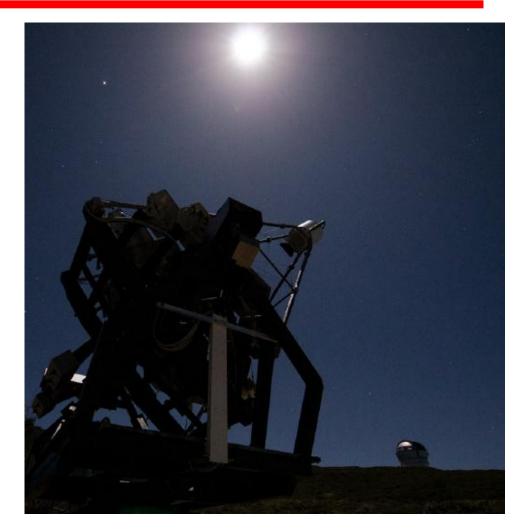
## SST Progress Report

#### Tim Greenshaw for SST group

- Davies-Cotton SST.
  - 7 m structure.
  - 4 m design.
- Progress with Dual Mirror designs:
  - ♦ SST-GATE.
  - ASTRI.
- Camera developments:
  - ASTRI camera and SiPMs.
- Calibration and pointing.
- Schedule.
- Summary.



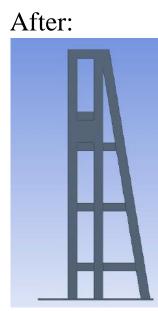
## Davies-Cotton SST – 7 m structure

Jacek Niemiec

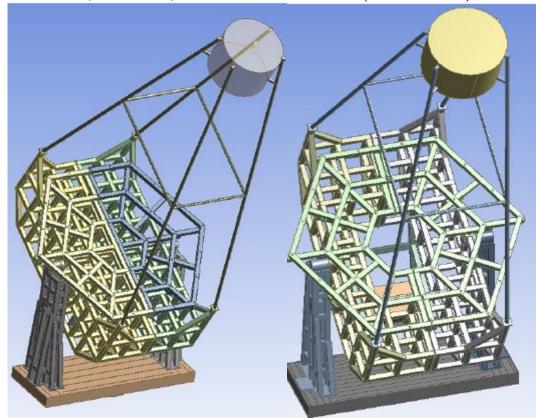
- Recommendation of SST review was that IFJ-PAN steel structure be simplified.
- Done for dish support structure, counterweight, dish and columns:

Before:





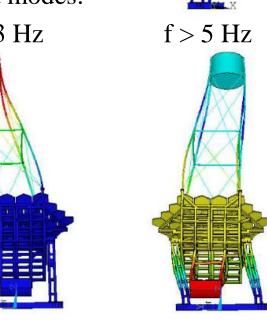
- Telescope:
- Before (27 tons):
  - After (24.4 tons):



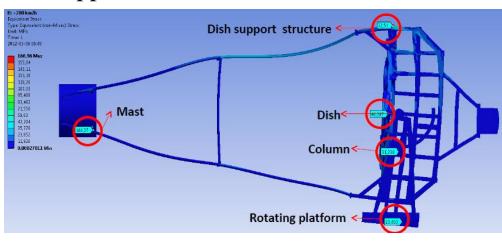
Costs of material and welding reduced.

### DC SST -7 m structure

- FEA of optimised structure completed.
- Lowest frequency oscillations 2.5 Hz at elevation of 60°:
- Next modes:
- f > 3 Hz



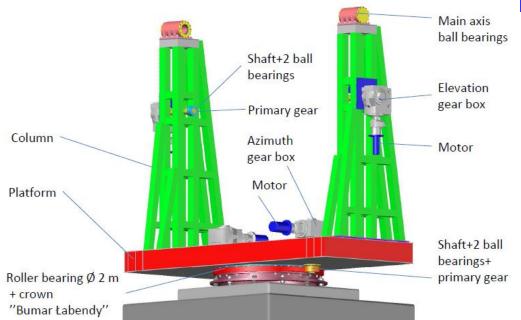
- Max. camera displacement under gravitational and 50 km/h wind load is 15 mm at an elevation of 30°.
- Less than 1/3 pixel size.
- Stresses under 200 km/h wind load in parking position are largest in dish support structure, 73 MPa:



• Steel yield strength 250 MPa.

