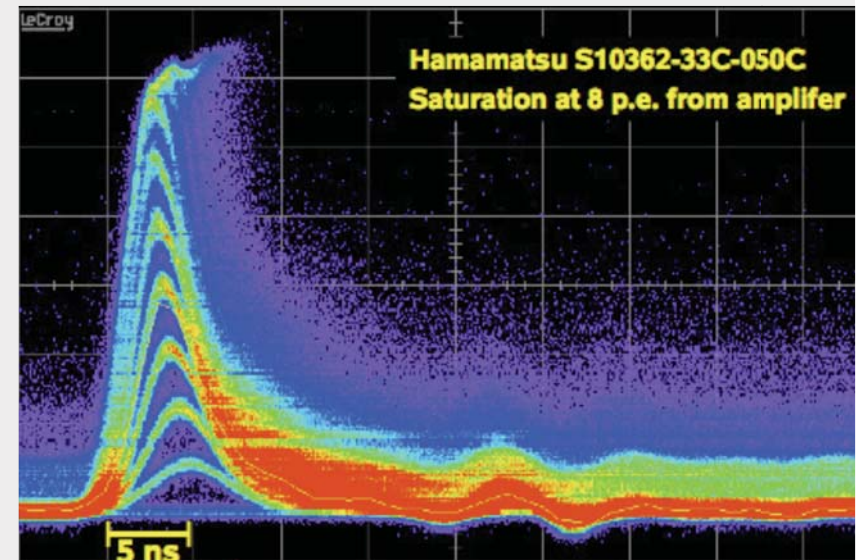
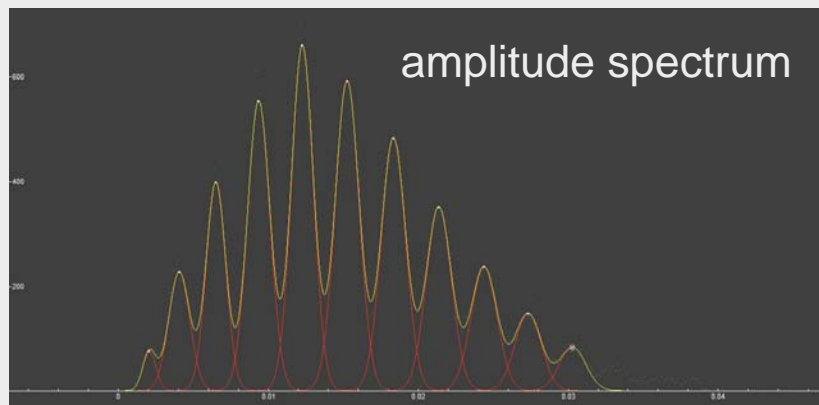
- length 812mm
- diameter 532mm
- weight approx. 100 kg
- water cooled
- thermal separation between sensor and electronics compartment
- housing / connectors must be water tight (IP67)



G-APDs

- 1440 pieces
- Hamamatsu MPPC S10362-33-50-C
- active area $3 \times 3 \text{ mm}^2$
- $50 \text{ }\mu\text{m}$ cells (less saturation than $100 \text{ }\mu\text{m}$ version)
- package $5.9 \times 6.6 \text{ mm}^2$

- peak photo detection efficiency 33%
- operation voltage $\sim 70 \text{ V}$
- gain $7.5 \cdot 10^5$

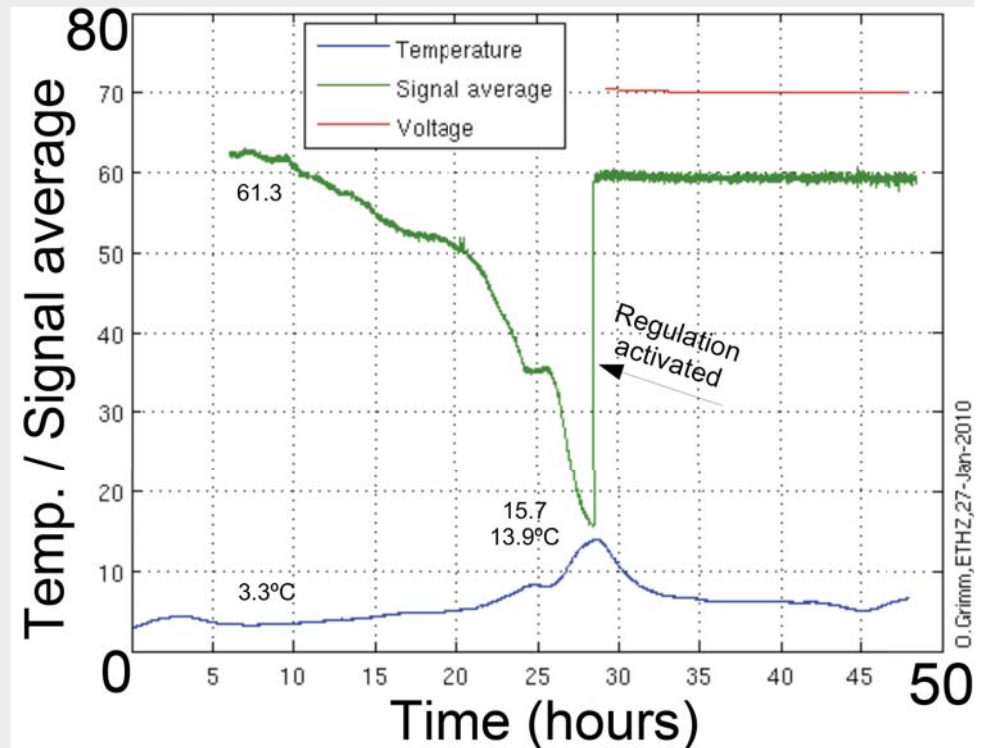


Bias Voltage Feedback System

- Temperature influences breakdown voltage -> overvoltage
- => changes photo detection efficiency, crosstalk probability, gain



