

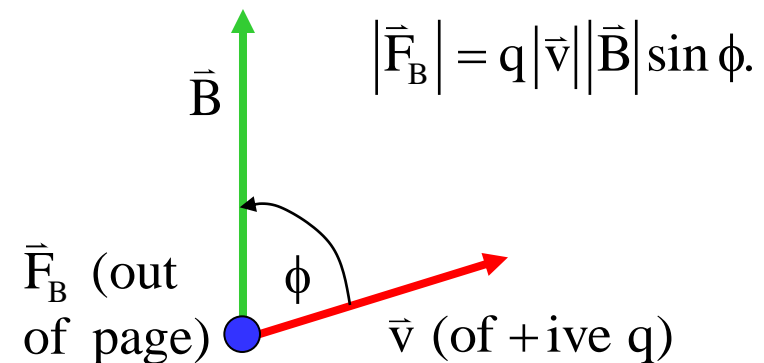
Lecture 12

- In this lecture we will look at:
 - ◆ Magnetic fields.
 - ◆ Force on a charged particle in a magnetic field.
 - ◆ Magnetic field lines.
 - ◆ Circulating charged particles.
 - ◆ Magnets in particle physics.
 - ◆ The Aurora Australis and the Aurora Borealis.
- After this lecture, you should be able to answer the following questions:
 - How does a magnetic field influence a) a stationary charged particle and b) a moving charged particle?
 - Use field lines to illustrate the field due to a bar magnet.
 - Explain how the mechanism that gives rise to the Aurora Borealis.

Magnetic Fields

- As soon as we start considering moving electric charges, we are faced with a new phenomenon: magnetism.
- Two comments:
 - ◆ Not really new phenomenon! (See “Relativity and Magnetism” in Lecture 14.)
 - ◆ Fields due to magnets – i.e. magnetic materials – are result of electric currents in the materials, e.g. electrons orbiting nuclei.
- The strength and direction of an electric field was defined in terms of the force it caused on a test charge:
$$\vec{F}_E = q\vec{E}.$$

- No magnetic monopoles, so cannot quite do same for magnetic field.
- Magnetic fields have no effect on stationary electric charges, but do cause force on moving charges.
- Magnitude and direction of magnetic force found to depend on velocity of charge (speed and direction!).