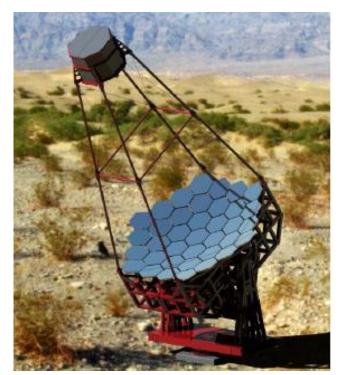
### DC SST -7 m structure

Detailed design of drive system performed:

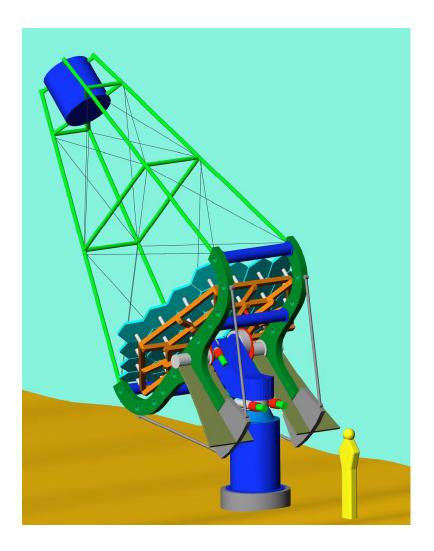


- Gear boxes, motors... chosen.
- Suppliers same as for MST in many cases.

- Design now considered complete enough to be shelved.
- Polish groups starting work on 4 m DC structure for which Swiss developing SiPM-based camera.

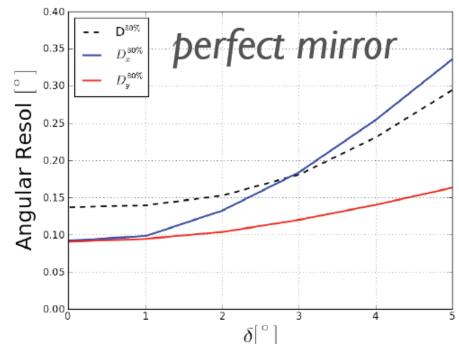


### 4 m DC SST – structure and mirrors



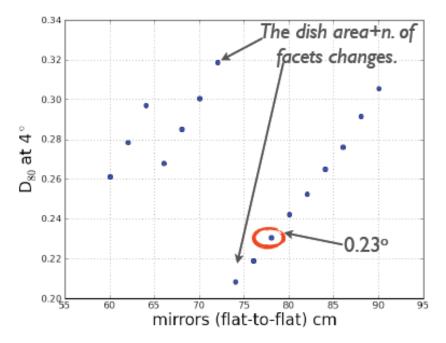
#### Teresa Montaruli, Jacek Niemiec

- Design DC telescope that has D ~ 4 m, focal ~ 5.6 m and 9° FoV.
- C.f. FACT, refurbished HEGRA CT3 telescope, D ~ 3.5 m, f ~ 5 m.
- Hexagonal mirrors of flat-to-flat size
  78 cm give required PSF:

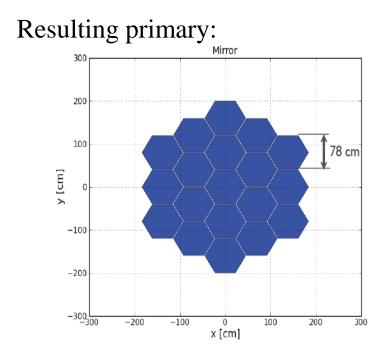


## 4 m DC SST – structure and mirrors

Facet sizes which give mirrors extending beyond D ~ 4 m result in poor PSF:

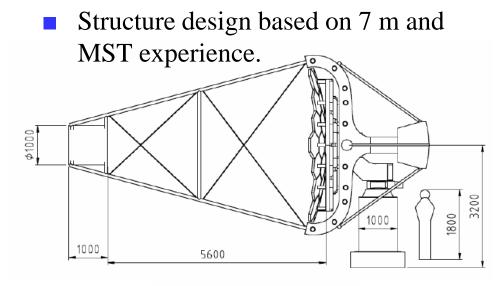


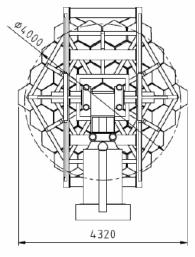
Optimum facet size 78 cm flat-to-flat.



