

The main circuits of the LHC (RB, RQD, RQF) have about 24000 splices.

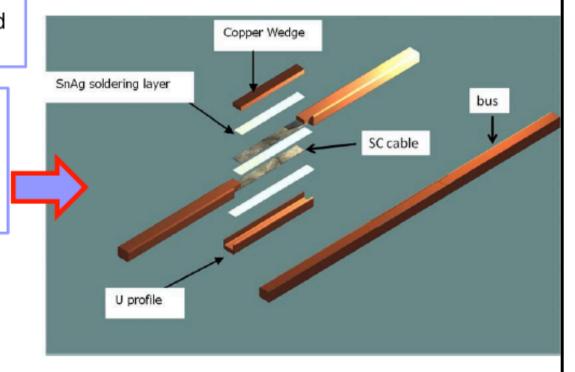
Out of these there are:

- •10170 interconnect splices and
- •13796 magnet splices

Interconnect splices are not protected by diodes and in the case of a problem all the current of the circuit passes through them

Nominal interconnect splice resistance:

- •At cold:  $300p\Omega$
- •At warm (300K): **10**μ**Ω**



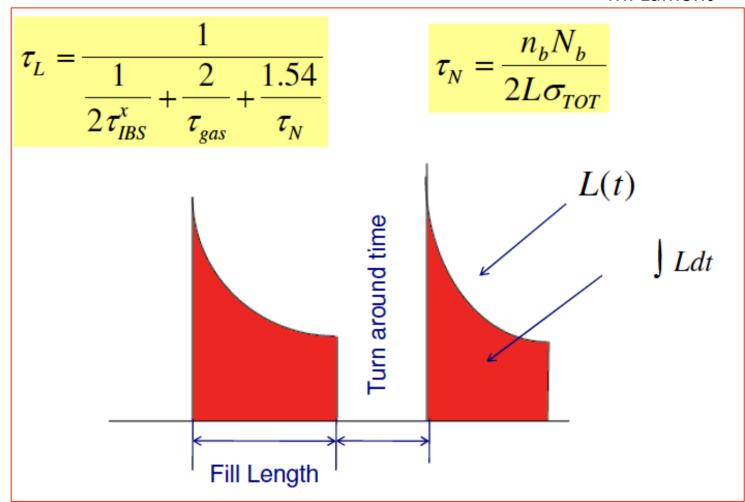
For the LHC to operate safely at a certain energy, there is a limit to how big a splice resistance can be

- Splices at cold (in the superconducting state) have been measured with excellent accuracy and do not pose a problem.
- Splices at warm (copper stabilizer) have been measured in part of the machine and extrapolated to the whole machine using statistical methods.
  - worse splice measured: 60±1μΩ
  - worse splice known to exist in the machine: 53±15μΩ
  - worse splice extrapolated: 90μΩ
- The current knowledge of the interconnect splices leaves no margin even for operation at 3.5TeV.
- 5TeV running is excluded without major repairs after a warm up.
- Two methods have been proposed to increase our knowledge of the interconnect splices
  - A low current method that can measure the RRR of the busbars
  - A high current method (the Thermal Amplifier) that is sensitive to the worst splices in all bus bar segments
- Using any of the above methods would allow us to either run at a higher energy around 4TeV and/or get a bigger margin at 3.5TeV.

# Two Possible Scenarios 2010-2011

- Run at 3.5 TeV/beam up to a predefined integrated luminosity with a date limit. Then consolidate the whole machine for 7TeV/beam.
  - · Need to determine the needs for the shutdown (resources, coactivity etc)
- 2. Run until second half 2010 then do minimum repair on splices to allow 5TeV/beam in 2011 (7TeV/beam comes much later)
  - ? Do DN200s at same time
  - ? Will we need to warm all sectors in order to re-measure (looks like yes to 7 RB octants from Mike's results, and 8 RQ)
  - ? How many splices will we need to repair to reach the "limit" copper stabilizer resistances (what about the RQs?)