- active HV Feedback!
- light pulse signal stabilized
- signal stable within $\sim 0.5\%$



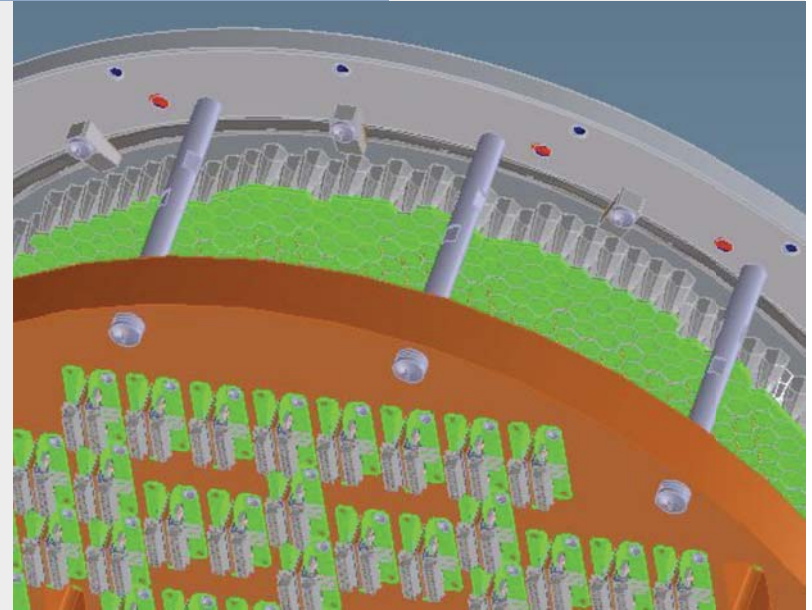
Solid Light Guides

Design:

- hexagon (9.5mm flat to flat) -> square (2.8mm sides)
($C_{\text{Area}} = 8.7$ rel. to G-APD)
- height 19.939mm
- wall shape parabolic (in our case 2% better than Winston)
- optimized for $S(\text{inner } 20^\circ)/\sqrt{N}$

Material:

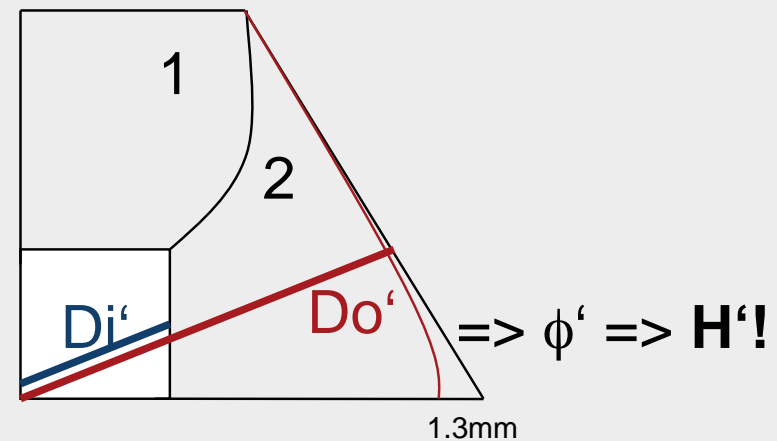
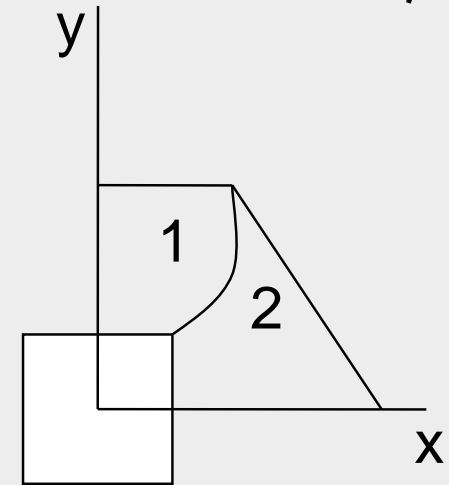
- UV transparent Plexiglas 7N
- injection moulding (IMOS Gubela GmbH)