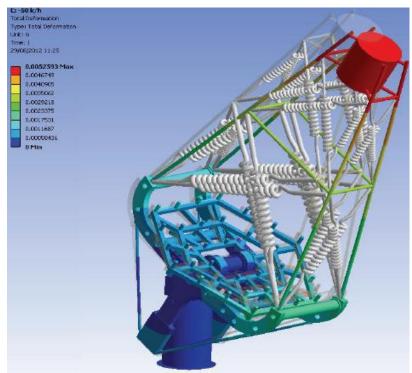
- Baseline: HESS mirror design.
- Galaktica (Armenia) can produce mirror for prototype at cost of €1700 for each of ~ 10 facets.
- Also looking at composite mirrors.

### 4 m DC SST – structure and mirrors





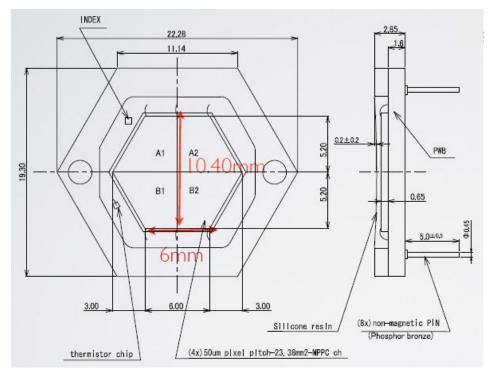
 Gravitational deformation plus effects of 50 km/h wind acceptable:



Lowest mode of oscillation ~ 6.5 Hz.

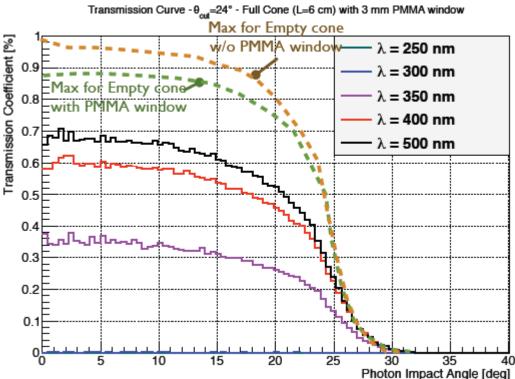
### 4 m DC SST – camera

SiPM development with Hamamatsu.

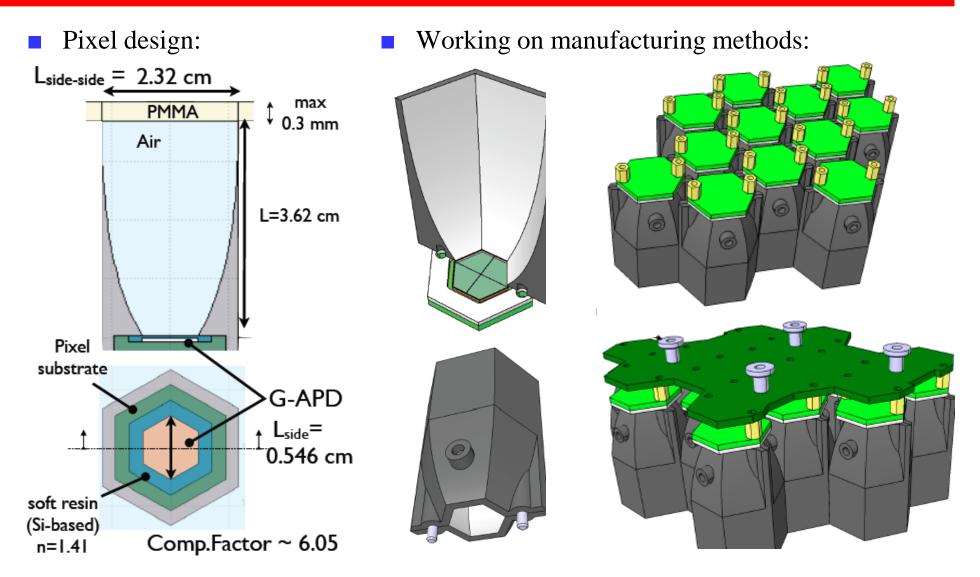


- Pixel is hexagonal, consists of 4 channels of area 23.4 mm<sup>2</sup>.
- Cells  $50 \times 50 \ \mu m^2$ , fill factor ~ 60%.

