

Laplace's equation in spherical polar coordinates

- In this lecture we will:

- ◆ See how Legendre polynomials arise in the solution of Laplace's equation in spherical polar coordinates.
- ◆ Introduce spherical harmonics.
- ◆ See how spherical harmonics are used in the quantum mechanical description of atoms.

- A comprehension question for this lecture:

- ◆ Prove that the function $G = r^{-l-1}$ is a solution of the equation

$$\frac{1}{G} \frac{d}{dr} \left(r^2 \frac{dG}{dr} \right) = l(l+1).$$

Laplace's equation in spherical polar coordinates

- In spherical polar coordinates, the gradient is: $\nabla V = \frac{\partial V}{\partial r} \hat{r} + \frac{1}{r} \frac{\partial V}{\partial \theta} \hat{\theta} + \frac{1}{r \sin \theta} \frac{\partial V}{\partial \phi} \hat{\phi}$.
- The divergence is: $\nabla \cdot \vec{A} = \frac{1}{r^2} \frac{\partial}{\partial r} (r^2 A_r) + \frac{1}{r \sin \theta} \frac{\partial}{\partial \theta} (\sin \theta A_\theta) + \frac{1}{r \sin \theta} \frac{\partial}{\partial \phi} (A_\phi)$.
- Putting them together, we get the Laplacian in spherical polar coordinates:
$$\nabla \cdot \nabla V = \nabla^2 V = \frac{1}{r^2} \left[\frac{\partial}{\partial r} \left(r^2 \frac{\partial V}{\partial r} \right) + \frac{1}{\sin \theta} \frac{\partial}{\partial \theta} \left(\sin \theta \frac{\partial V}{\partial \theta} \right) + \frac{1}{\sin^2 \theta} \frac{\partial^2 V}{\partial \phi^2} \right]$$
- Setting this expression equal to zero gives us Laplace's equation in spherical polar coordinates:
$$\nabla \cdot \nabla V = \nabla^2 V = 0.$$
- Lots of physical potentials are described by this equation and many of them depend on r and θ , but not on ϕ .
- Look for solutions to Laplace's equation that are independent of ϕ .
- Also assume we can solve by separating variables, i.e. that $V(r, \theta) = G(r)H(\theta)$.

Solving Laplace's equation by separating variables

- We can then rewrite the equation as:

$$\frac{1}{G} \frac{d}{dr} \left(r^2 \frac{dG}{dr} \right) = - \frac{1}{H \sin \theta} \frac{d}{d\theta} \left(\sin \theta \frac{dH}{d\theta} \right).$$

- The only way that a function of r and a function of θ can be equal for all values of r and θ is if they are both equal to the same constant.
- Write that constant as $l(l+1)$. (We will see later why this form is chosen!)
- We then have:

$$\frac{1}{G} \frac{d}{dr} \left(r^2 \frac{dG}{dr} \right) = l(l+1).$$

- Two solutions of this equation are:

$$G = r^l \quad \text{and} \quad G = r^{-l-1}.$$

- Prove that $G = r^l$ is a solution of

$$\frac{1}{G} \frac{d}{dr} \left(r^2 \frac{dG}{dr} \right) = l(l+1).$$

Solving Laplace's equation by separating variables

- Also: $-\frac{1}{H \sin \theta} \frac{d}{d\theta} \left(\sin \theta \frac{d}{d\theta} H \right) = l(l+1)$. Change variables by setting $w = \cos \theta$.
- This gives: $\frac{d}{d\theta} = \frac{dw}{d\theta} \frac{d}{dw} = -\sin \theta \frac{d}{dw}$ and $-\frac{1}{\sin \theta} \frac{d}{d\theta} = \frac{d}{dw}$.
- We then have: $-\frac{1}{H} \frac{d}{dw} \left(\sin^2 \theta \frac{dH}{dw} \right) = l(l+1)$ or $\frac{d}{dw} \left(\sin^2 \theta \frac{dH}{dw} \right) = -l(l+1)H$
- Rearranging: $\frac{d}{dw} \left(\sin^2 \theta \frac{dH}{dw} \right) + l(l+1)H = 0 \Rightarrow \frac{d}{dw} \left((1-w^2) \frac{dH}{dw} \right) + l(l+1)H = 0$.
- Differentiating w.r.t. w gives: $(1-w^2) \frac{d^2 H}{dw^2} - 2w \frac{dH}{dw} + l(l+1)H = 0$.
- This is Legendre's equation (with l instead of n)!
- The solutions of this equation are the Legendre polynomials $P_l(w) = P_l(\cos \theta)$.
- The solutions of the Laplace equation (without ϕ dependence) are therefore:
 $G(r)H(\theta) = r^l P_l(\cos \theta)$ and $G(r)H(\theta) = r^{-l-1} P_l(\cos \theta)$.

