#### Optical Calibration and Scintillator Plant Construction on the SNO+ experiment

Christmas 2014

Rob Stainforth University of Liverpool

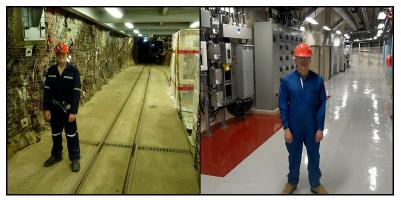
11th December, 2014

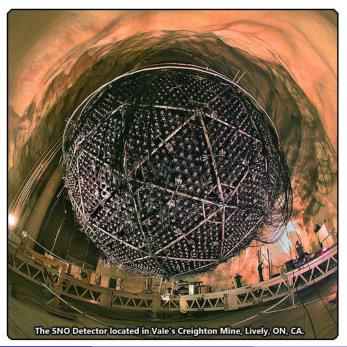


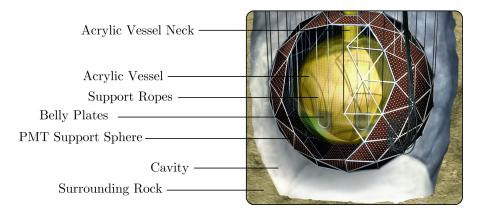
#### Future Work - LAST YEAR [2013]

- Continue work on LOCAS in time for laserball runs (early 2014) Optical Calibration
- Work onsite with other members of the collaboration in 2014 in Sudbury.

#### Construction







# Acrylic vessel to be filled with scintillator in 2015/6. Scintillator loaded (0.3 % g/g) with Tellurium [Telluric acid in solution] sometime during 2016.

# SNO+ Cavity Access

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# SNO+ Cavity Access

# SNO Cavity Access

#### **Optical Calibration**

#### • An optical calibration of the SNO+ detector:

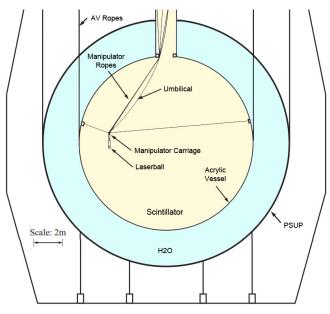
- In-situ measurements from a controlled source to calibrate detector
- Characterise the optical response of the detector. i.e. How the scintillator, acrylic of the AV, water and PMT response affect scintillation light and its detection.

# Controlled: [position, intensity, wavelength] light source, ideally isotropic . . .

Calibration data collected using an in-situ light source; the Laserball.

# Perform many runs with the laserball in different positions and at different intensities and wavelengths.

#### The Laserball



# Inside the Acrylic Vessel

#### **Optical Model**

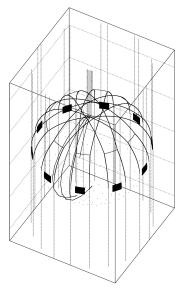
Occupancy at PMT *j* from laserball run *i* relative to central laserball run, modelled as follows:

$$\mathsf{R}_{\text{Occ},ij} = \frac{\mathsf{Occ}_{ij}}{\mathsf{Occ}_{0j}} = \frac{\Omega_{ij}\mathsf{R}_{ij}\mathsf{T}_{ij}\mathsf{L}_{ij}e^{-(d_{ij,scint}/\alpha_{scint}+d_{ij,av}/\alpha_{av}+d_{ij,H2O}/\alpha_{H2O})}}{\Omega_{0j}\mathsf{R}_{0j}\mathsf{T}_{0j}\mathsf{L}_{0j}e^{-(d_{0j,scint}/\alpha_{scint}+d_{0j,av}/\alpha_{av}+d_{0j,H2O}/\alpha_{H2O})}}$$

- $\Omega_{ij}$ : Solid angle subtended by the PMT
- $R_{ij}(\theta_{PMT})$ : PMT angular response beyond the solid angle
- $T_{ij}$ : Transmission coefficient for material boundaries
- $L_{ij}(\theta_{LB}, \phi_{LB})$ : Isotropy distribution
- $\bullet~d_{scint}, d_{av}, d_{h2o}\text{:}$  Light distance in scintillator, acrylic vessel and water
- $\alpha_{scint}, \alpha_{av}, \alpha_{h2o}$ : Attenuation lengths

#### black: calculated values red: fitted values

#### Revised PMT Shadowing Calculation



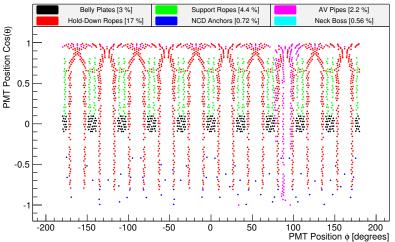
LightPathCalculator extrapolated geometry positions.

Rob Stainforth (Liverpool)

#### Revised PMT Shadowing Calculation

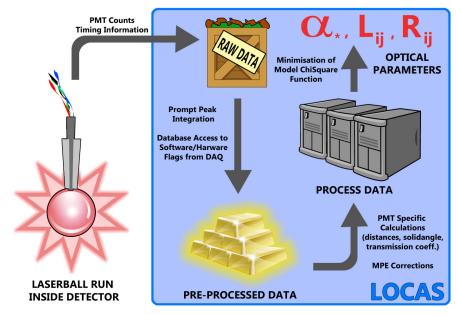
#### Shadowed PMTs from Laserball Position (0.0, 0.0, 0.0) cm

Total PMT Shadowing: 28 %



Example output of **new** PMT shadowing calculation technique.