- Cost of (small numbers) of these SiPMs high, but price per mm<sup>2</sup> reasonable for large numbers.
- Use open Winston cone to limit absorption:

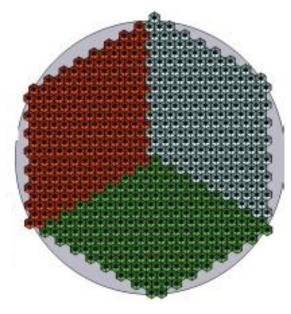


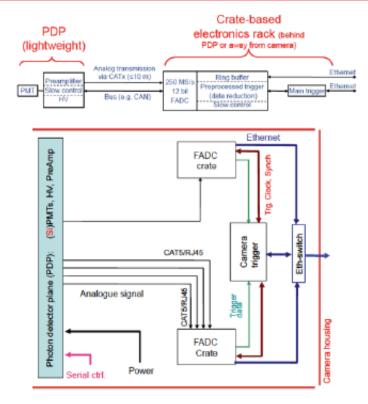
### 4 m DC SST – camera



## 4 m DC SST – camera

- DPNC proposal is to develop prototype of detector plane using 12 pixel units.
- Preamps and front end electronics will be developed with ETHZ.
- Will be matched to FlashCam...

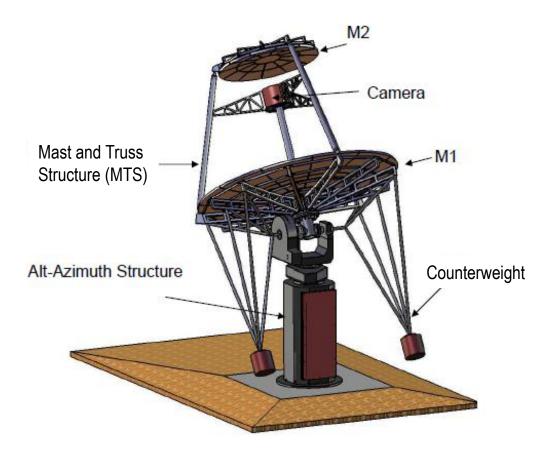




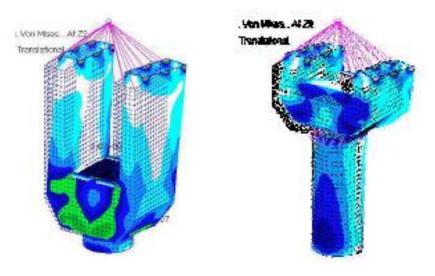
- ...but will also investigate e.g.
  EASIROC-based readout.
- Look at use of commercial cooling system.

#### Delphine Dumas

Structure design evolving:



Now use "short" fork as opposed to long design originally intended to accommodate camera electronics.



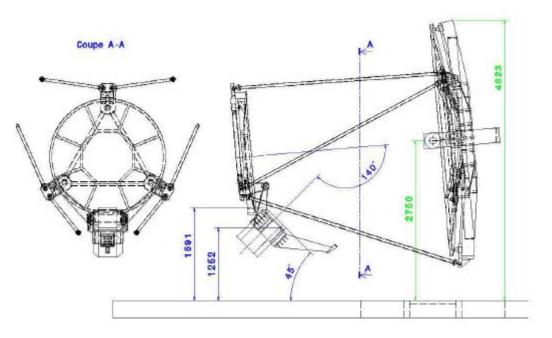
- Improves stiffness and reduces mass from 1.9 to 1.0 tons.
- Steel chosen for MTS.
- Support from technical division of INSU for drives.