Unanimous at Chamonix for 1)  $\rightarrow$  can expect CERN decision, watch for 'date'



Turnaround time about 3h; 30 days, 3 days stop, 2 days MD, 60% up (!) after 1year

#### 2010

Step	E [TeV]	Fill scheme	N	β* [m] IP1 / 2 / 5 / 8	Run time (indicative)	
1	0.45	2x2	5x10 <sup>10</sup>	11 / 10 / 11 / 10		
2	3.5	2x2	2 - 5x10 <sup>10</sup>	11 / 10 / 11 / 10	Weeks	
3	3.5	2x2*	2 - 5x10 <sup>10</sup>	2/10/2/2		
4	3.5	43x43	5x10 <sup>10</sup>	2/10/2/2	Weeks/Months	
5	3.5	156x156	5x10 <sup>10</sup>	2/10/2/2	weeks/ Piontins	
6	3.5	156x156	9x10 <sup>10</sup>	2/10/2/2		
7	3.5	50 ns - 144**	7x10 <sup>10</sup>	2.5 / 3 / 2.5 / 3	Months	
8	3.5	50 ns - 288	7x10 <sup>10</sup>	2.5 / 3 / 2.5 / 3		
9	3.5	50 ns - 720	7x10 <sup>10</sup>	2.5 / 3 / 2.5 / 3	Months	

20pb<sup>-1</sup> per month

One month: 720 bunches of 7 e10 at beta\* = 2.5 m gives a peak luminosity of 1.2 e32 cm<sup>-2</sup>s<sup>-1</sup> and an integrated of about 105 pb<sup>-1</sup> per month

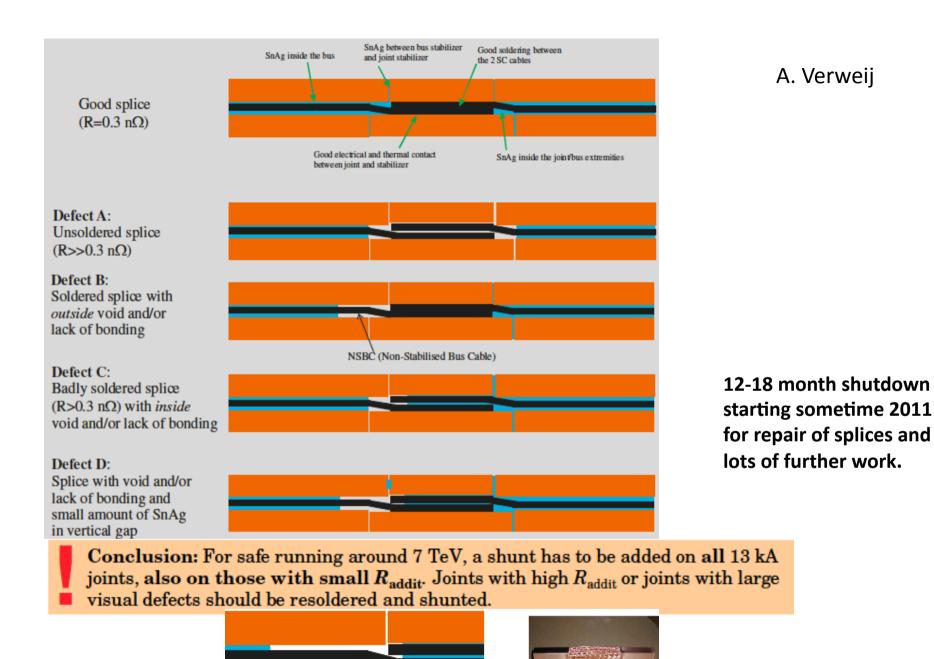
[15% nominal – 28 MJ]

2011: "flat out at 100 pb<sup>-1</sup> per month"

My guess: by June have ~30pb<sup>-1</sup> for ICHEP

<sup>\*</sup> Turn on crossing angle at IP1.

<sup>\*\*</sup>Turn on crossing angle at all IPs.