Challenges

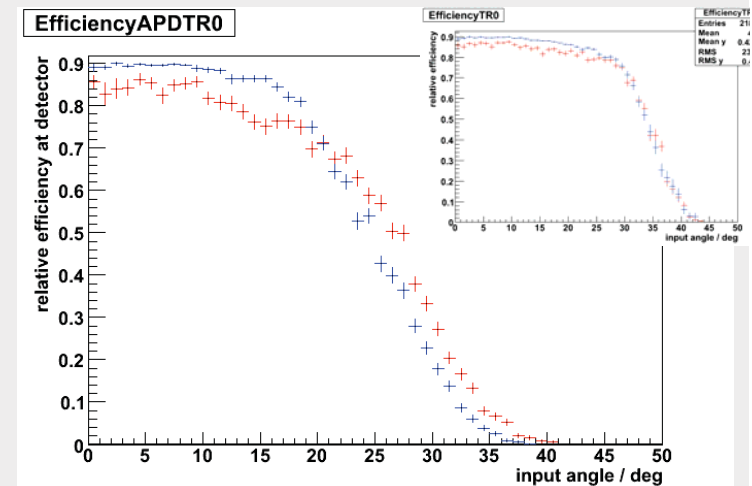
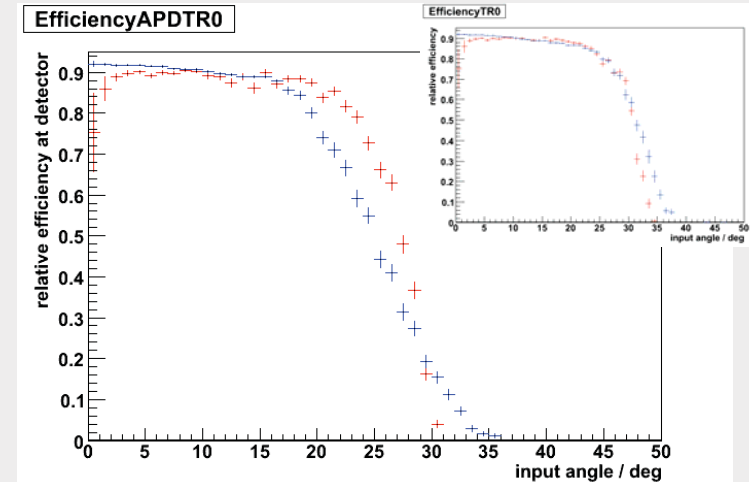
a few slides for Jim...

- Shape
 - type 1 without problem
 - edge (type 2) has changing input/output ratio, hence changing angle (OK) and changing height (not OK!)
- Compromise: Part-Winston
 - type 1: Winston, type 2: regular parabola
 - loose 2% efficiency
- Or back to full Parabola?
 - very first shape tested
 - UZH prototypes



3D Cone Types Compared

- Differences in 3D
 - Winston Cone not 'ideal' for non-center rays
 - **Winston edge** sharper than in parabola
 - consider difference at detector
- Parabola is better than its reputation!
 - in circ-circ Winston is superior
 - but already for square-square parabola wins!
- Options
 - full-Winston
 - part-Winston
 - full parabola ($z(F)=-1.5\text{mm}$)

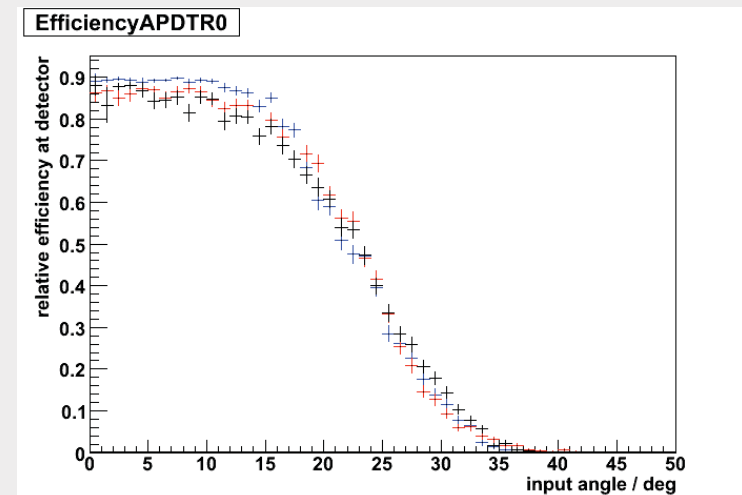
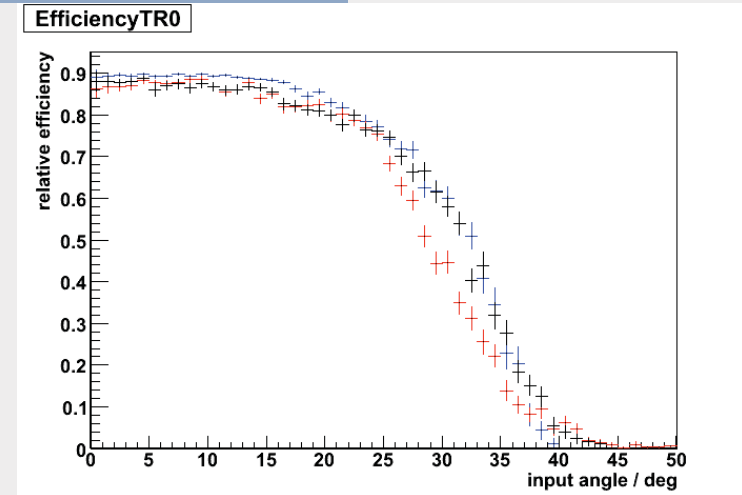