Three arm Serrurier support for secondary mirror and camera under study:



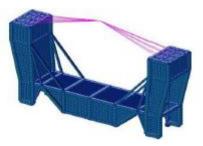
Significantly stiffer than previous design.

Camera removal using rotating support:

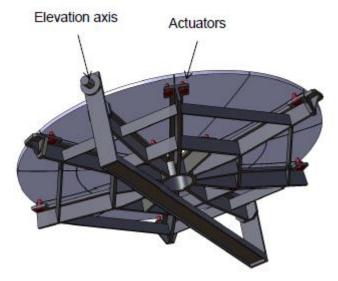


FEA indicates need to increase stiffness of dish.

Back up solution, wide fork...



...which carries dish at outer extremes:



Shelter will be used, provided by Losberger.



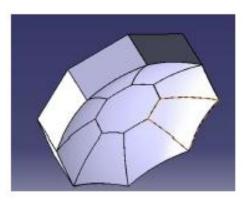
- Planning permission sought for erection of shelter in Meudon.
- Procurement planned for November.

- Secondary mirror will be constructed by CEA IRFU.
- Feasibility demonstrated.

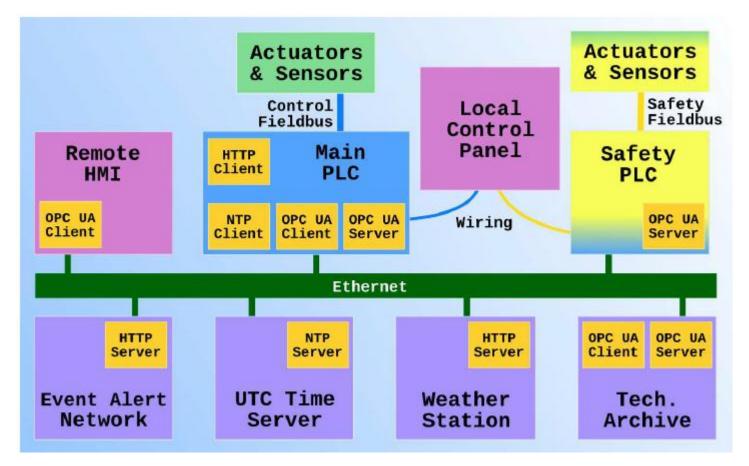


- Mould purchased.
- First panel expected December 2012.
- Complete mirror summer 2013.

- July 2012, talking to several possible suppliers of primary mirror.
  - Optic Technium (Wales).
  - Schott (Germany).
  - LZOS (Russia).
  - LAM (France).
- Developing collaboration with Nexeya for actuators.

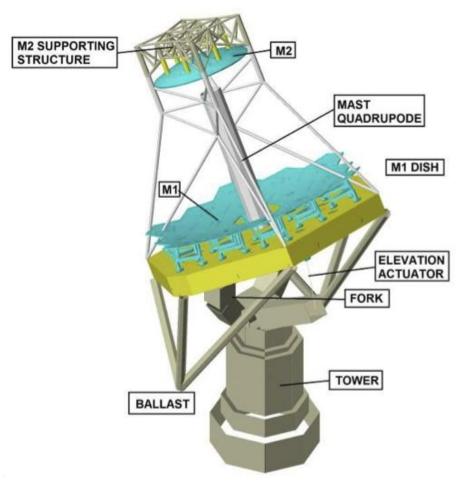


- Software and communication layout being designed.
- Main and safety Programmable Logic Controllers probably Siemens.