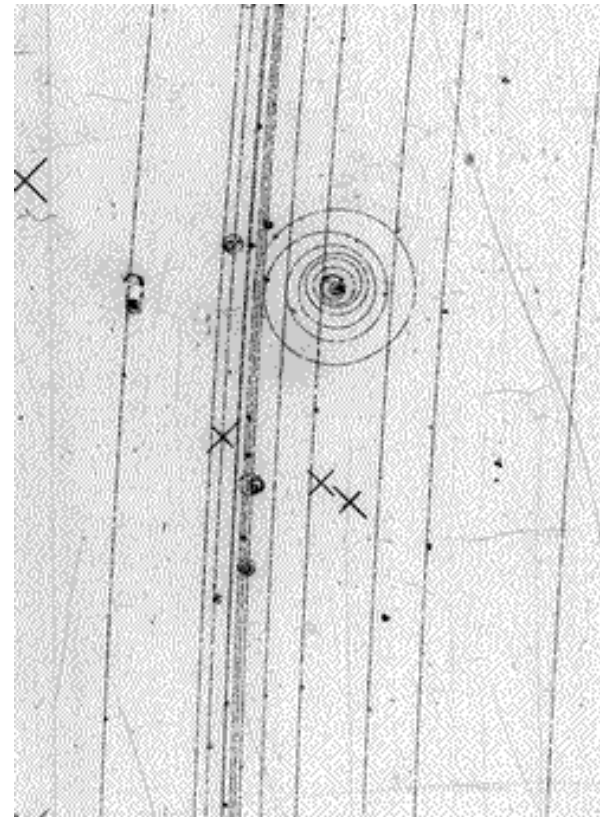


- See is cross product:
$$\vec{F}_B = q\vec{v} \times \vec{B} \quad [12.1]$$

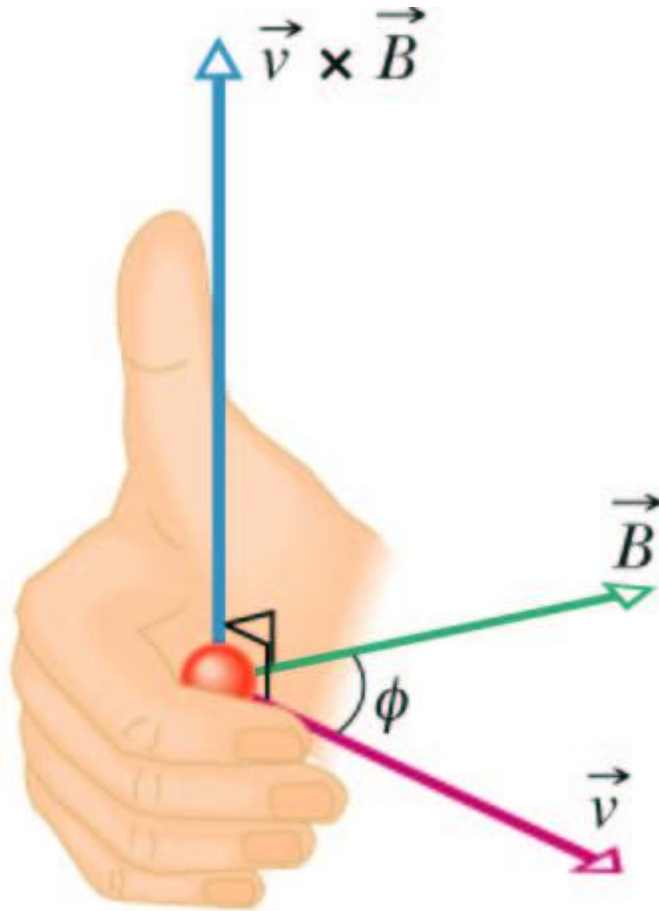
Force on a Charged Particle in a Magnetic Field

- This defines strength and direction of the magnetic (“B”) field.
- More descriptively, the direction of the B field is that in which the charged particle experiences no force.
- The strength of the B field is
$$|\vec{B}| = \frac{F_B}{|q|v}, \text{ when } \vec{v} \perp \vec{B}.$$
- Sense (which way along direction vector) from right hand rule:
 - ◆ 1st finger: velocity (+ive charge).
 - ◆ 2nd finger: B field.
 - ◆ Thumb: motion (i.e. force).

- Work out the direction of the B field in this bubble chamber photograph, assuming that the spiralling track is that of an electron:



Force on a Charged Particle in a Magnetic Field



- Another way of thinking of the RH rule:
 - ◆ Hold your right hand so your extended fingers point along the +ive particle's velocity.
 - ◆ Turn your hand so that when you curl your fingers they move towards the direction of the B field.
 - ◆ Your thumb is then pointing in the direction of the force.
- And another...
 - ◆ Rotate a RH screw from the direction of the velocity towards that of the B field.
 - ◆ The screw moves in the direction of the force.

Force on a Charged Particle: Typical Magnetic Fields

- Some comments:

- ◆ Force is proportional to charge, force on -ive particle opposite to that on +ive particle.
- ◆ Force always perpendicular to both particle's velocity and direction of the B field.
- ◆ The force doesn't change the speed (and hence kinetic energy) of the particle.

- The unit of magnetic field strength is the Tesla:

$$1\text{ T} = 1 \frac{\text{N}}{\text{C m s}^{-1}} = 1 \frac{\text{N}}{\text{C s}^{-1} \text{ m}} = 1 \frac{\text{N}}{\text{A m}}.$$

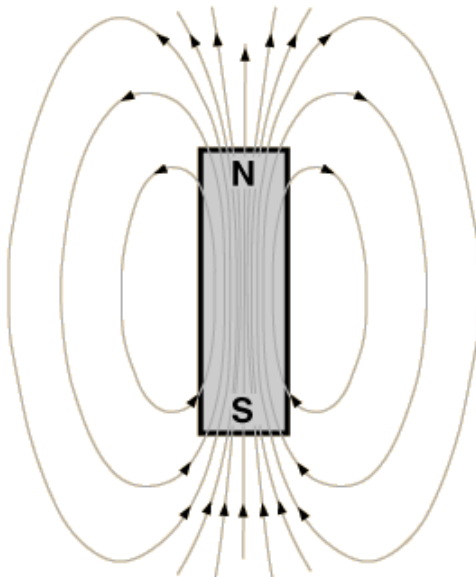
- Older unit, the gauss, $1\text{ T} = 10\text{ kG}$.