3D Cone Types Compared

Options

- full-Winston 85.0% / 79.0%
- part-Winston 84.9% / 76.9%
- full parabola 88.0% / 81.4%

- ... produced parabola after checks



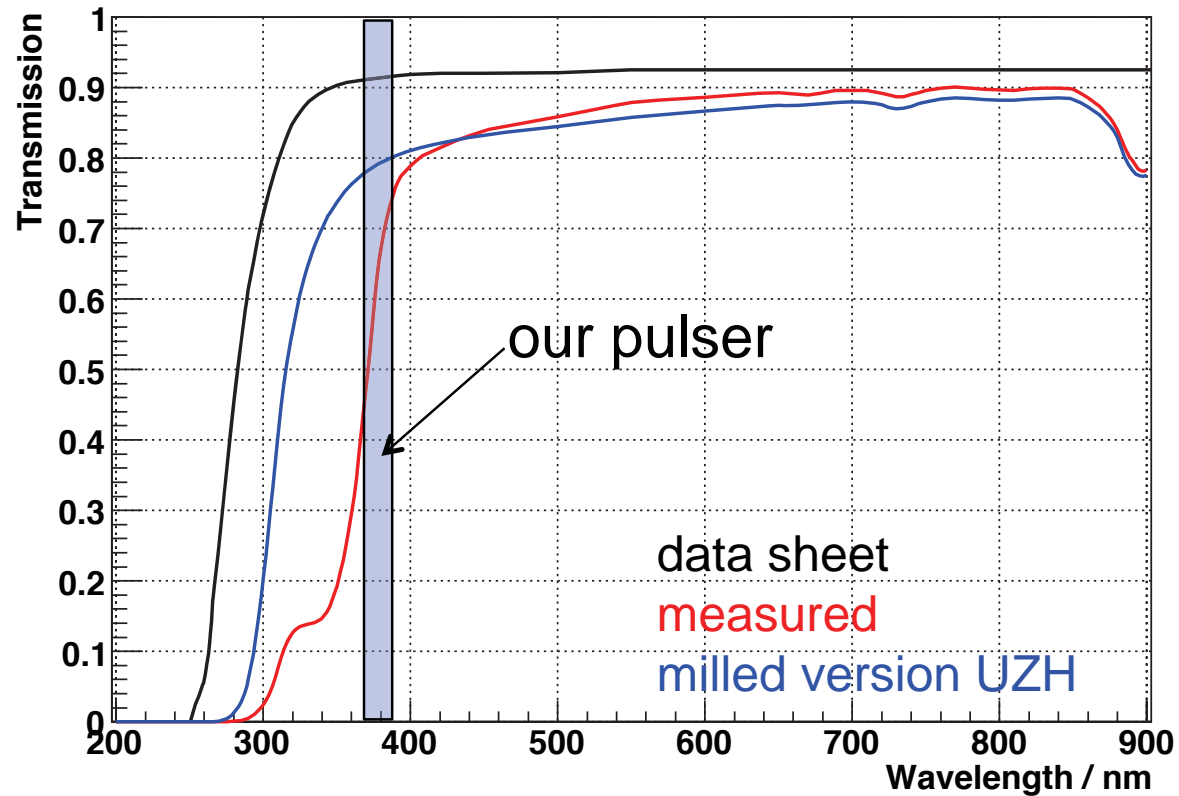
Prototypes arrived...

- Total efficiency for vertical light in first UZH and ETH measurements too low (expected 83-92% (roughness))
- angular behaviour as expected
- Inspection of the cone shows some surface structures (flow lines, injection hole)
- Transmission? Surface? Light coupling?