# By 2014 may have 10-30 fb<sup>-1</sup>

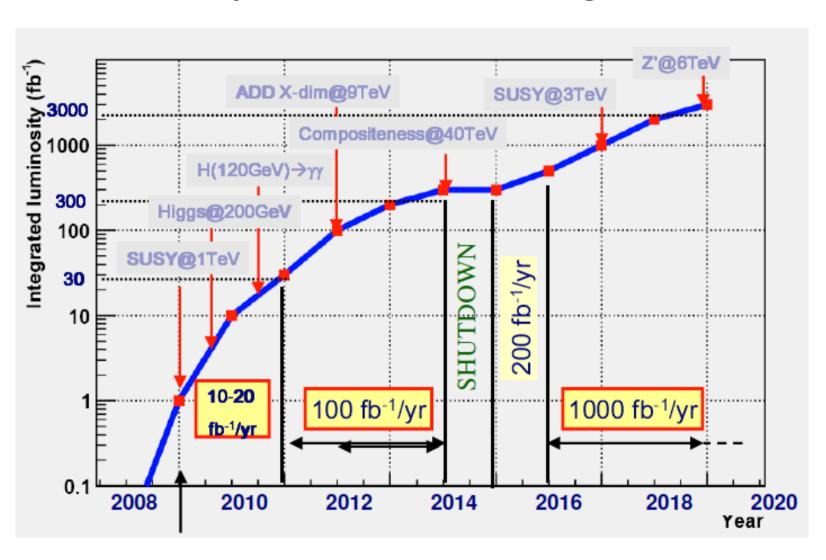
M.Lamont

Year	Months	energy	beta	ib	nb	Peak Lumi	Lumi per month	Int Lumi Year	Int Lumi Cul
2010	8	3.5	2.5	7 e10	720	1.2 e32	-	0.2	0.2
2011	8	3.5	2.5	7 e10	720	1.2 e32	0.1	0.8	1.0
2012									
2013	6	6.5	1	1.1 e11	720	1.4 e33	1.1	7	8
2014	7	7	1	1.1 e11	1404	3.0 e33	2.3	16	24

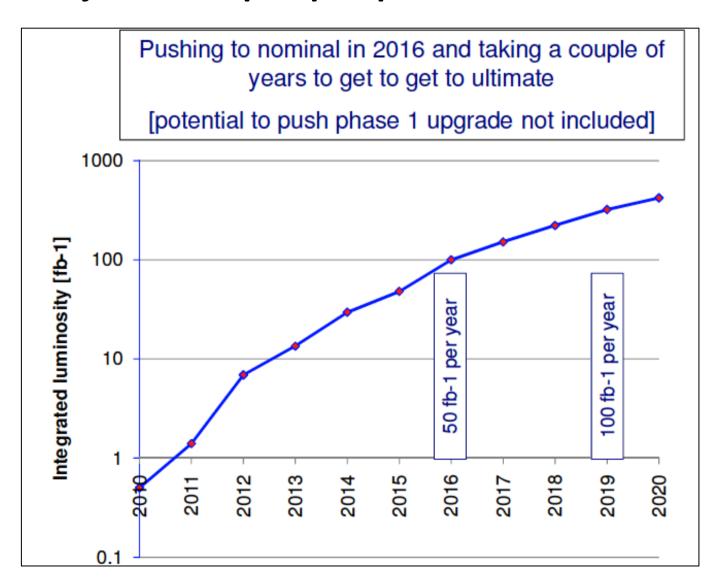
#### M. Ferro Luzzi

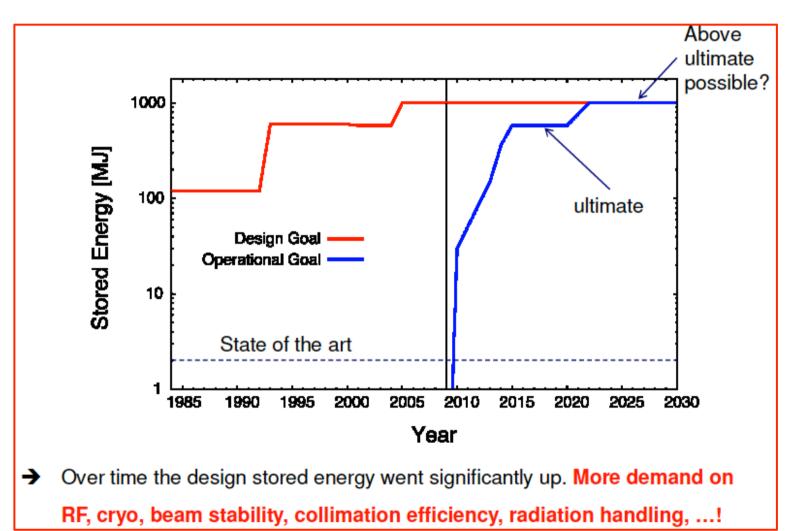
Year	Months	energy	beta	ib	nb	Peak Lumi	Lumi per month	Int Lumi Year	Int Lumi Cul
2010	6	3.5	2.5	7 e10	720	1.0 e32	-	0.1	0.1
2011	9	3.5	2.5	9 e10	720	2.0 e32	0.1	1	1.1
2012									
2013	6	6.5	1	9 e10	720	9 e32	0.45	2.7	3.8
2014	9	6.5	1	9 e10	1404	1.7 e33	0.6	5.3	9.1