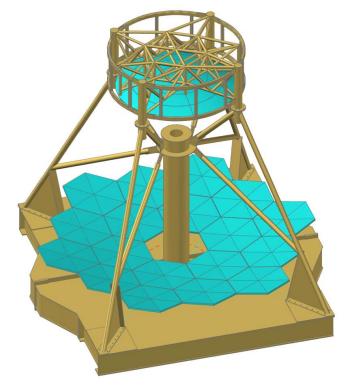


#### Rodolfo Canestrari, Giovanni Pareschi

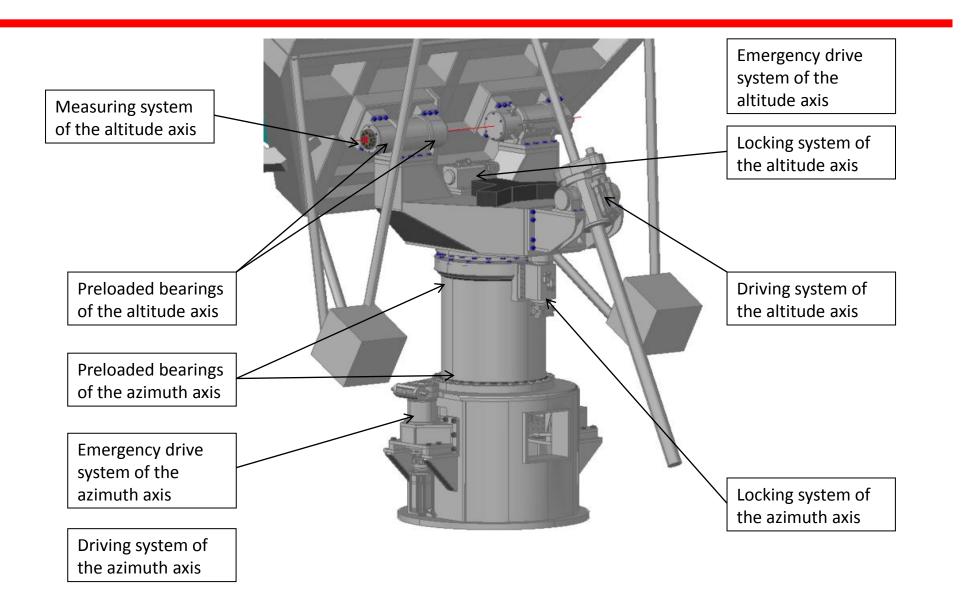
Structural design essentially complete:



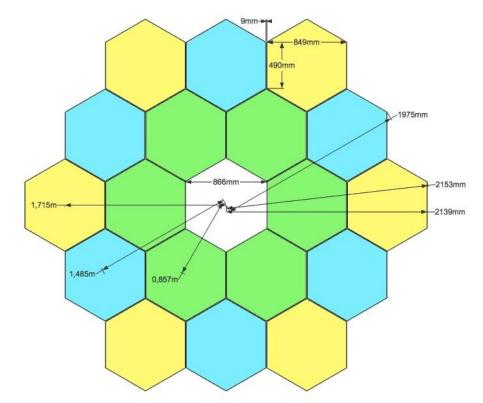
• Quadrupod/tube for secondary and camera support:



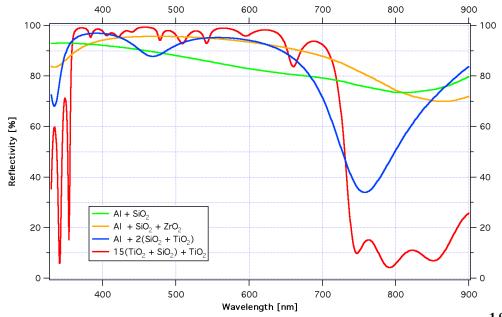
- Legs counteract lateral deformation.
- Tube increases torsional stiffness.



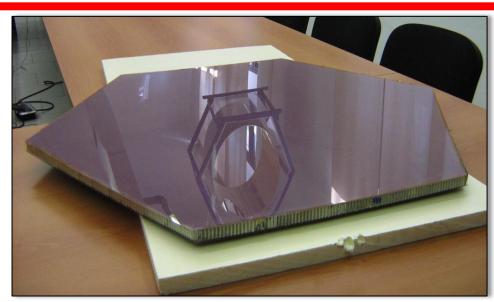
### Primary mirror:

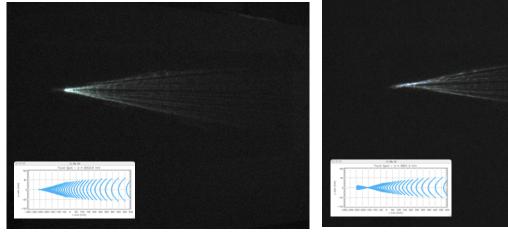


- Solution one, sandwich structure using thin (2 mm) hot slumped glass.
- Solution two, thick (10 mm) hot slumped glass.
- Dielectric coating procedure developed.