simulated loss at hole < 0.5%



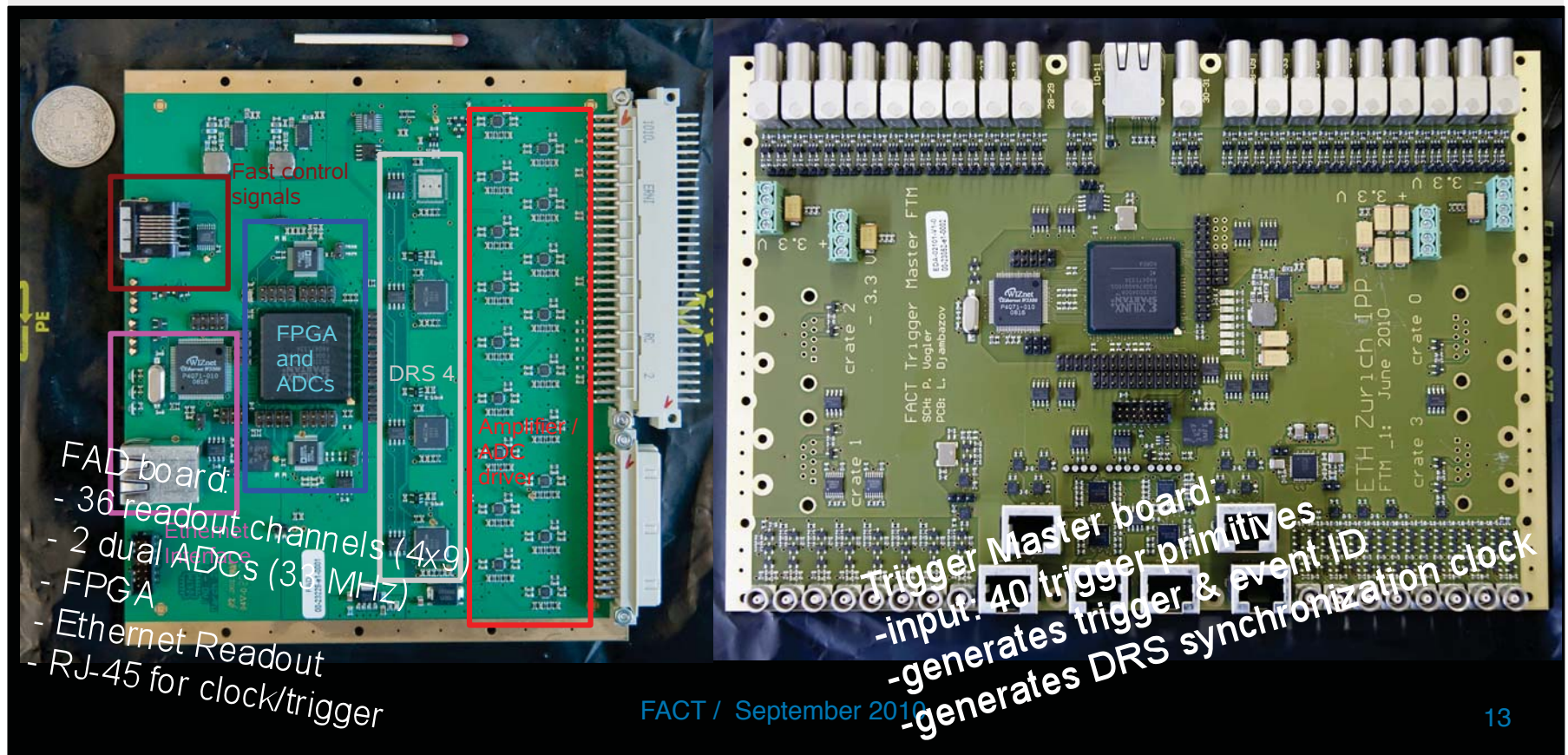
Measured Transmission



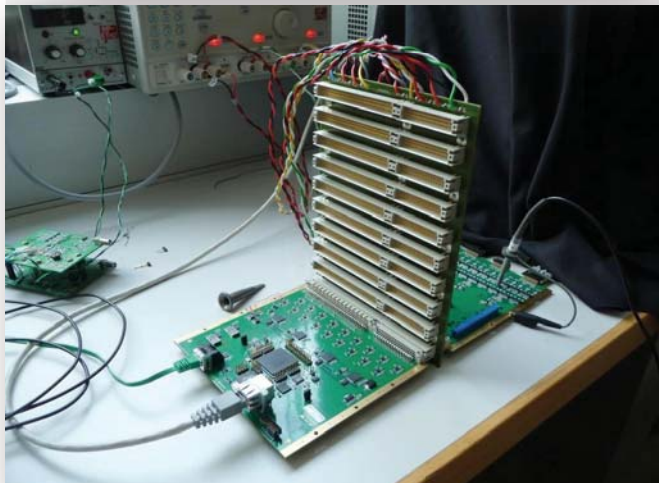
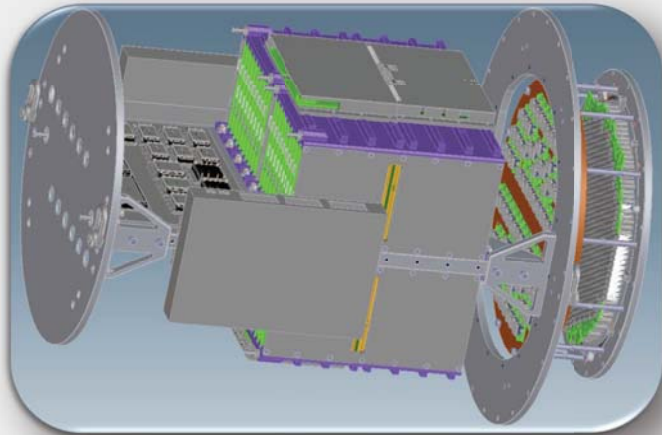
- apparently wrong material -> investigating...

Other Components

- Readout: DRS 4, probably @ 2 GHz
- Trigger: threshold on sum of 9 pixels
- Water Cooling: only electronics compartment, 1 kW