- Some typical magnetic field values:

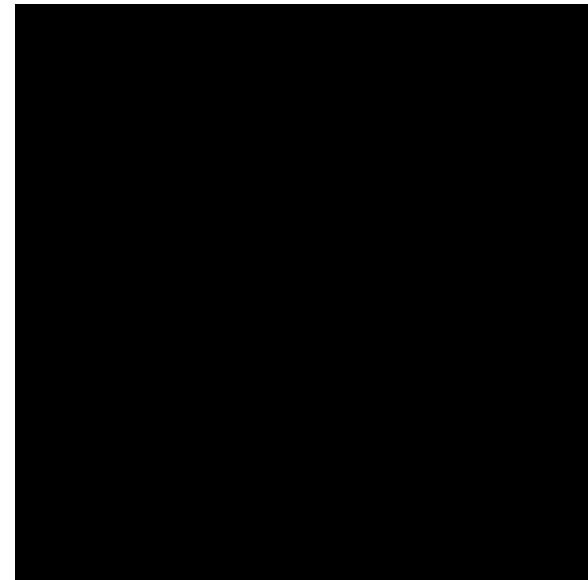
Surface of neutron star	10^8 T
Superconducting magnet	10 T
Small bar magnet	10^{-2} T
At earth's surface	$10^{-4}\text{ T} = 1\text{ G}$
Interstellar space	10^{-10} T
Magnetically shielded room	10^{-14} T

Magnetic Field Lines

- Represent magnetic field using field lines.
- Direction of the field line gives the direction of the B field.
- Density of the lines represents the magnitude of the field: the closer the lines the stronger the field.



- Video of simulation of one of periodic flips of earth's B field direction (flip takes ~ 1000 years!).
- Earth's magnetic field represented using field lines, blue = inward, yellow = outward lines.



Circulating Charged Particle

- If project beam of particles, into magnetic field with $\vec{v} \perp \vec{B}$, magnetic force causes particles to follow circular path, radius r .

- Magnetic force $F_B = qvB$.

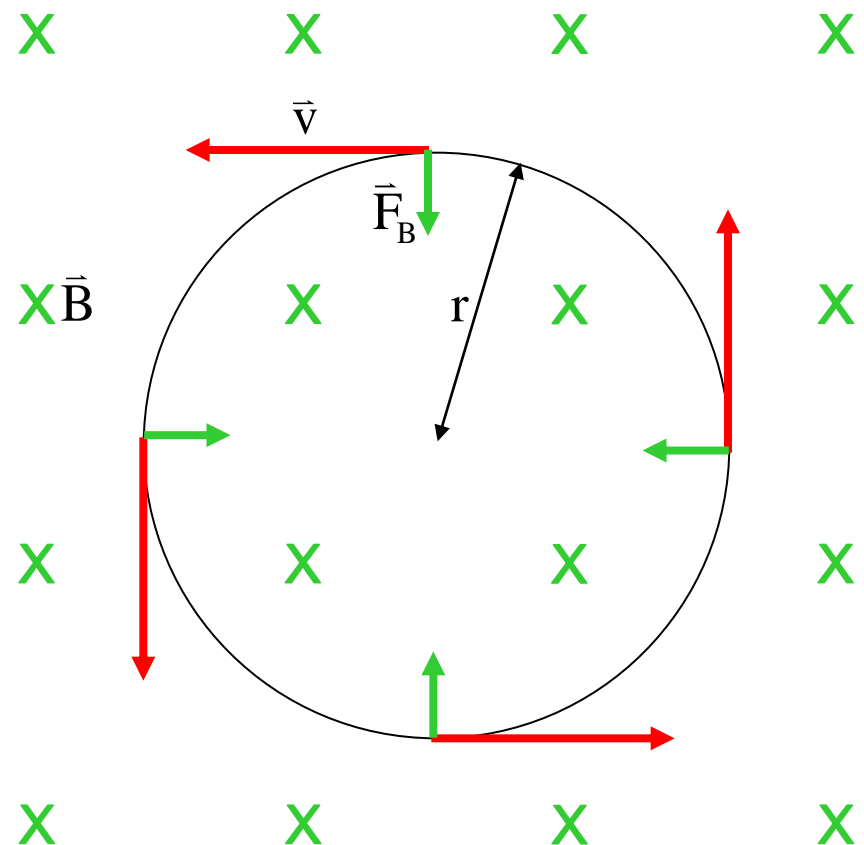
- Centripetal force:
$$F = mr\omega^2 = m \frac{v^2}{r}.$$

- Equating these:

$$qvB = m \frac{v^2}{r}$$

$$\Rightarrow r = \frac{mv}{qB} = \frac{p}{qB} \quad [12.2]$$

- B field directed into transparency.