- First "yellow" hexagonal primary mirror panel delivered.
- Dimensions 850 mm flat-to-flat.
- Thickness 2 mm.
- Radius of curvature
  ~ 9.4 m.
- Aspherical profile.
- "Almost within specification"

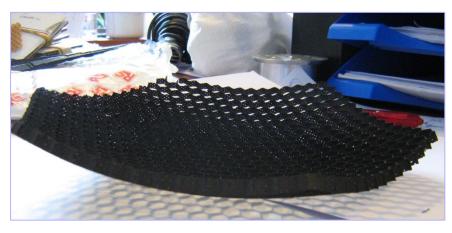




PSF, 0 mm off axis

PSF, 500 mm off axis

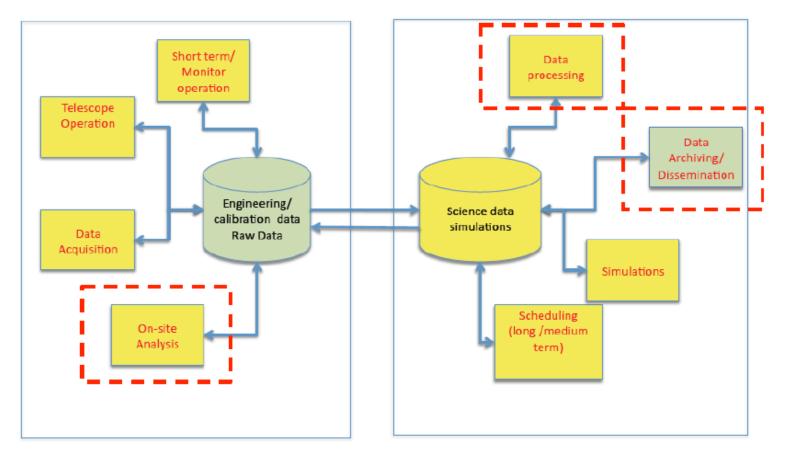
- Secondary mirror.
- Monolithic.
- Will use pre-shaped honeycomb:



- Optimization of design now underway with US manufacturer.
- Reduces stresses in structure which tend to cause astigmatism.
- Glass, developments with FLABEG.
- Secondary support structure will include three actuators and lateral constraints:
- Actuators will be connected via whiffletrees to spread load.

#### Massimo Trifoglio, Angelo Antonelli

 Control framework, use ALMA Common Software (ACS). Data handling and archiving under development.



## Camera development – ASTRI

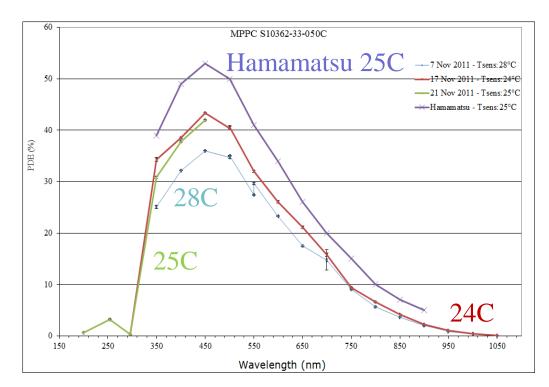
#### Giovanni La Rosa, Giovanni Bonanno

- Camera based on Hamamatsu SiPMs.
- Readout using EASIROC chip.
- Mechanical and electronic design progressing.



Camera mass about 25 kg.

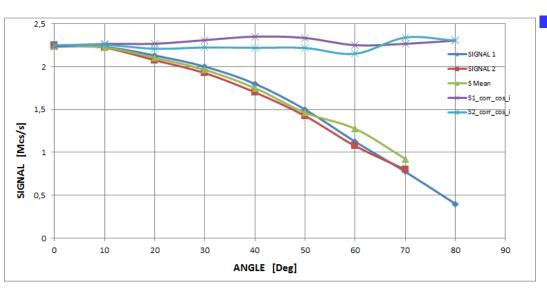
Characterizing SiPMs, e.g. PDE:



- Catania measurements corrected for cross talk.
- Peak PDE at ~  $450 \text{ nm of} \sim 40\%$ .