Status

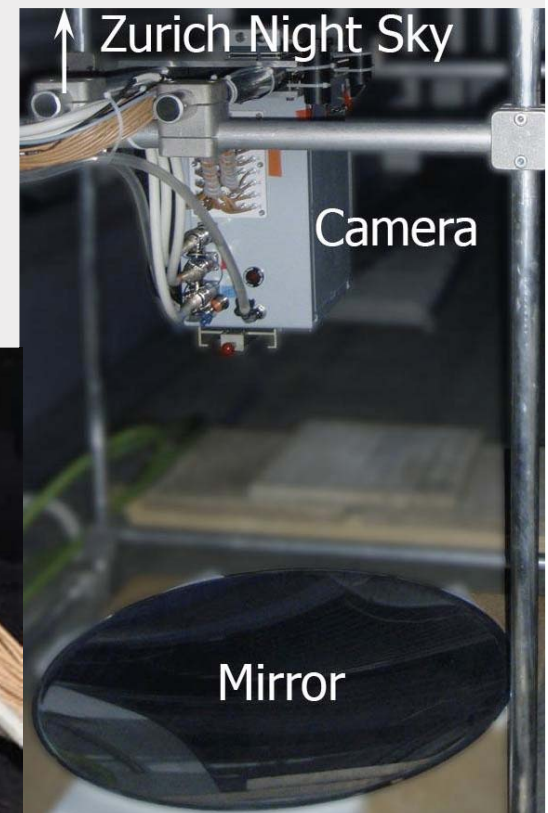
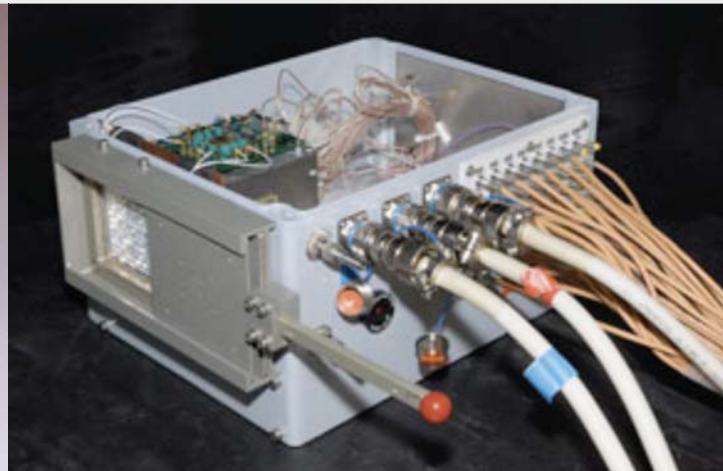
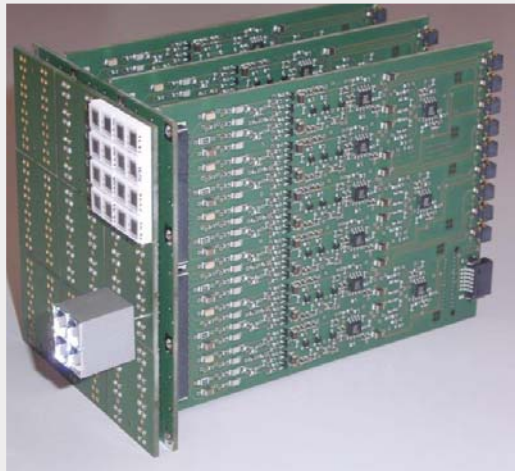


- **Telescope:**
 - mirrors and drive system ready for installation
- **Mechanics:**
 - design incl. cooling finished
 - fabrication of components started
- **Photo detection:**
 - all G-APDs available
 - light guide fabrication ongoing
- **Electronics:**
 - prototypes of all PCBs fabricated
 - tests ongoing