#### LOCAS: Lisbon, Liverpool, Optical Calibration Sussex



Rob Stainforth (Liverpool)

SNO+

#### LOCAS: Optics Fit with original SNO data

- October 2003, 420 nm laserball waterfill data set.
- 27 off-axis (i.e. not central) runs. All using the same normalising central run.
- Fit using aforementioned occupancy ratio method as input data points.

#### LOCAS: Optics Fit

Minimise over the chi-square function:

$$\chi^2 = \sum_{i=1}^{N_{\rm runs}} \sum_{j=1}^{N_{\rm pmts}} \frac{\left(R_{\rm Occ,\ ij}^{\rm data} - R_{\rm Occ,\ ij}^{\rm model}\right)^2}{\sigma^2 + \sigma_{\rm PMT}^2},$$

using Levenberg-Marquardt algorithm.

- $R_{Occ, ij}^{data}$ ,  $R_{Occ, ij}^{model}$  Observed and respective model predicted occupancy ratios.
- $\sigma$  Statistical error on measured occupancies from off-axis and central run.
- $\sigma_{\rm PMT}^2$  PMT variability correction based on  $\theta_{\rm PMT}$ .

#### LOCAS: Optics Fit - Minimisation and Results

Perform 4 fits + 1 nominal fit using following chisquare limits for each data point at the top-level of each fit:

 $\chi^2_{\max}: 1000, 100, 25, 16+16$ 

LOCAS (SNO)

- $\chi^2_{\rm reduced} = 108259.1016/(144364 510) = 0.7526$
- $\alpha_{\text{D}_2\text{O}} = (6.72 \pm 0.01) \times 10^{-5} \text{ cm}^{-1}$
- $\alpha_{\rm H_2O} = (1.50 \pm 0.29) \times 10^{-4} \ {\rm cm^{-1}}$

#### LOCAS (SNO+)

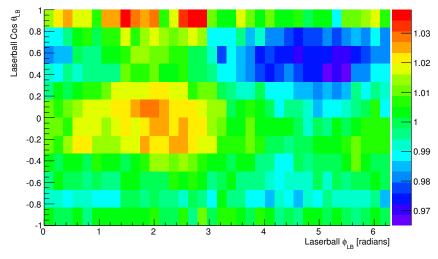
• 
$$\chi^2_{\rm reduced} = 165855.1719/(171059 - 510) = 0.9725$$

• 
$$\alpha_{\text{D}_2\text{O}} = (6.20 \pm 0.01) \times 10^{-5} \text{ cm}^{-1}$$

•  $\alpha_{\rm H_2O} = (2.20 \pm 0.25) \times 10^{-4} \ {\rm cm^{-1}}$ 

#### LOCAS: Optics Fit - Results

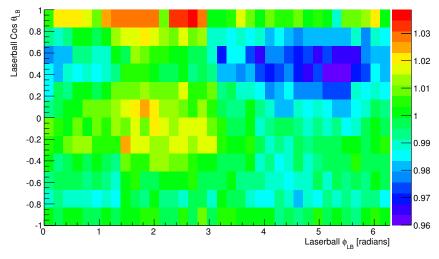
#### Laserball Distribution Histogram [Binned]



LOCAS (SNO) laserball distribution histogram

#### LOCAS: Optics Fit - Results

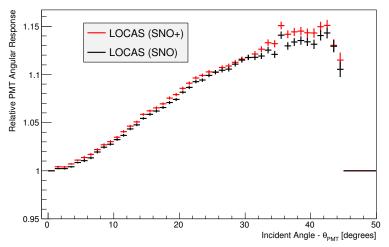
#### Laserball Distribution Histogram [Binned]



LOCAS (SNO+) laserball distribution histogram

#### LOCAS: Optics Fit - Results

**PMT Angular Response** 



Comparison of the fitted PMT angular response (histogram) distributions from LOCAS (SNO) and LOCAS (SNO+) using the same shadowing values.

#### Long Term Stay in Sudbury

- 8 months on site working with various members of the collaboration.
- Gained valuable experience and skills from a variety of tasks.

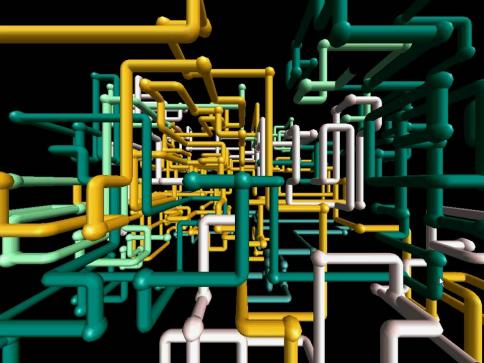


Working with John Walker

# SNO+ Scintillator Plant - Upper Level



# SNO+ Scintillator Plant - Lower Level



#### Helium Leak Checking and Construction

- Plant consists of 13 Vessels, 17 Kettles, 9 Columns and  $\sim$  2 km of piping.
- Purpose is to test seal/weld integrity of all volumes in the plant.
- SNO+ requires < 10<sup>-9</sup> mBar.l.s<sup>-1</sup> leak rate to beat naturally occuring Radon background levels in the mine.



#### Rough Pump

# Helium Leak Detector

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#### **Helium Source**

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TEL UNIDIONER OFFICE

HELIUM FOR BALLOONS

Highland

# Inject Helium in Flange/Instrument

### Holium leak checking crew - November

### Helium leak checking crew

Vovember

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#### **Future Work**

- Continue work on LOCAS in time for laserball data runs (mid-late 2015).
- Write thesis.



Moose

END.