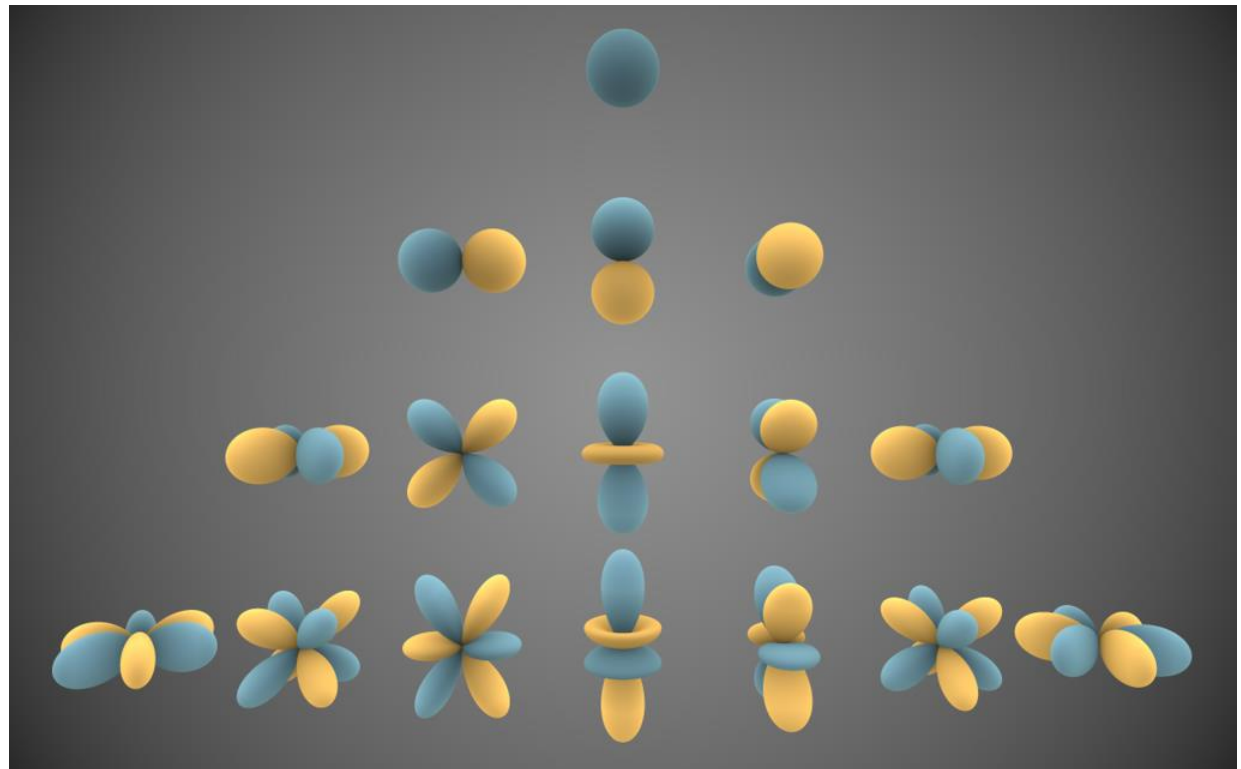
Spherical harmonics

- If we allow ϕ dependence, the Laplace equation can still be solved by separating variables; the angular part of the solution is given by the *spherical harmonics*:

$$Y_l^m(\theta, \phi) \propto \sin^m \theta \frac{d^m}{d(\cos \theta)^m} P_l(\cos \theta) \exp[im\phi], \text{ with } -l \leq m \leq l.$$

Wikimedia

- The picture shows the first few real spherical harmonics ($m = 0 \dots 3$).
- The distance from the origin shows the value of $Y_l^m(\theta, \phi)$ in the (θ, ϕ) direction, with blue being positive and yellow negative.