Circulating Charged Particle

- The period (time for one revolution) is given by speed/distance, i.e.

$$T = \frac{2\pi r}{v} = \frac{2\pi m v}{v q B} = \frac{2\pi m}{q B}.$$

- Frequency given by:

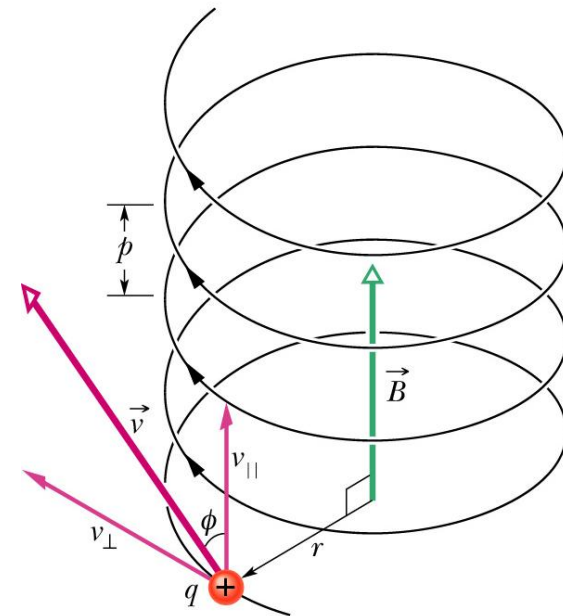
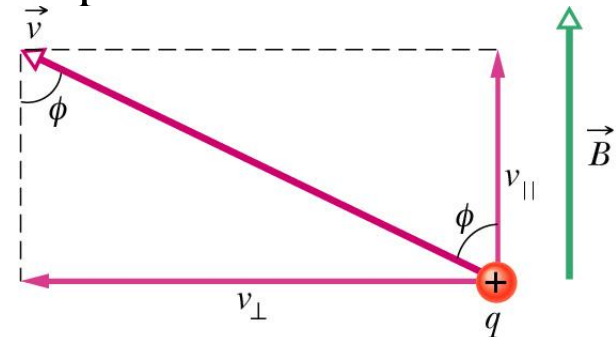
$$f = \frac{1}{T} = \frac{q B}{2\pi m}.$$

- Angular frequency:

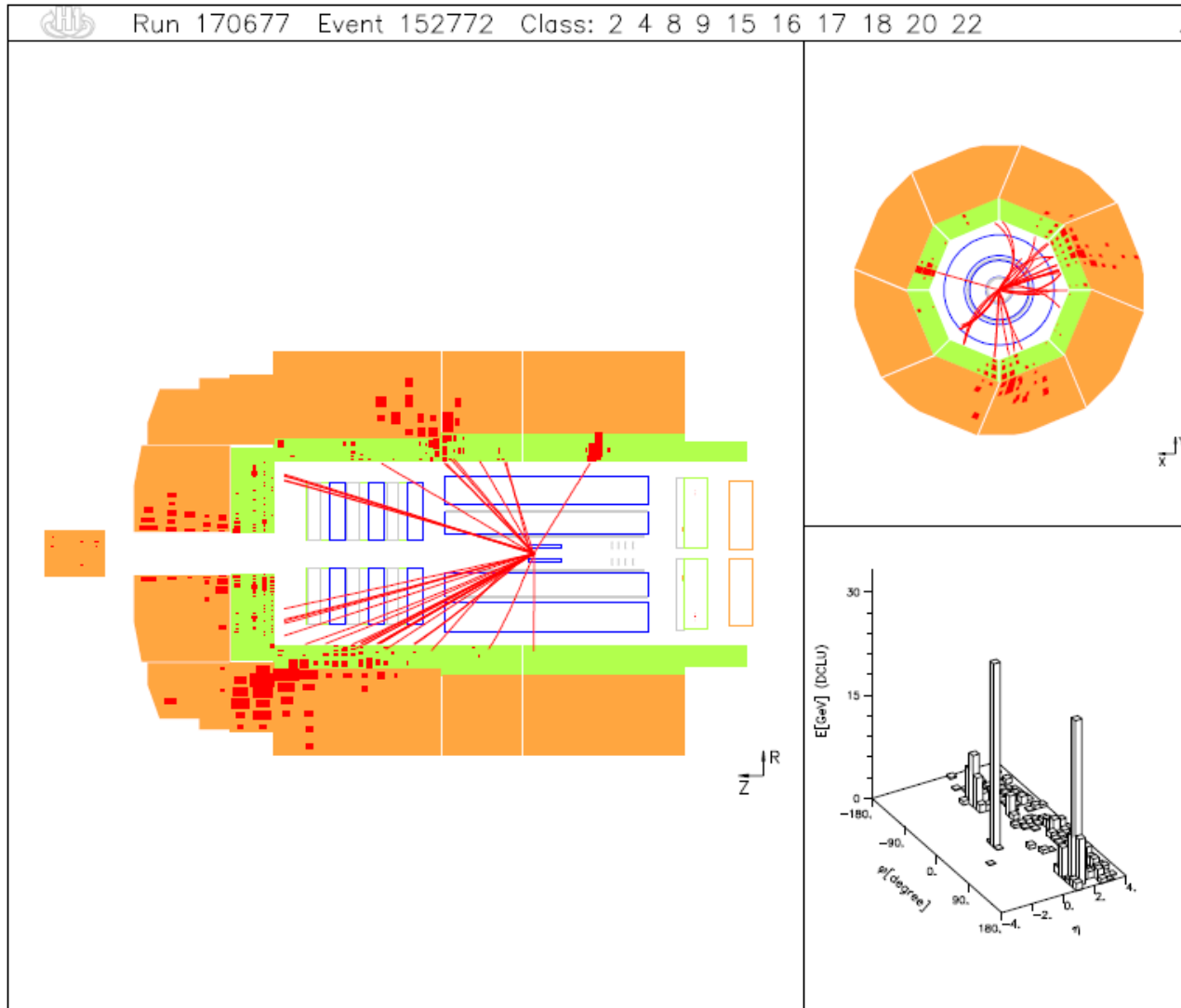
$$\omega = 2\pi f = \frac{q B}{m} \quad [12.3]$$

- If there is a component of particle's velocity, v_L , along direction of B field, particles follow helical path.
- Radius of helix given by v_T , transverse component of velocity, pitch by v_L .

- Helical path:

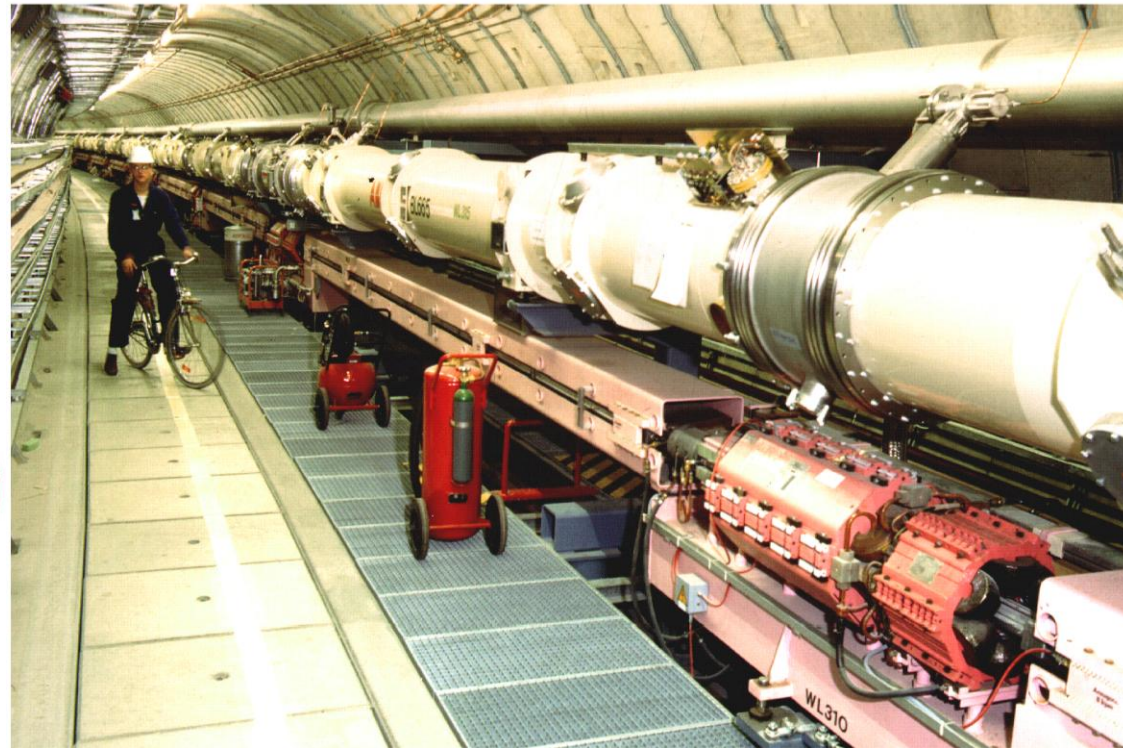
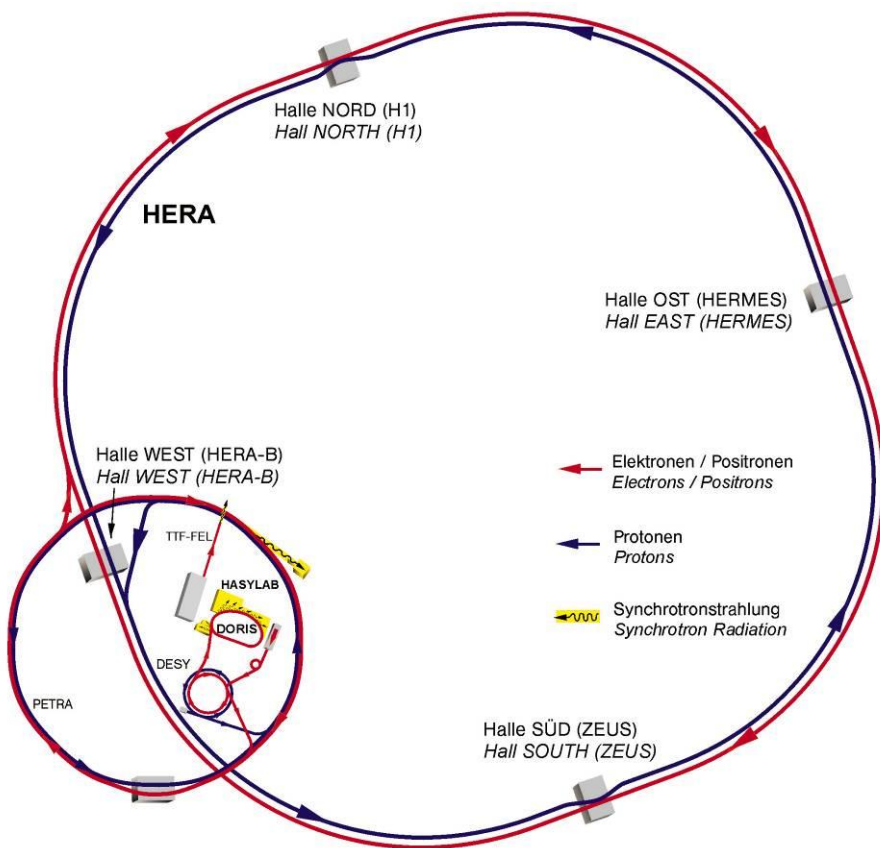


Magnets in Particle Physics Detectors



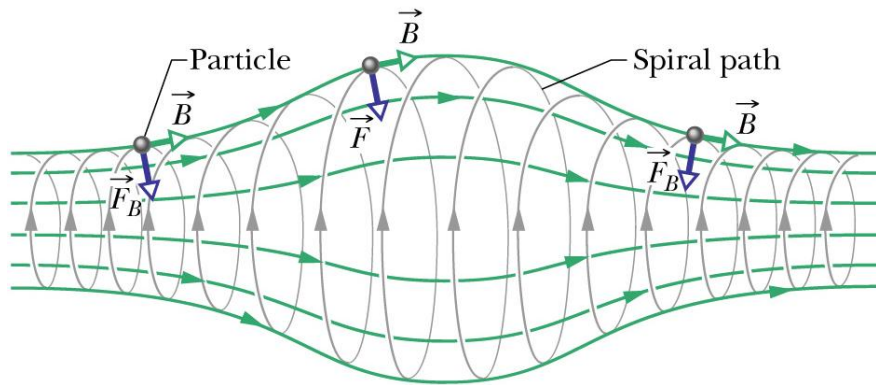
Magnets in Particle Accelerators

- Dipole magnets steer particles around accelerators and quadrupole magnets provide focussing.