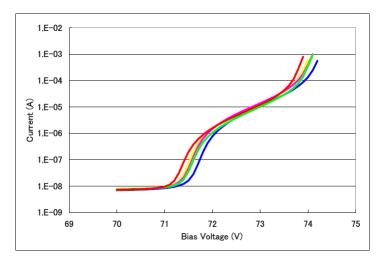
## Camera development – ASTRI

Signal as func. of incidence angle of light on SiPM:



 Signal decreases with cos θ, i.e. in proportion to light hitting sensor, to angles of incidence of (at least) 80°.

- ASTRI consortium has received 530 SiPM.
- I-V tests at Hamamatsu being repeated in Catania.



 Establishing quality control procedures for camera construction.

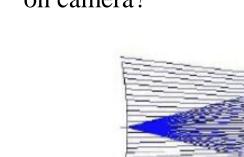
## Pointing

# Calibration

### Jacco Vink

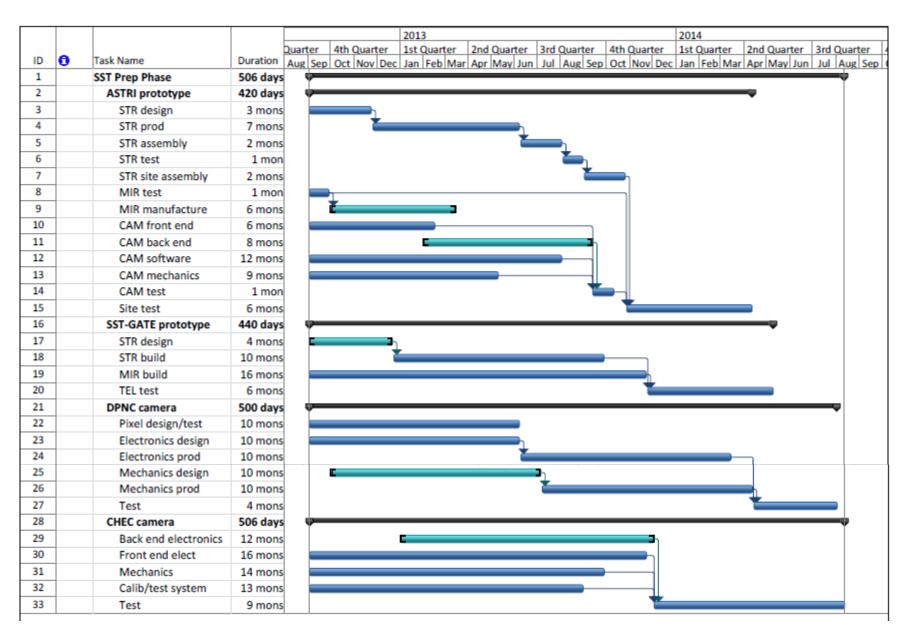
- Dutch contribution to CTA.
- Determine position of images to high accuracy.
- E.g. by imaging stars during observations.
- Use camera?
  - Saturation?
  - How many stars/field needed?
- Alter camera?
  - Replace one module/pixel by CCD camera?
  - Use light guides between modules?
- Funding sought.

- Use "photon statistics": variance of NSB images.
- Verified with UVscope at Auger.
- Use calibration system; need flat, uniform, fast pulse at right wavelength.
- Look e.g. at LEDs with suitable driver circuit.
- With monolithic secondary, can mount on camera?



Alberto Segreto, Michael Daniel

## Developing overall SST project planning



## Summary

- We have a 7 m DC SST structure "ready to go".
- Ideas on how to mass produce a SiPM-based camera for a Davies-Cotton telescope are being developing.
- Funding is being sort to pursue this option and build a prototype device.
- A first design of a telescope of 4 m diameter has been produced to match this camera.
- Dual Mirror prototypes suitable for MAPM-based and SiPM-based cameras are progressing well.

- The design and prototyping of the mirrors for these telescopes is going well.
- The design of cameras for the DM prototypes is progressing.
- Calibration and pointing concepts are being developed.
- Much will be known at the end of 2013 that will allow the writing of the SST section of the Technical Design Report for CTA.