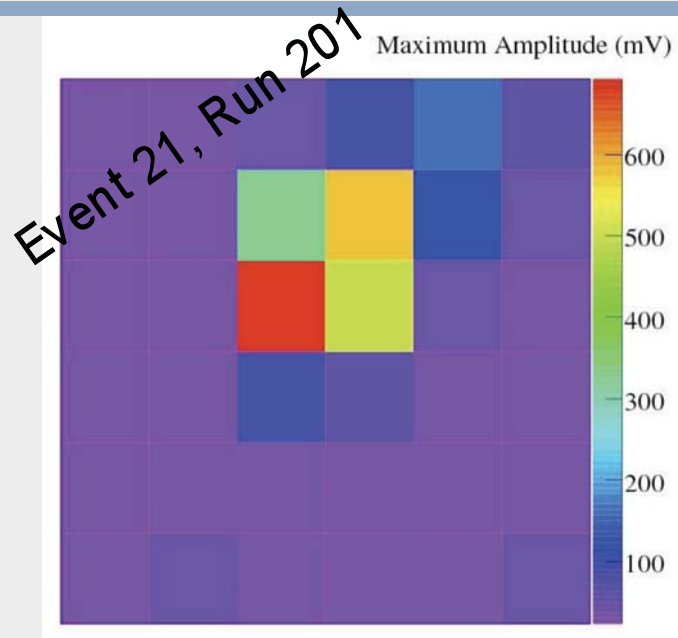


Prototype Measurements

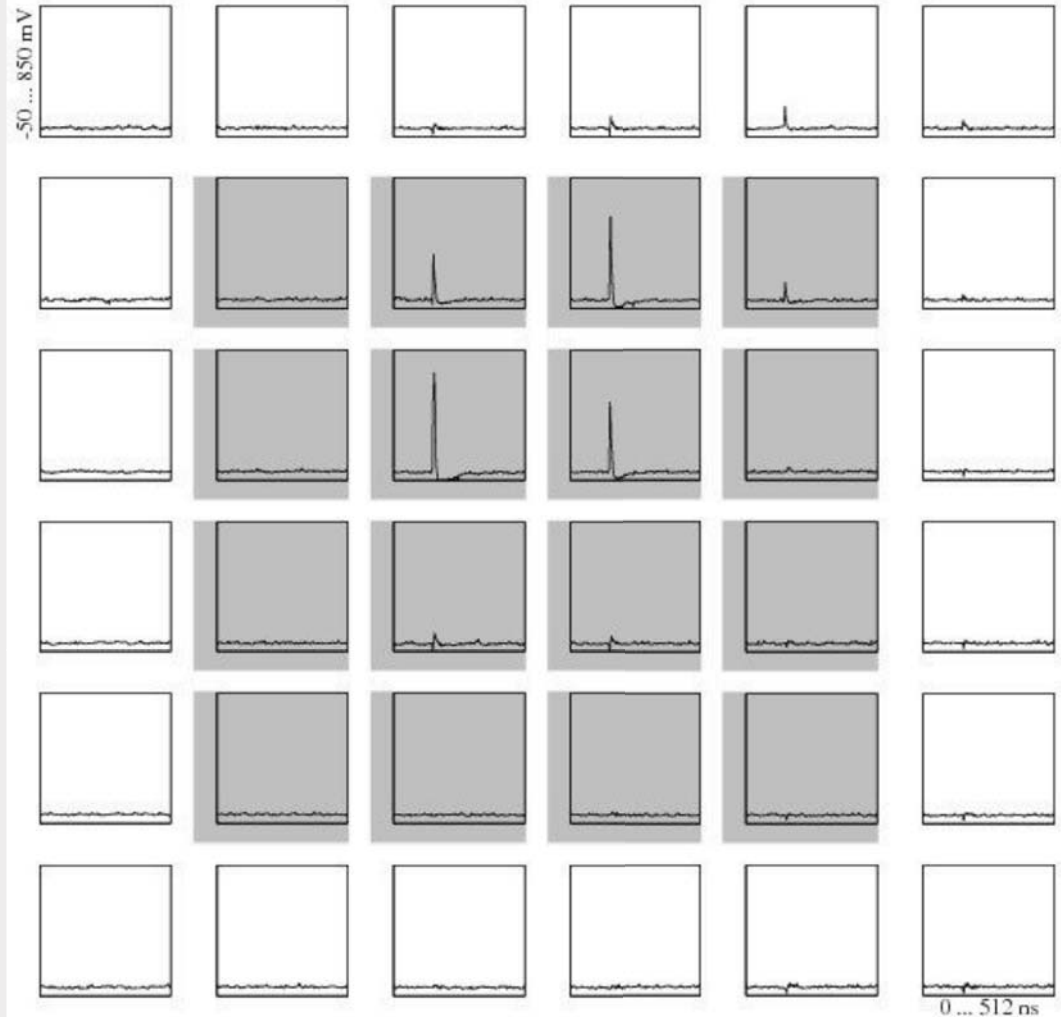
- June 2nd/3rd 2009 on roof of IPP in Zurich (~ 1 GHz NSB rate / pixel)
- 80 cm focal length
- 1° / pixel
- 144 G-APDs (4 G-APDs /pixel)
- simple light collectors (pyramidic, Vikuiti 3M)



Air Shower Images with G-APDs



- 25mV single pixel threshold (4 p.e.)
- Majority 4 out of 16
- 100 kHz trigger rate per pixel
- 0.01 Hz total shower rate
- 2 GHz sampling frequency DRS2



Summary

- 1440 channels (G-APDs)
- solid light concentrators (plexiglas)

- field of view 0.11° / pixel, 4.5° total
- integrated electronics
- ~ 100 kg!

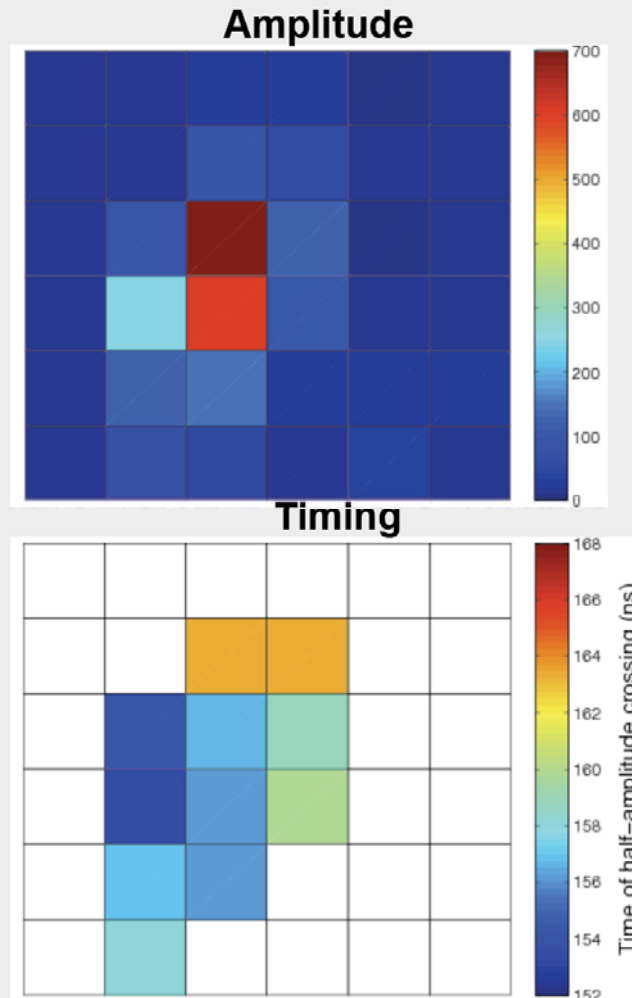


- 2009: prototype measurements with 36 pixels recorded air showers
- 2011: we will present Crab analysis
- will become part of DWARF physics program