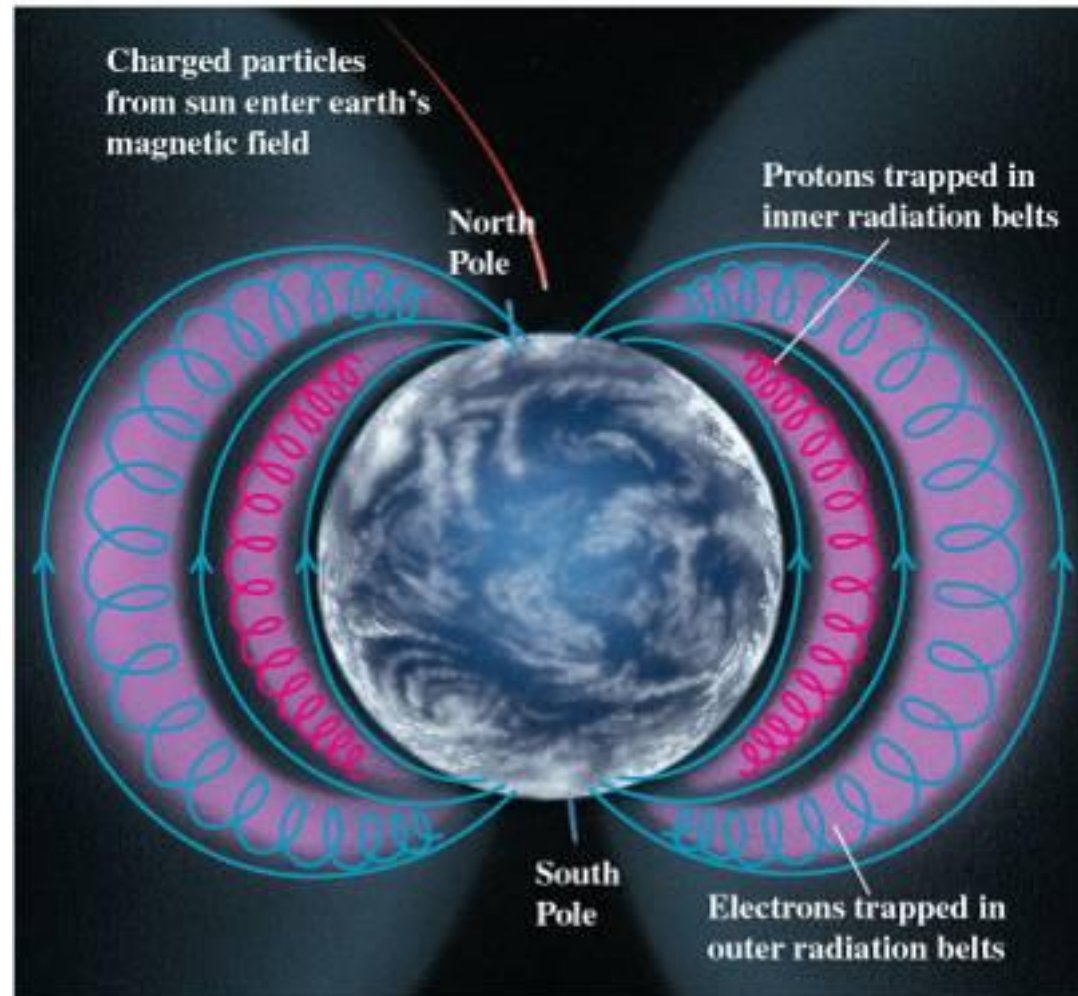


Magnetic Bottle

- Charged particle can become trapped in non-uniform B field:



- This happens in the magnetic field of the earth, gives rise to Van Allen Radiation Belts (right), Aurora Australis and Aurora Borealis (next slides) .



Aurora Australis, Photo from Discovery



Aurora Borealis



Photo: Pekka Parviainen



Photo: Pekka Parviainen

Aurora Borealis



Photo: Pekka Parviainen



Photo: Pekka Parviainen

Aurora Borealis



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