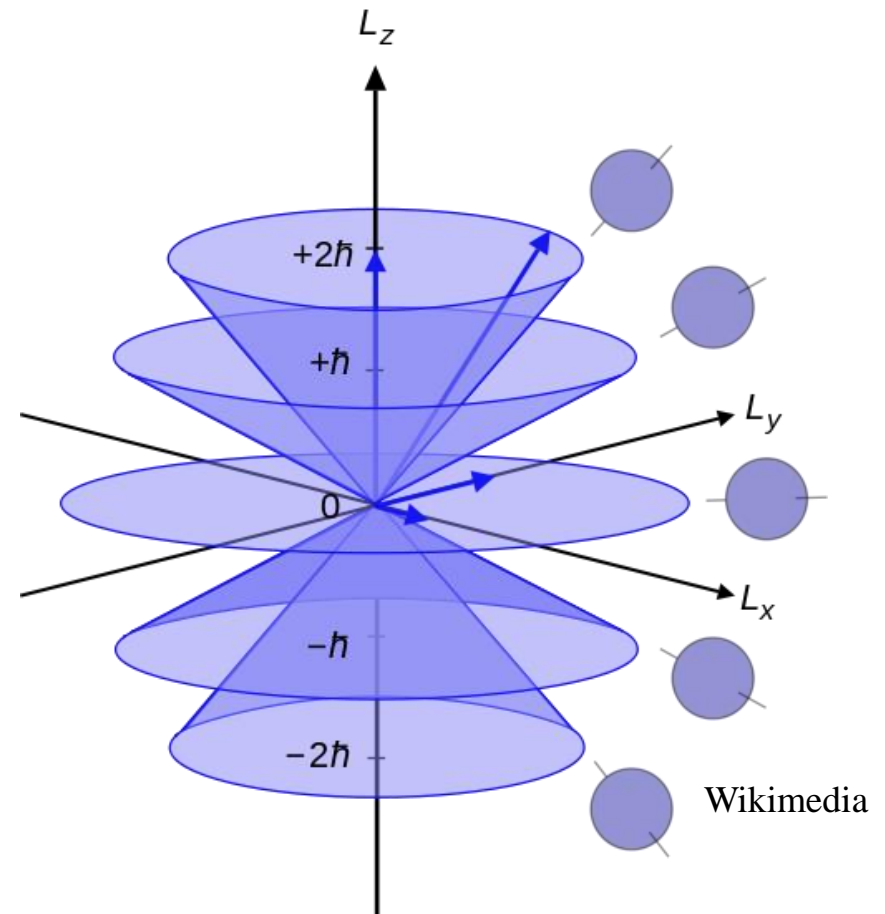
Schrödinger's equation for an H-like atom

- Schrödinger's equation describing an electron moving around a nucleus is:

$$-\frac{\hbar^2}{2m}\nabla^2\psi + V(r)\psi = E\psi.$$

- The solutions are of the form:
$$\psi_{nlm}(r, \theta, \phi) = R_{nl}(r)Y_l^m(\theta, \phi).$$
- The energy $E_n \propto 1/n^2$, i.e. it can only take on discrete values.
- The value of l is limited by $l \leq n - 1$.
- The magnitude of the orbital angular momentum of the electron is given by $L = \sqrt{l(l+1)}\hbar$.
- The z component of the orbital angular momentum is given by $L_z = m\hbar$.

- The *magnetic* quantum number m is restricted to the range $-l \leq m \leq l$.



Schrödinger's equation for an H-like atom

- The value of n , is called the principal quantum number.
- If an electron shifts from an orbit with $n = n_1$ to one with $n = n_2$, it emits (or absorbs) an energy:
$$E \propto \frac{1}{n_2} - \frac{1}{n_1}.$$
- As $E = hf = hc/\lambda$, this means energy is emitted from atoms at particular frequencies/wavelengths.
- As the nuclear charge of (and the number of electrons in) an atom influence the energy levels, this gives rise to distinctive spectra which allow atoms to be identified.
- Note that, in this solution, the energy is independent of l and m .
- The independence of the energy on the magnitude of the angular momentum vanishes when relativistic effects are considered.
- These effects introduce *fine structure* to the spectra.
- A further l dependence is also introduced if the atom is placed in a magnetic field, the *Zeeman effect*.
- This latter effect is used in nuclear magnetic resonance spectroscopy (NMR) and magnetic resonance imaging (MRI).