# **Projections – some time ago**

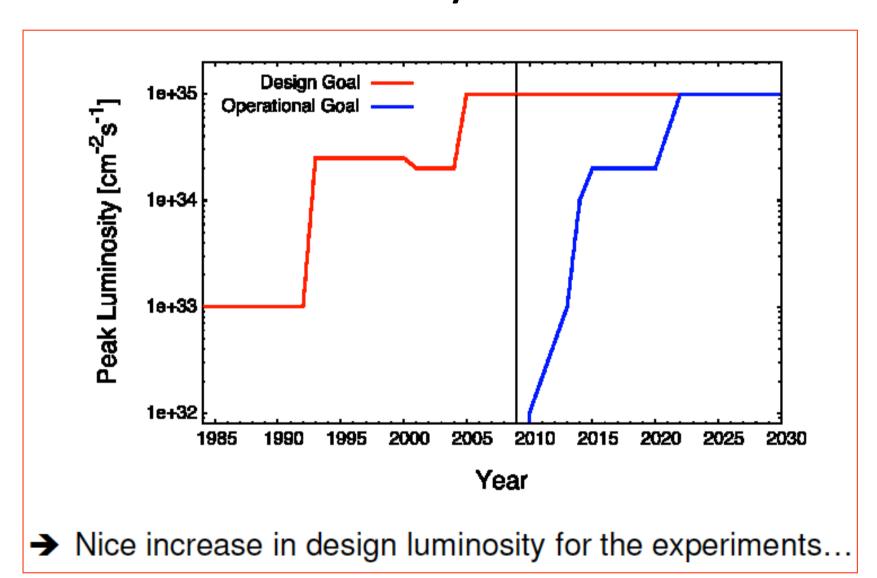


# Projections: open perspective until 2030

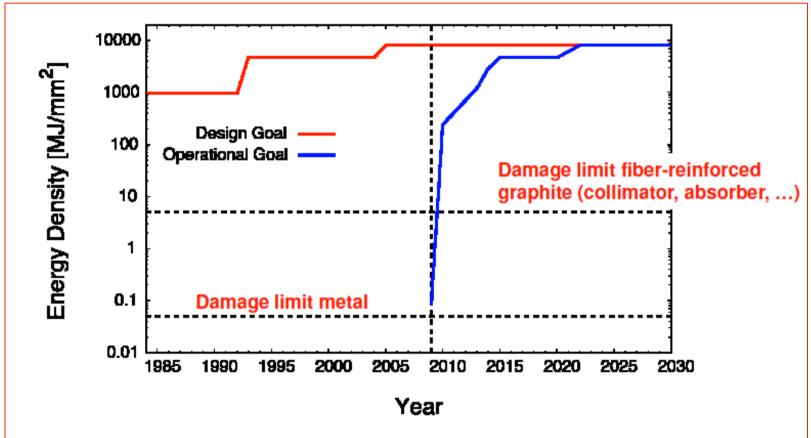




# **History**



### **History**



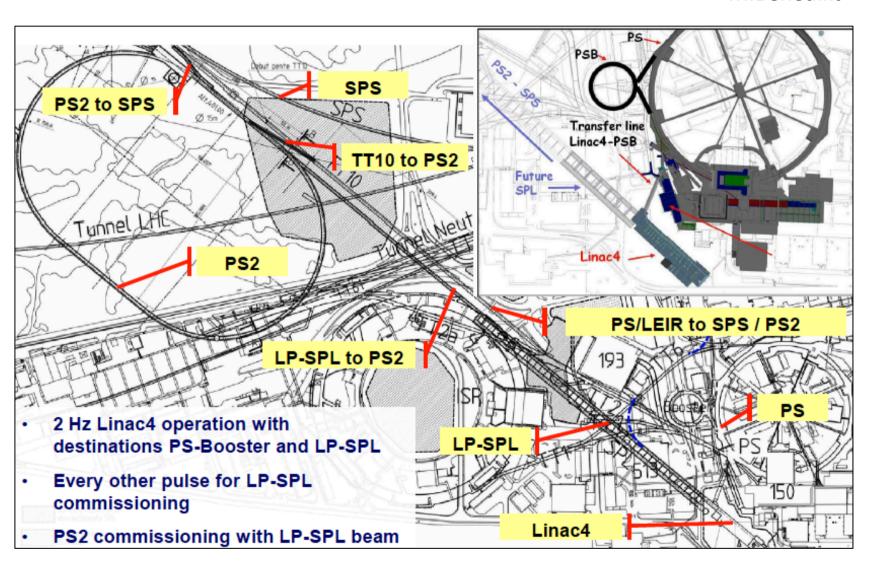
→ Transverse energy density is pushed further, way above damage limits of materials! At some point classical protection is not feasible. Must look at advanced technologies (e.g. SLAC rotatable collimator).

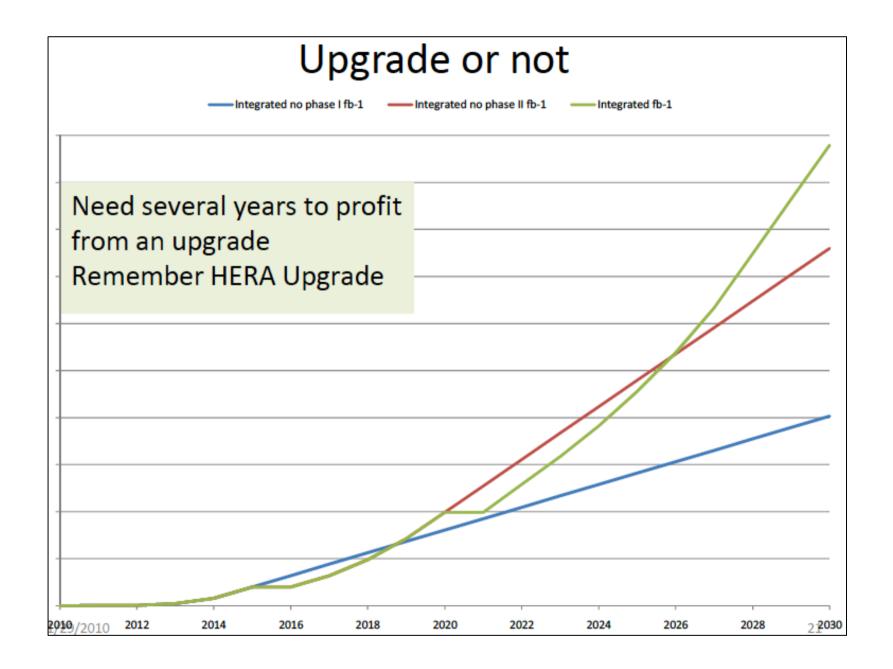
- Ultimate intensity is challenging for the LHC. Many systems at technological limits with little or no margin.
- Long (incomplete) list of required LHC work collected:
  - □ "New" RF system, possibly requiring civil engineering.
  - □ New DSL in IR3, review of potted magnets, radiation damage.
  - ☐ Two new cryoplants (assuming one installed for ultimate).
  - Essentially all protection devices to be replaced with more robust designs, possibly requiring also layout changes.
  - Upgrade of the beam dump system. Additional hardware.
  - ☐ Half of the phase 1 collimation system to be reviewed (replaced).
  - □ Remote handling mandatory in parts of the machine.
  - □ Additional service galleries?
  - □ Absolute filters and modifications of ventilation system.
  - □ Additional shielding in some areas.
  - Upgrade of permanent vacuum bake-out system.
- Nobody argued that an LHC intensity upgrade to beyond ultimate is impossible.
- "With enough money everything is possible...;-)"
- Yes, but effort and cost might be significant...