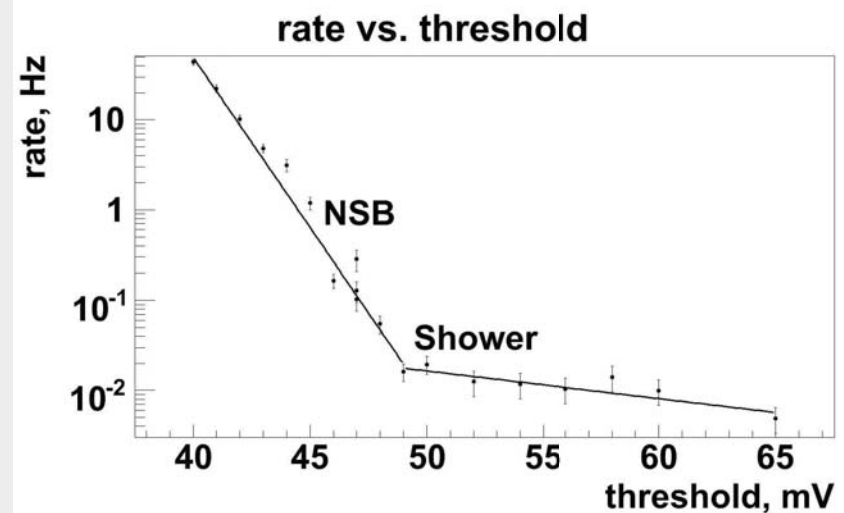
Backup



Air Shower Images with G-APDs

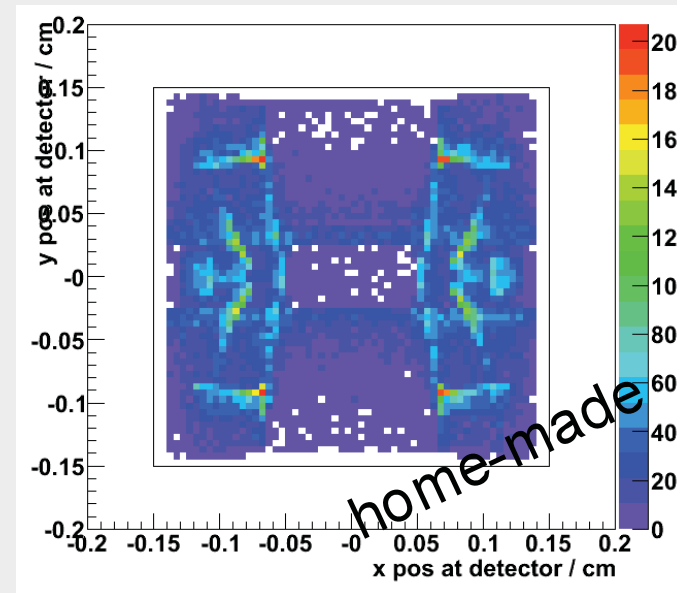
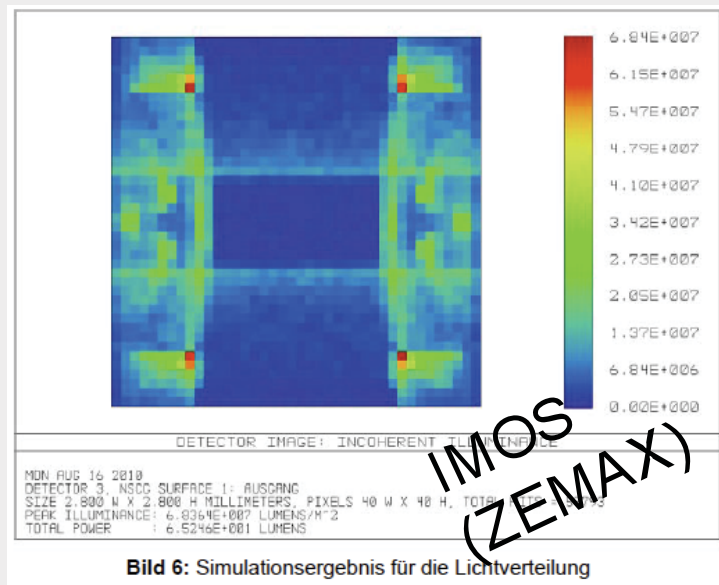


- G-APD Camera records first air showers
- rate scan behaves as expected



Possible explanations

- The light distribution for vertical incidence is not homogeneous => the G-APD has locally saturated?
=> maybe, but no influence on Poisson mean calculation!



- but: ETH light pulser operates at 380 nm...