Thank you Ralf for bringing some sanity into this business. SM

#### **PS2** and **SPL**

#### M.Benedikt





Since PS2 would not come before 2020, the existing chain must be maintained anyhow

# May expect decision on PS2/SPL

Intensity Limitations (10 <sup>11</sup> protons per bunch)							
	Present	SPL-PS2	2GeV in PS				
LINAC4	4.0	4.0	4.0				
PSB or SPL	3.6	4.0	3.6				
PS or PS2	1.7	4.0	3.0				
SPS	1.2	1.2	1.2				
LHC	?	?	?				

S.Myers

Faster and Cheaper

# Then how come to highest Lumi?

Crab cavities (R.Calaga)

Luminosity levelling (JP. Koutchouk)

Stable operation (many)

Upgrade of PSB to 2 GeV (M.Giovannozi)

Inner triplet (phase 2 or/and 1)?

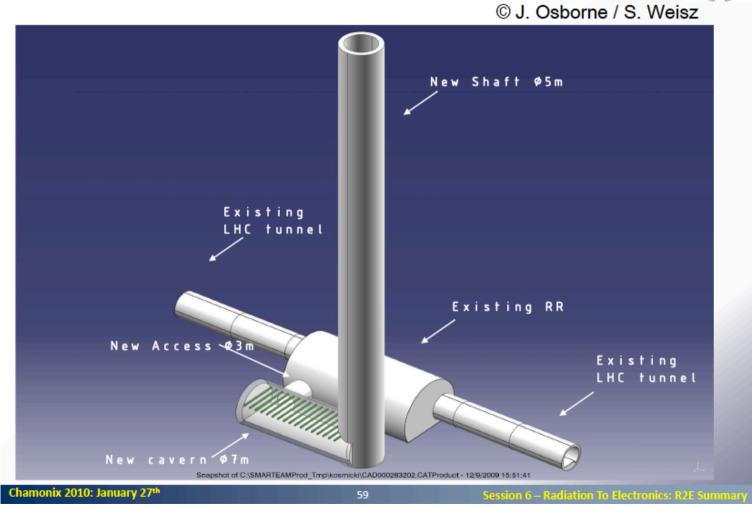
Radiation hardness (M.Bugger)

SPS coating of ~700 dipoles against e cloud (Shaposhnikova)

• • •

# **Making LHC Radiation Hard**





Relocation Recabling New electronics

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Time constrains

M. Bugger

# Then how come to highest Lumi?

$$L = \frac{1}{4\pi} f_{rev} n_b \gamma \frac{1}{\beta^* (\gamma \varepsilon)} N_b^2 \frac{1}{\sqrt{1 + \phi_{piw}^2}}$$
**L** [10<sup>34</sup> cm<sup>-2</sup>s<sup>-1</sup>]
$$0.5 N_b = 2.3 \times 10^{11}$$

$$0.5 N_b = 1.7 \times 10^{11}$$

$$0.5 N_b = 1.15 \times 10^{11}$$

### **Personal Summary**

- 1. The LHC is a complex machine which in early design phases was pushed hard.
- 2. If the splices stand it and all runs well, we get max 50pb<sup>-1</sup> for ICHEP and 1fb<sup>-1</sup> in 11.
- 3. A long shutdown then follows to repair what was overlooked and to improve.
- 4. By 2014 we may then have 10-30 fb<sup>-1</sup>, less than projected before, but at 6.5-7 TeV.
- 5. The investments in stabilizing, maintaining and upgrading the machine will continue to be highly demanding for CERN.
- 6. Given the huge cost and manpower consumption and its merits I expect the PS2/SPL to not go ahead.
- 7. The LHC will be with us until 2030 and is run by impressive people.

We did not see the Mont Blanc but we trusted it exists..